



Some Thoughts on the Luminosity Spectrum

ALCPG Linear Collider Workshop

January 9th, 2004

SLAC

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University of Oregon

- What do we need
- Simple fits
- Goals for Paris



Who Cares?



What motivates the need to know dL/dE to high precision?

- Direct Reconstruction - m_H
- Threshold Scans - m_t

Bias to $\langle s' \rangle$ which counts
(same motivation for beam energy)

Others less important (i.e.: SUSY endpoints)
due to lower event rates

Beware of thinking you know the answer
before you begin! Nature may have surprises...

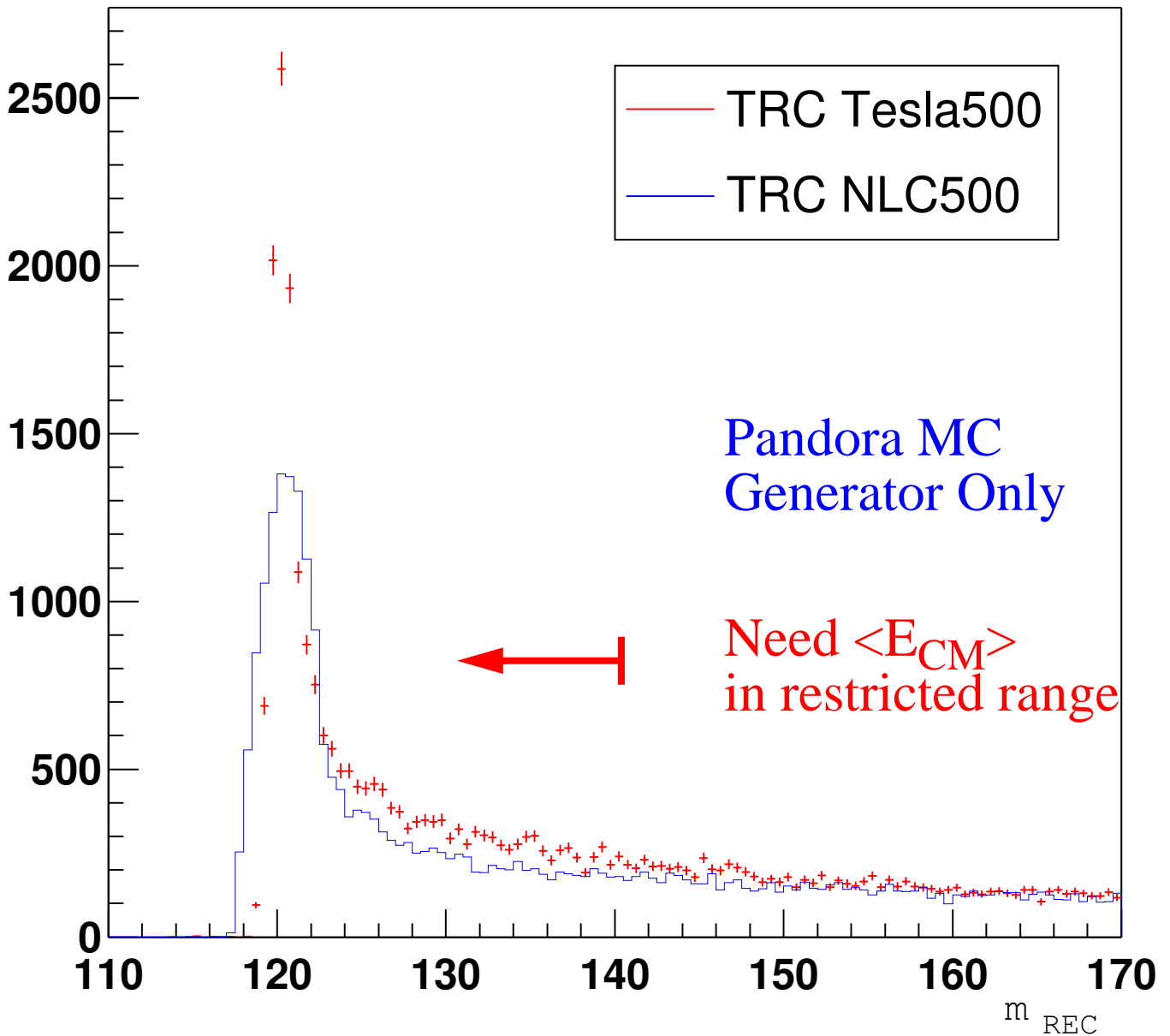
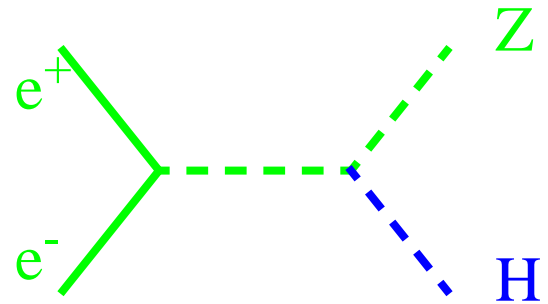


Higgs Example



Reconstruct recoil mass
from Z decay products

$$m_R^2 \approx (E_{CM} - E_Z)^2 - |\vec{p}_Z|^2$$
$$= E_{CM}^2 - 2E_{CM}E_Z + M_Z^2$$

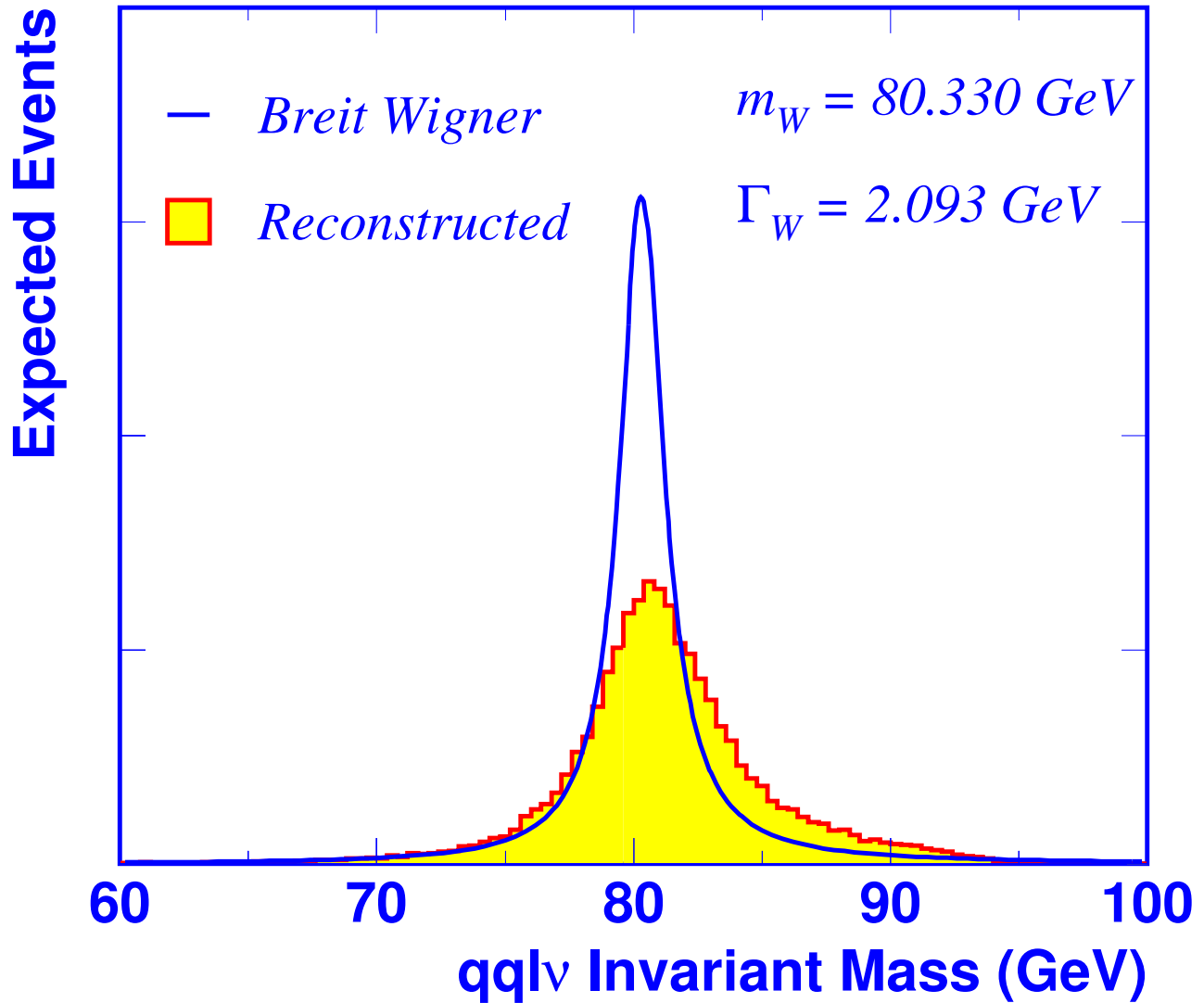




LEP II Example



W Boson Mass



Substantial bias due to ISR

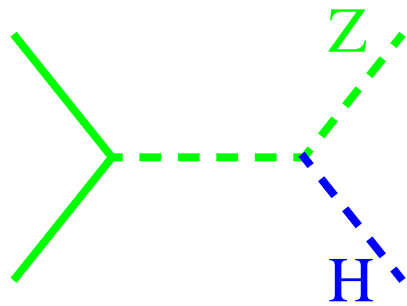
Can correct to high accuracy
using precision MC generator



Precision Analyses



Physics
Process

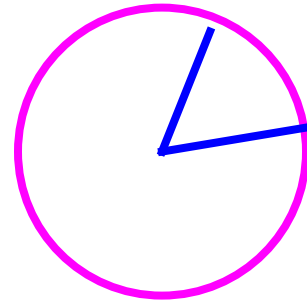


m_H

Simulation

ISR
BSL
Detector
...

Observable



$Z \rightarrow \mu^+\mu^-$

↑
must go here!

Must plug measured beam-beam effects
directly into precision MC description

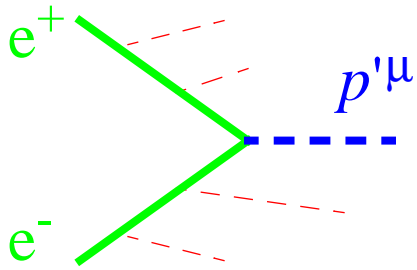
Bootstrap using some other physics process
(akin to Parton Density Functions...)

MC tools for extraction of luminosity spectra at LC,
S. Jadach, SLAC, 24th Oct. 2002

<http://home.cern.ch/jadach/public/LumLCslac.ps.gz>



What do we want?



4-vector producing physics

$$d^4L/dp'$$

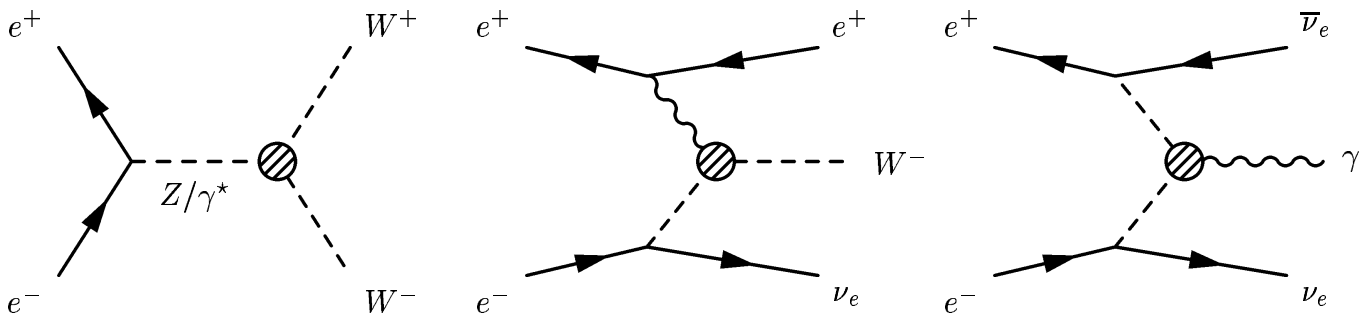
Note: s-channel bias...

Assume transverse components small...

Two parameters: E', p_z' or s', β or E^+, E^-

We talk about dL/dE ,

but at minimum we need $\frac{d^2L}{ds'd\beta}$

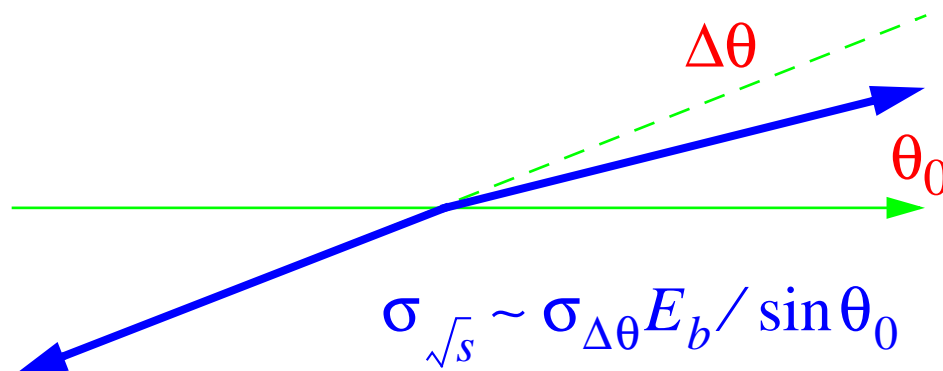




Acolinearity



What's the problem with acolinearity?



$$\frac{s'}{s} = \frac{\sin \theta_1 + \sin \theta_2 - |\sin(\theta_1 + \theta_2)|}{\sin \theta_1 + \sin \theta_2 + |\sin(\theta_1 + \theta_2)|}$$

First-order approximation!

$$\Delta\theta = f(\beta) \text{ where } \beta \approx \frac{E^+ - E^-}{E^+ + E^-}$$

Physics actually depends more

$$\text{on } \frac{dL}{ds'} \text{ than } \frac{dL}{d\beta}$$



What to do?



Strategy #1

'Measure' $\frac{d^2L}{ds'd\beta}$ (or at least $\frac{dL}{ds'}$)

from observable $\frac{dL}{d\beta}$

Implicitly assumes $\frac{dL}{dE^+} = \frac{dL}{dE^-}$

(or at least that the relation is known)

Further assumes that the correlations are understood:

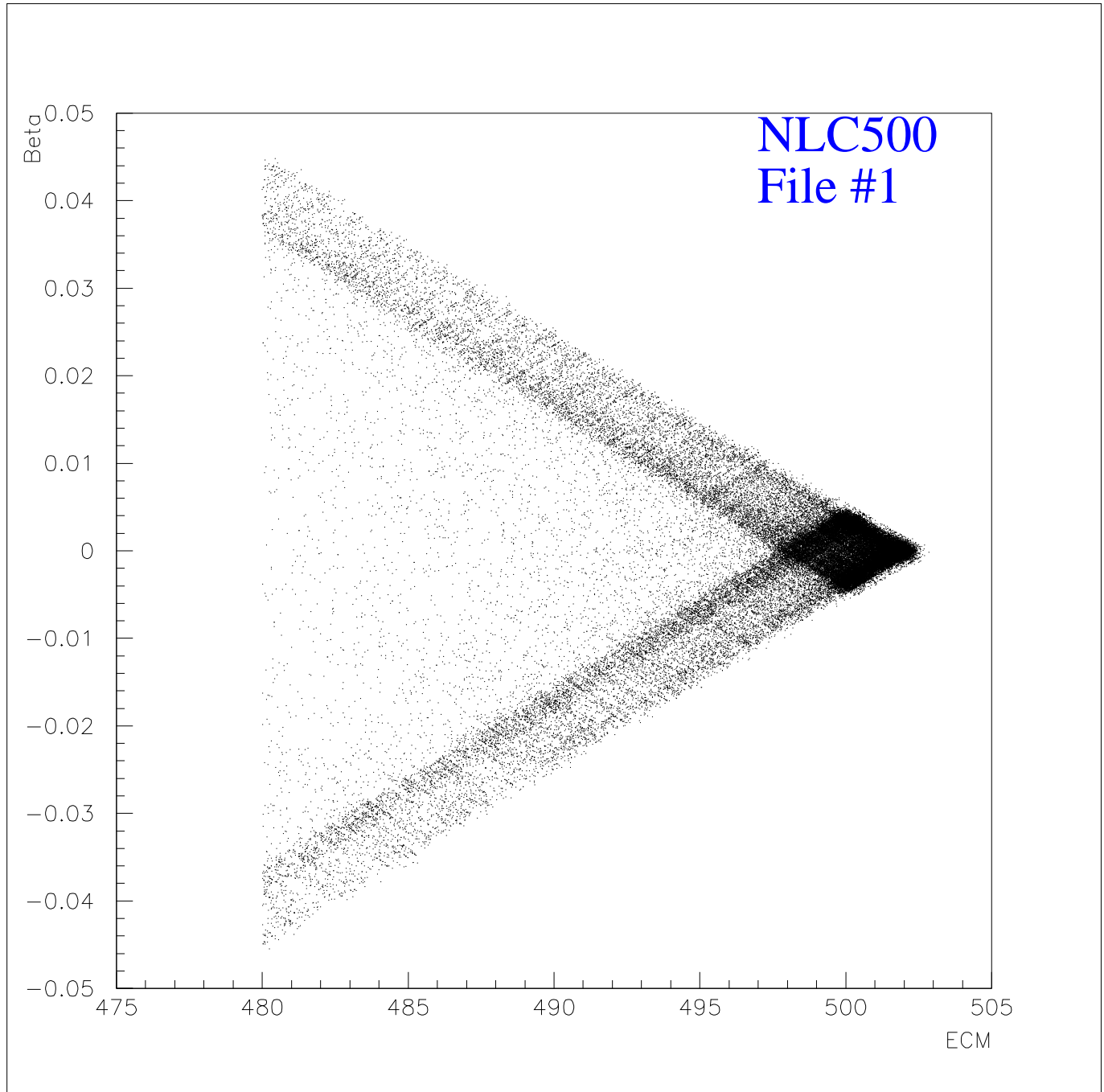
$$\frac{d^2L}{dE^+dE^-} = f\left(\frac{dL}{dE^+}, \frac{dL}{dE^-}, \rho(E^+, E^-)\right)$$

Mike has shown that these are potentially large and variable, must be measured

Don't trust simulations too much...



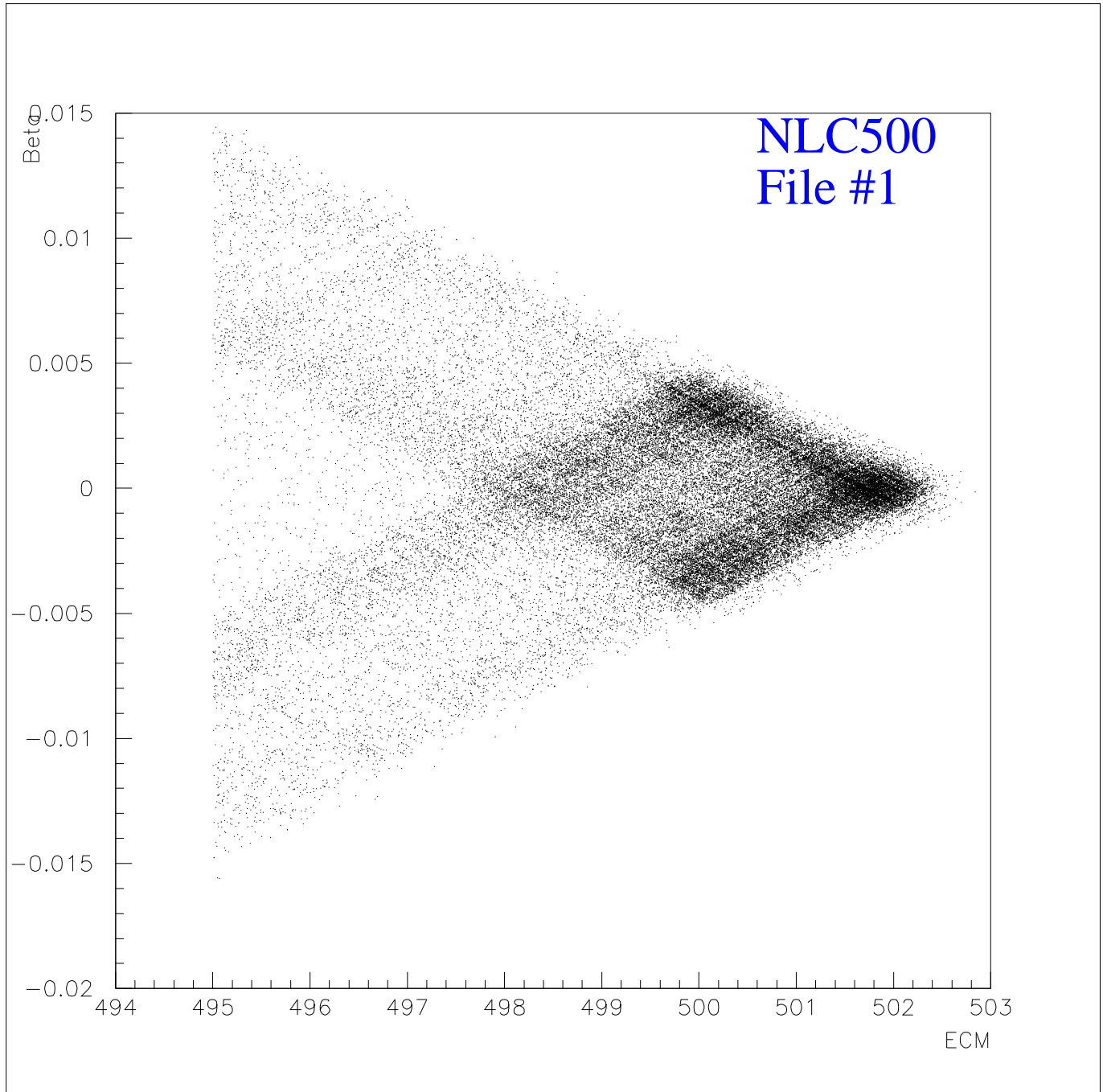
Beta vs. ECM I



Guinea Pig using A. Seyri 'realistic' inputs



Beta vs. ECM II



Guinea Pig using A. Seyri 'realistic' inputs



What to do?



Strategy #2

Use all information available in $e^+e^- \rightarrow e^+e^-$

$E^+, E^-, \theta^+, \theta^-, \phi^+, \phi^-$ or $E_0, \Delta E, \theta_0, \Delta\theta, \phi_0, \Delta\phi$

(ignore transverse info in ϕ for now)

Using $\frac{d\sigma}{d\theta_0}$ probably hopeless...

➤ 3 observables

$\Delta E, \Delta\theta$ redundant: cross-check!

Comparing $\frac{d^2L}{dE^+dE^-}$ vs. $\frac{dL}{dE^+}$ or $\frac{dL}{dE^-}$

is a direct measure of correlations!

(comment by Klaus Monig)

Only way to get what we truly want!



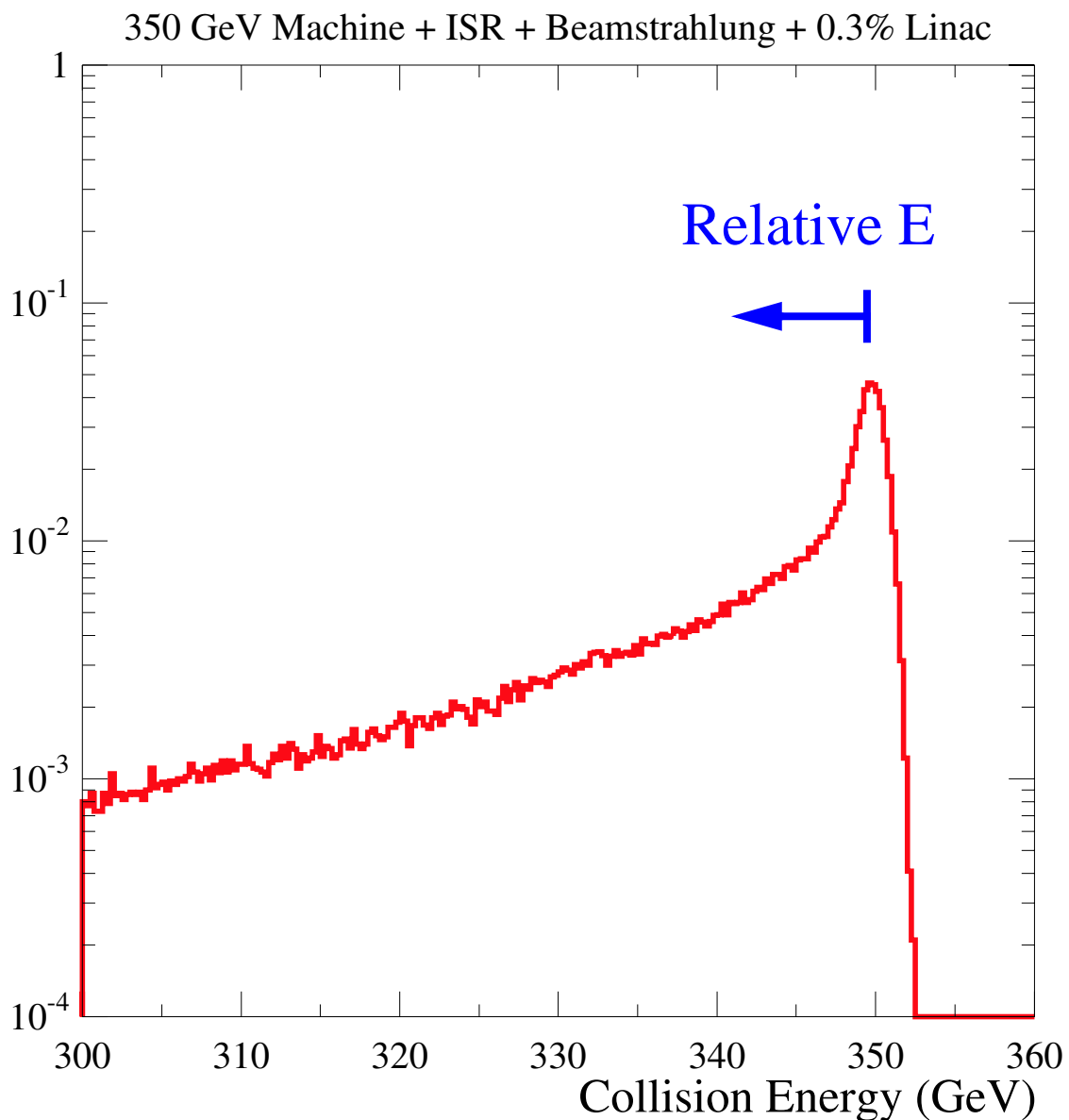
Calorimeters?



Conventional wisdom is that calorimeters
can't do better than $\sim 1\%$ on E

For the spectrum we want relative energy
from the peak, not absolute energy from zero.

Easier?





Montpellier

Talks by S. Bogart and E. Poirer about
'unfolding' dL/dE spectrum

Really want a 'simple' function with
few (<100) parameters to plug into MC

Take one step backwards

How bad do you do by fitting
Gaussian convoluted with Circe
if you had perfect data?

Tools

- A. Seyri 'realistic beam' files (NLC 500)
- Guinea Pig
- RooFit

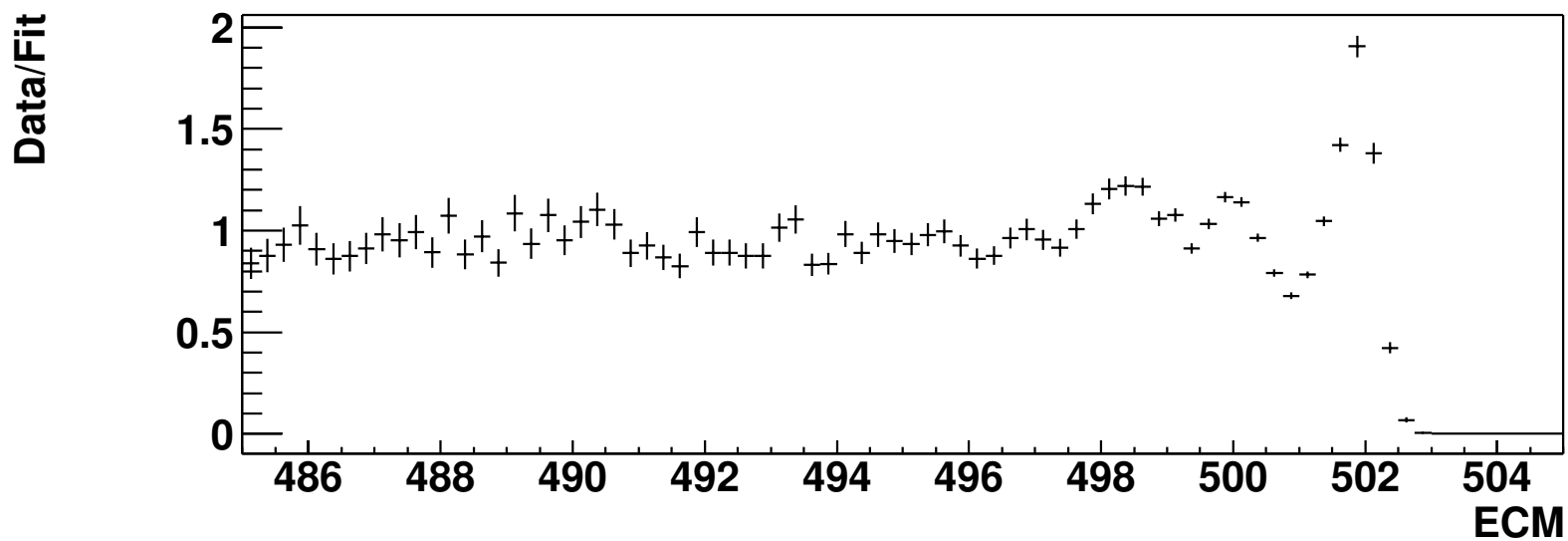
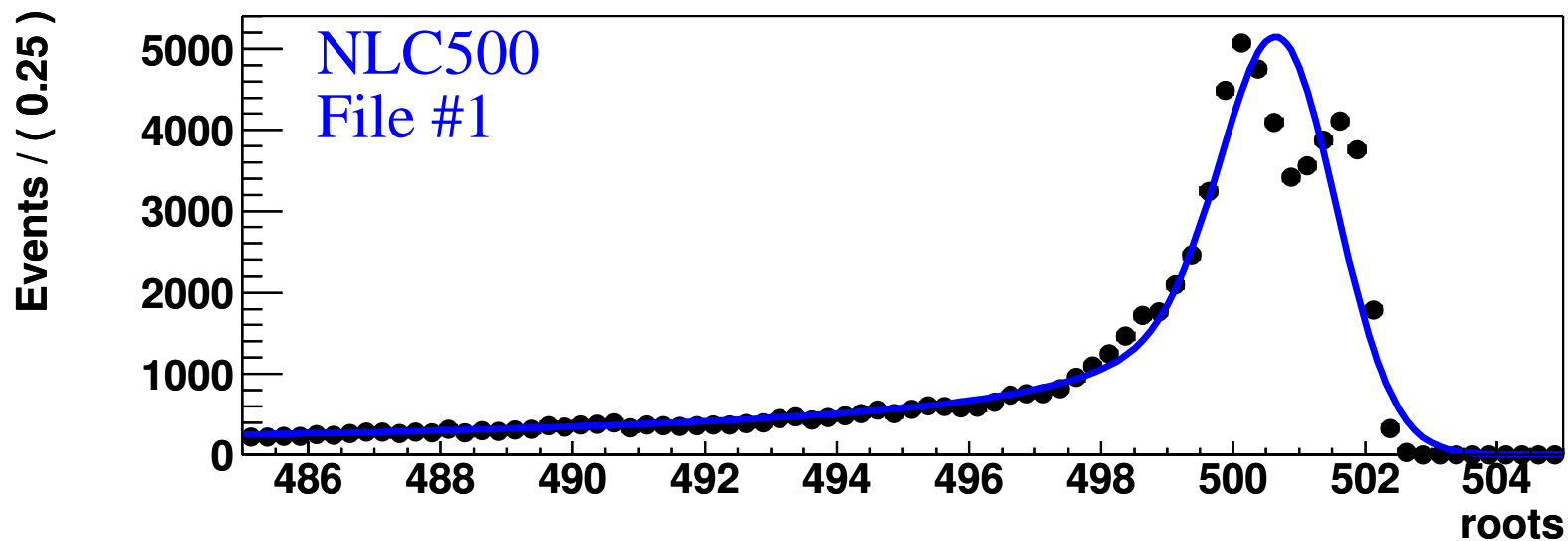
$$P(E) = g(E; E_0, \sigma) + a_0 \int_E x^{a_1} (1-x)^{a_2} g(E') dE'$$

where $x = E/E'$

5 parameters!

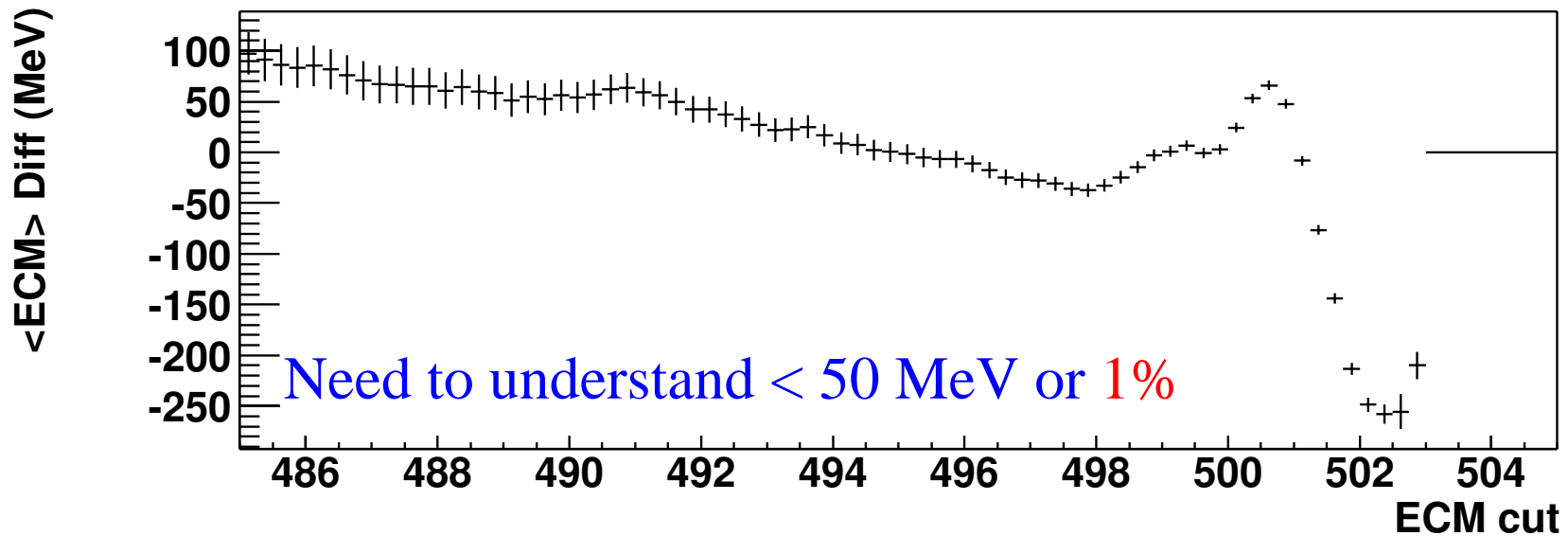
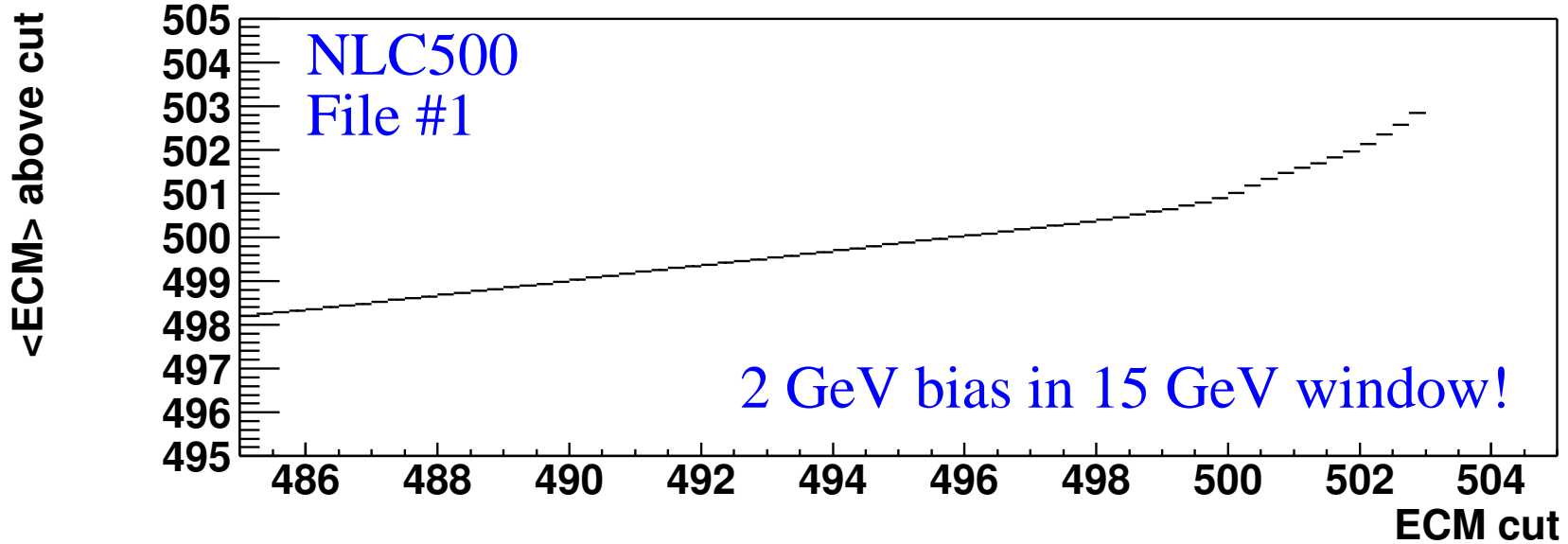


Good Fit?



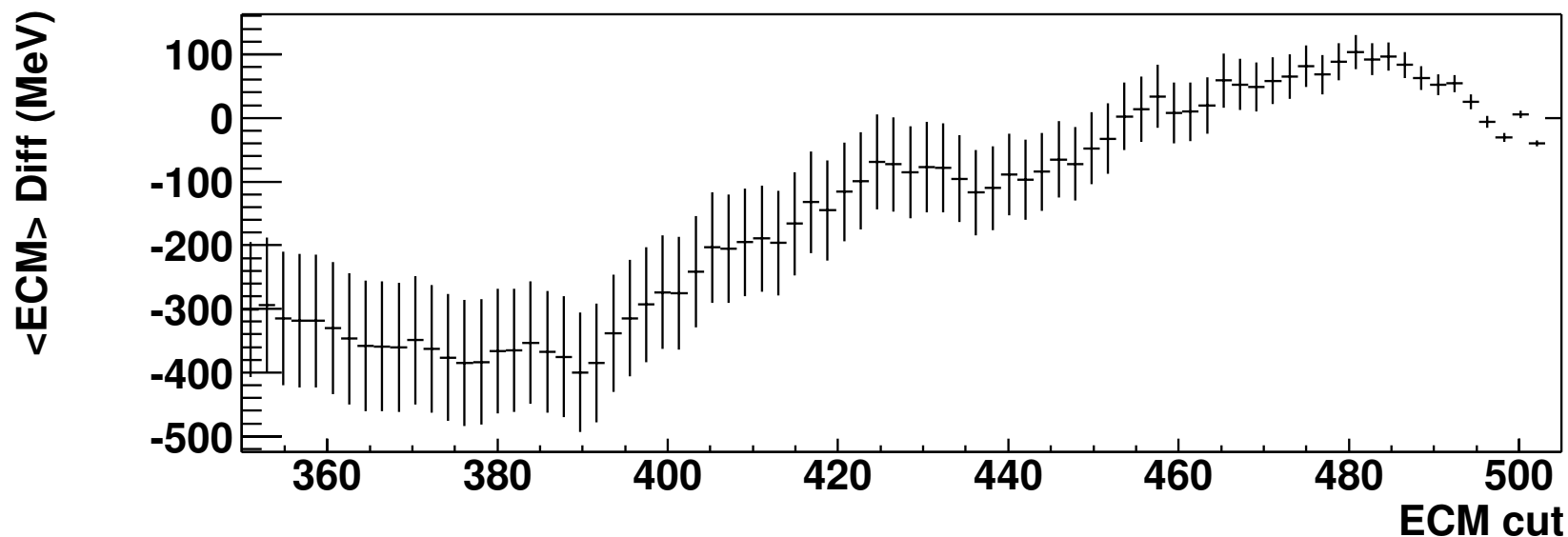
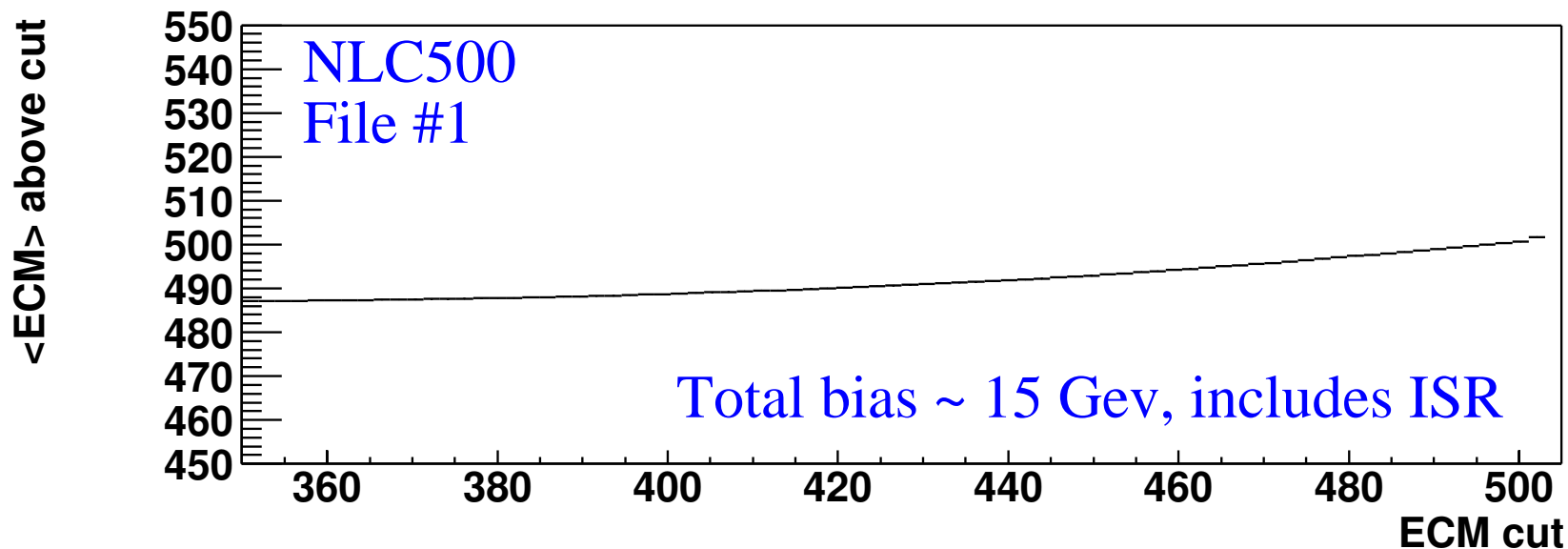


Does it matter?





Wider Range





Results



One more parameter might do it.

What about variation?

$\sqrt{s'}$ above	$\Delta\langle\sqrt{s'}\rangle$ (in MeV) for file				
	1	2	3	5	6
0 GeV	+11	+4	-128	-93	-8
250 GeV	-41	-4	-143	-130	-25
350 GeV	-300	-140	-305	-320	-198
450 GeV	-35	+90	+15	-61	-154
475 GeV	+77	+60	+97	+63	124

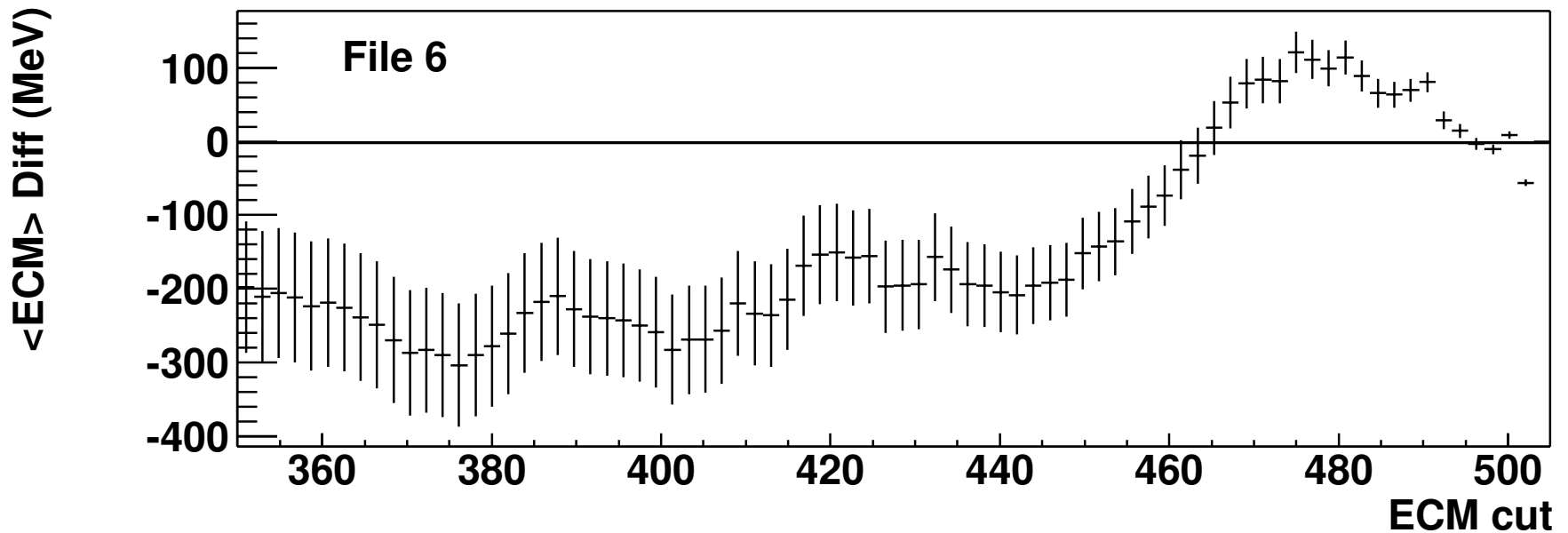
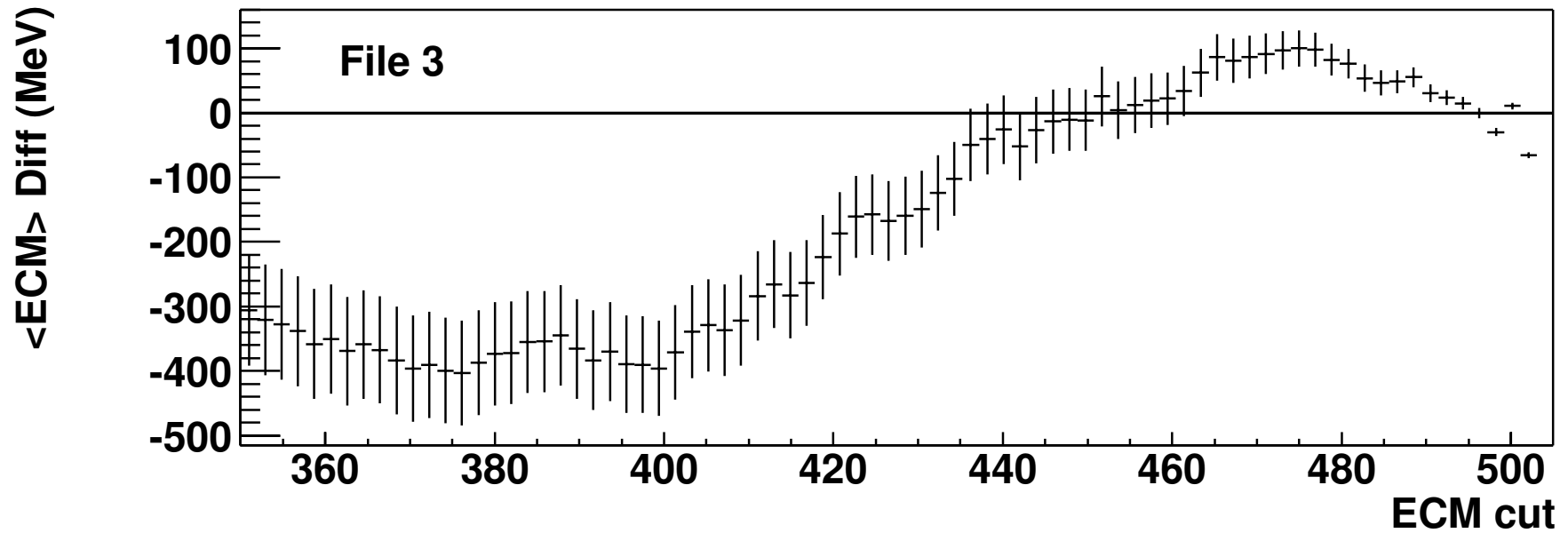
Errors range from 30 to 120 MeV

Bias in $\Delta\langle\sqrt{s'}\rangle$ from direct fit to data
varies significantly (100s MeV)

Simple function not adequate to track
'realistic' variations!
(and not just core, but tails also)



Another Look





Final Thoughts



Peak-region looks problematic for warm machine

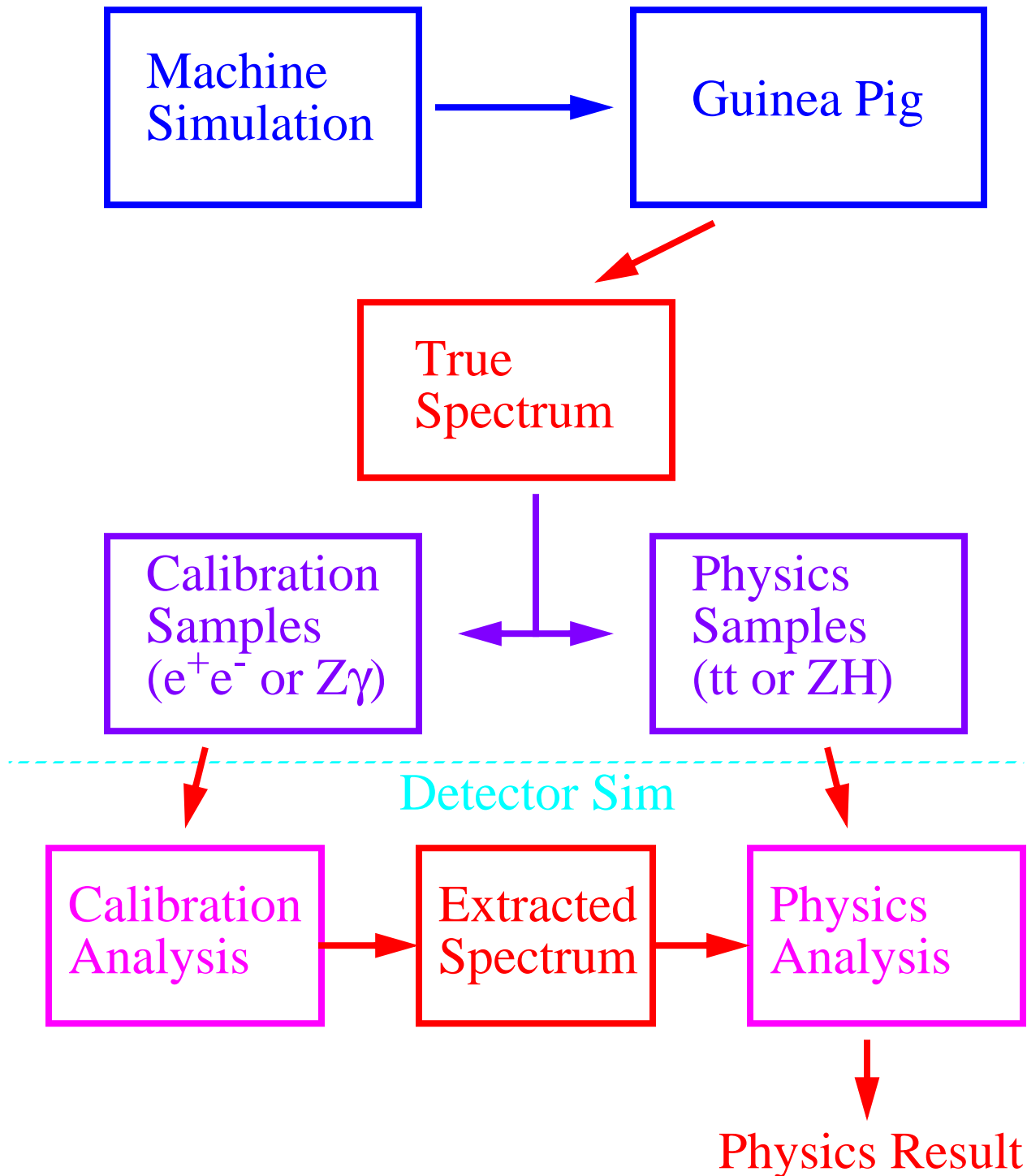
Detector resolution will tend to wash this out

Need proper study with Guinea Pig,
detector simulation, and realistic analysis method

Cold machine isn't off the hook, either!
Vertical offsets lead to uncomfortable numbers

Only way to show there isn't a problem
is to measure it to sufficient precision...

Quantify warm/cold difference for Paris?



Repeat under different machine assumptions
(failure modes?) to get variation of extracted result



The way ahead?



Luminosity spectrum and beam energy
are two sides of the same problem

Need full “cradle-to-grave” analysis
pulling all the pieces together!

