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# Computational Adaptive Mesh Refinement

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# Overview

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- \* Astrophysics is a highly multi-scale problem!
- \* Sophisticated physics needed on all scales
- \* Eulerian methods with large dynamical range in space and time are possible with embedded meshes
- \* With appropriate on-the-fly criteria, meshes can be embedded adaptively
- \* Extreme dynamic range now possible

# Multi-Scale Physical Problems

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- \* Primordial stars have formation efficiency of 0.03%!
- \* Galaxies shaped by stars, star clusters, merger history
- \* Must be able to adequately resolve large scale structure and small scale in order to develop next-generation subgrid models!

# Dynamic Range

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	Size	Relative Size
Visible Universe	$1.3 \times 10^{23}$ km	1.00
Galaxy Cluster	$3 \times 10^{19}$ km	$2 \times 10^{-4}$
Galaxy	$6 \times 10^{17}$ km	$5 \times 10^{-6}$
Star Cluster	$2 \times 10^{15}$ km	$2 \times 10^{-8}$
Star	700,000 km	$5 \times 10^{-18}$
Earth	6,000 km	$5 \times 10^{-20}$
Us	0.002 km	$1.5 \times 10^{-26}$

# Physics Necessary

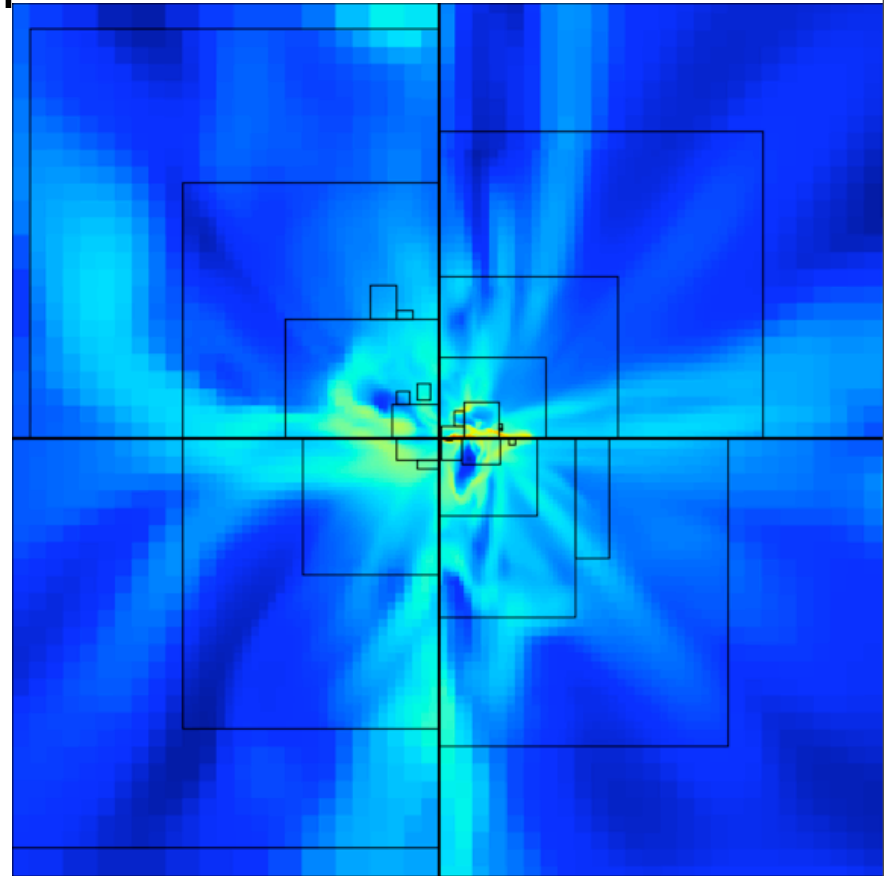
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- \* Radiative cooling
- \* Ionization physics
- \* Background radiation
- \* Multi-species fluids
- \* Extensible to new subgrid models

# Adaptive Mesh Refinement

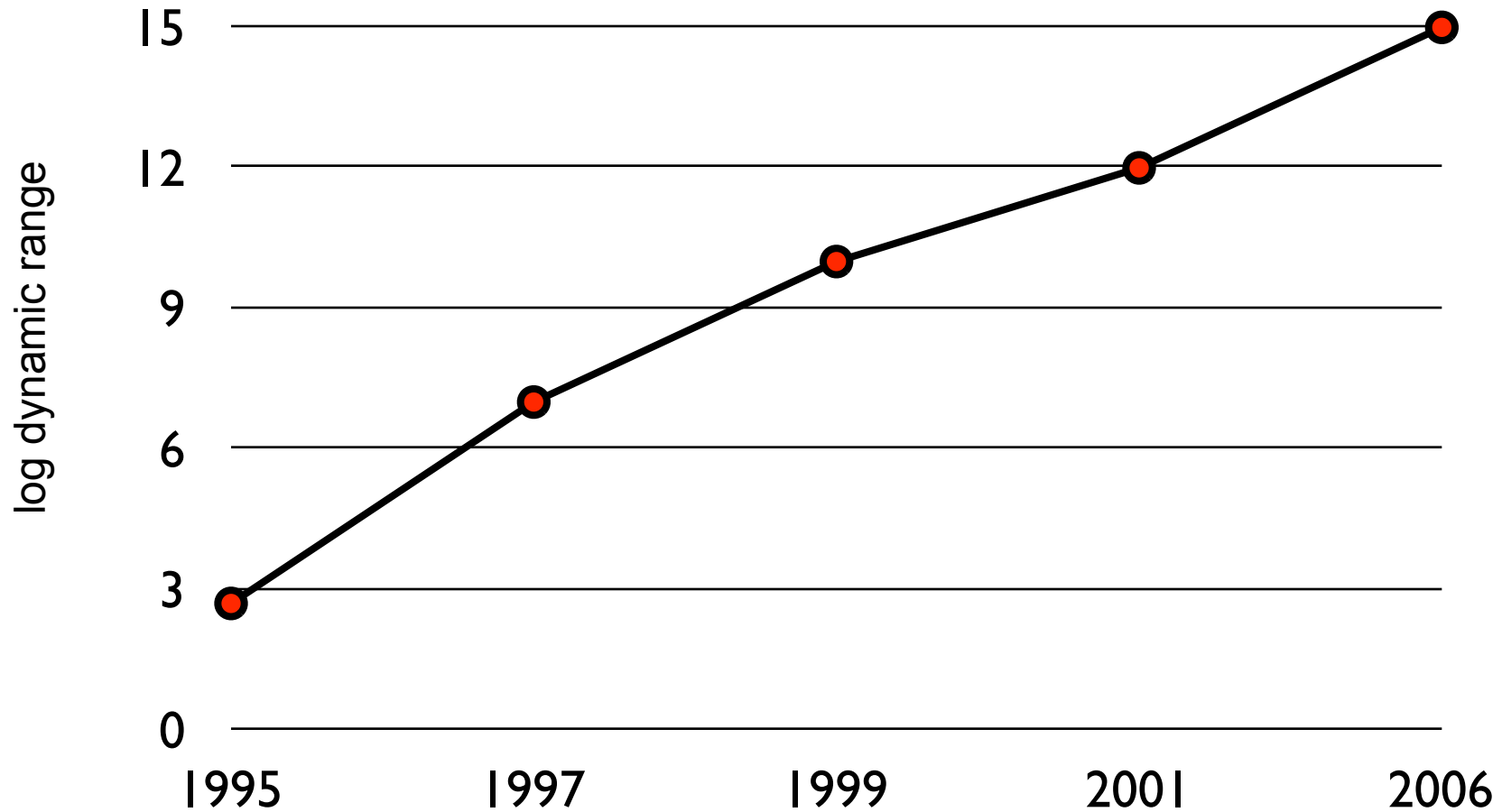
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- \* Higher-order hydrodynamic schemes
- \* Interpolate to higher resolution
- \* Correct fluxes of conserved quantities across boundaries



# Adaptive Mesh Refinement

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# Codes Available

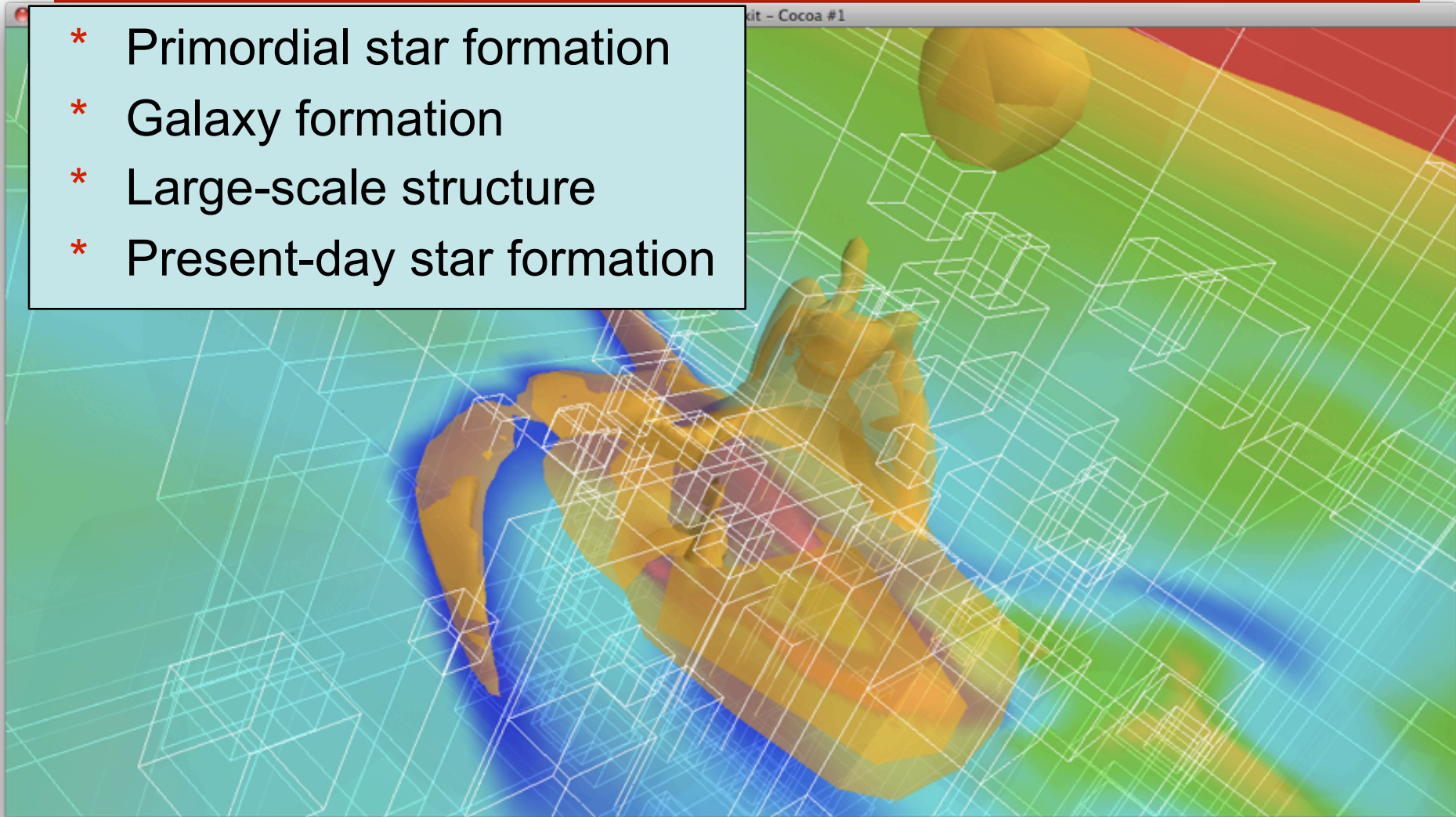
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- \* Diversity of codes:
  - Enzo (SLAC, UCSD, Colorado, Columbia)
  - Orion (LBL, LLNL, Princeton, UCSC)
  - ART (UChicago, Fermilab)
  - Chombo (LBL)
- \* Enzo code:
  - Wide use in cosmology
  - Freely distributed, community developed
  - Patch-based AMR
  - Piecewise Parabolic Mesh hydro reconstruction
  - 12-species chemistry network



# Simulational Domain

- \* Primordial star formation
- \* Galaxy formation
- \* Large-scale structure
- \* Present-day star formation



# Computational Domain Expanding

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- \* Modern simulations run on hundreds if not thousands of processors
- \* Computational infrastructure at SLAC supports large-scale simulations
- \* Studying formation of large scale structure with all attendant physics (galaxy feedback, cosmic rays, chemistry, star formation) is nearer than ever
- \* Galaxy catalogs, simulated observations, lensing studies