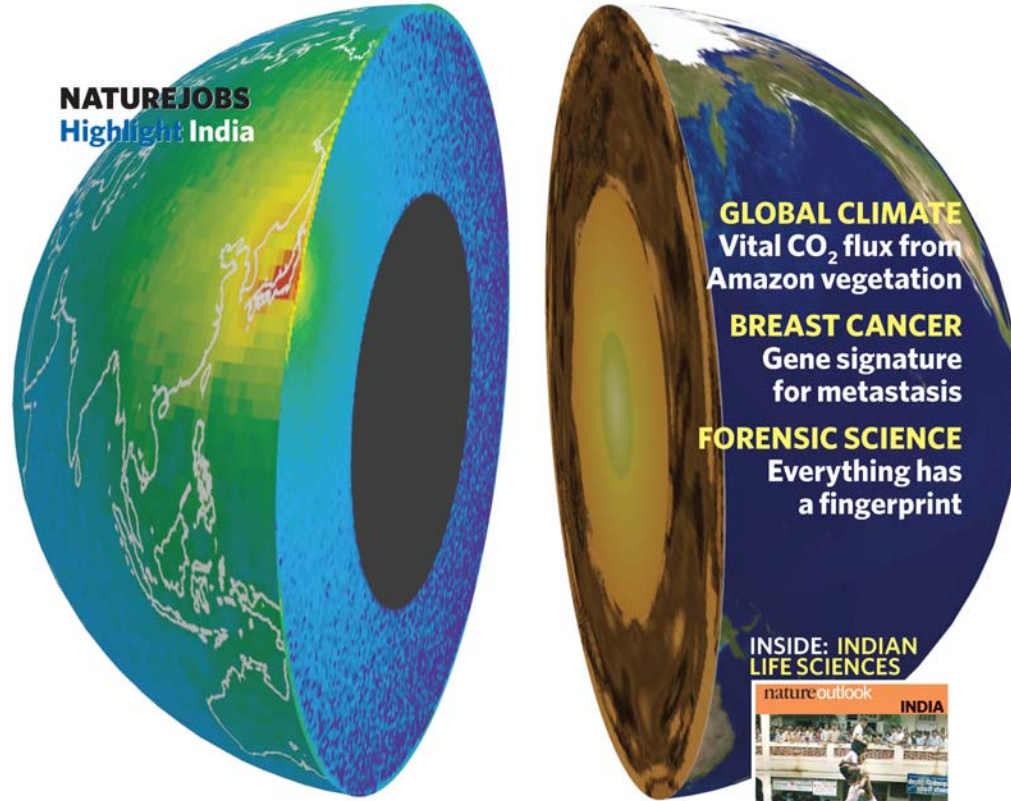


nature

NATUREJOBS
Highlight India



GLOBAL CLIMATE
Vital CO₂ flux from
Amazon vegetation

BREAST CANCER
Gene signature
for metastasis

FORENSIC SCIENCE
Everything has
a fingerprint

INSIDE: INDIAN
LIFE SCIENCES



EARTHLY POWERS

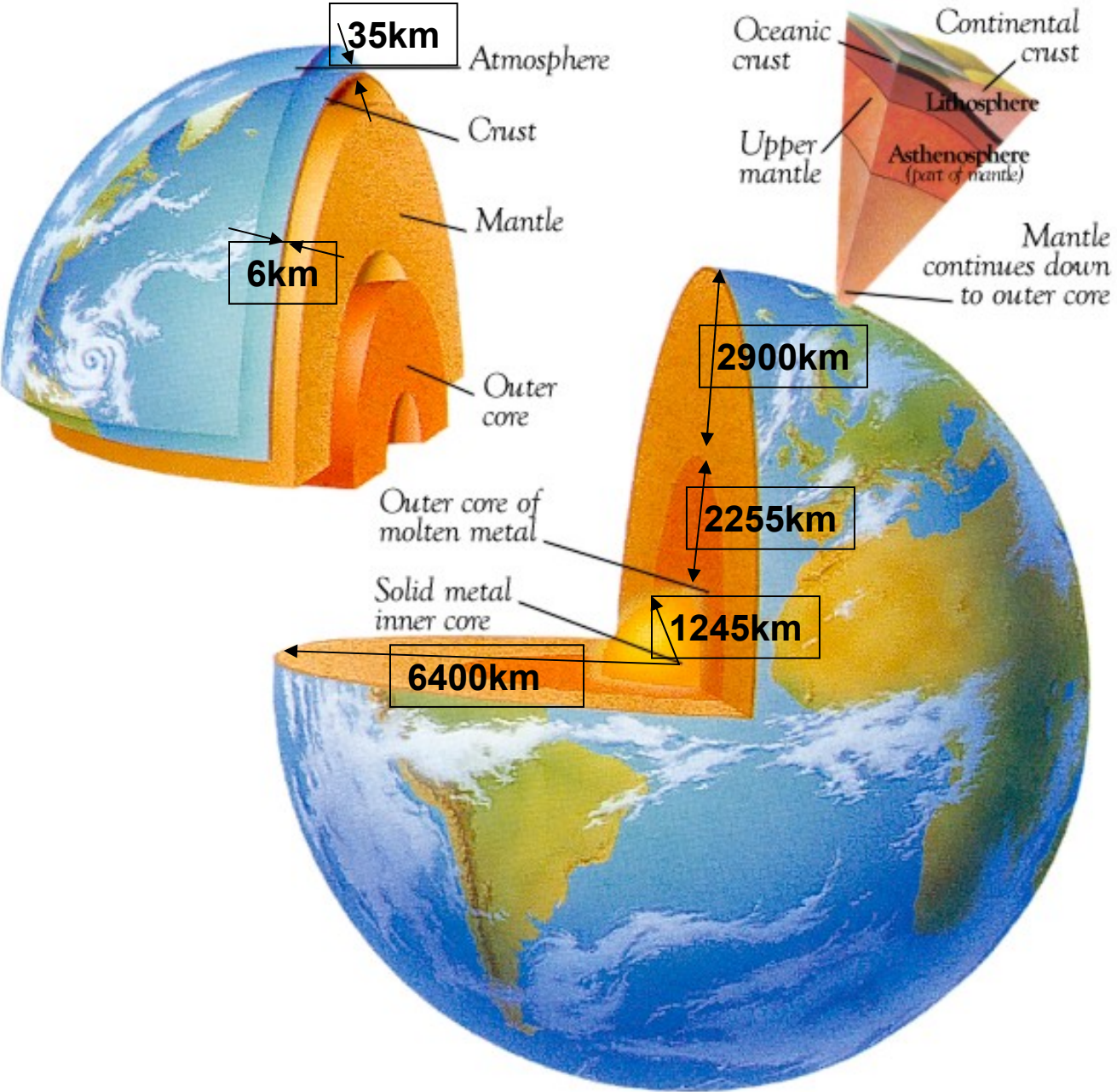
Geoneutrinos reveal Earth's inner secrets



The dawn of neutrino geophysics

G. Gratta
Physics Department
Stanford University

Structure of the Earth



- From seismic data 5 basic regions:
 - inner core,
 - outer core,
 - mantle,
 - oceanic crust,
 - continental crust and sediments
- All these regions behave like solids except the outer core.

Convection in the Earth

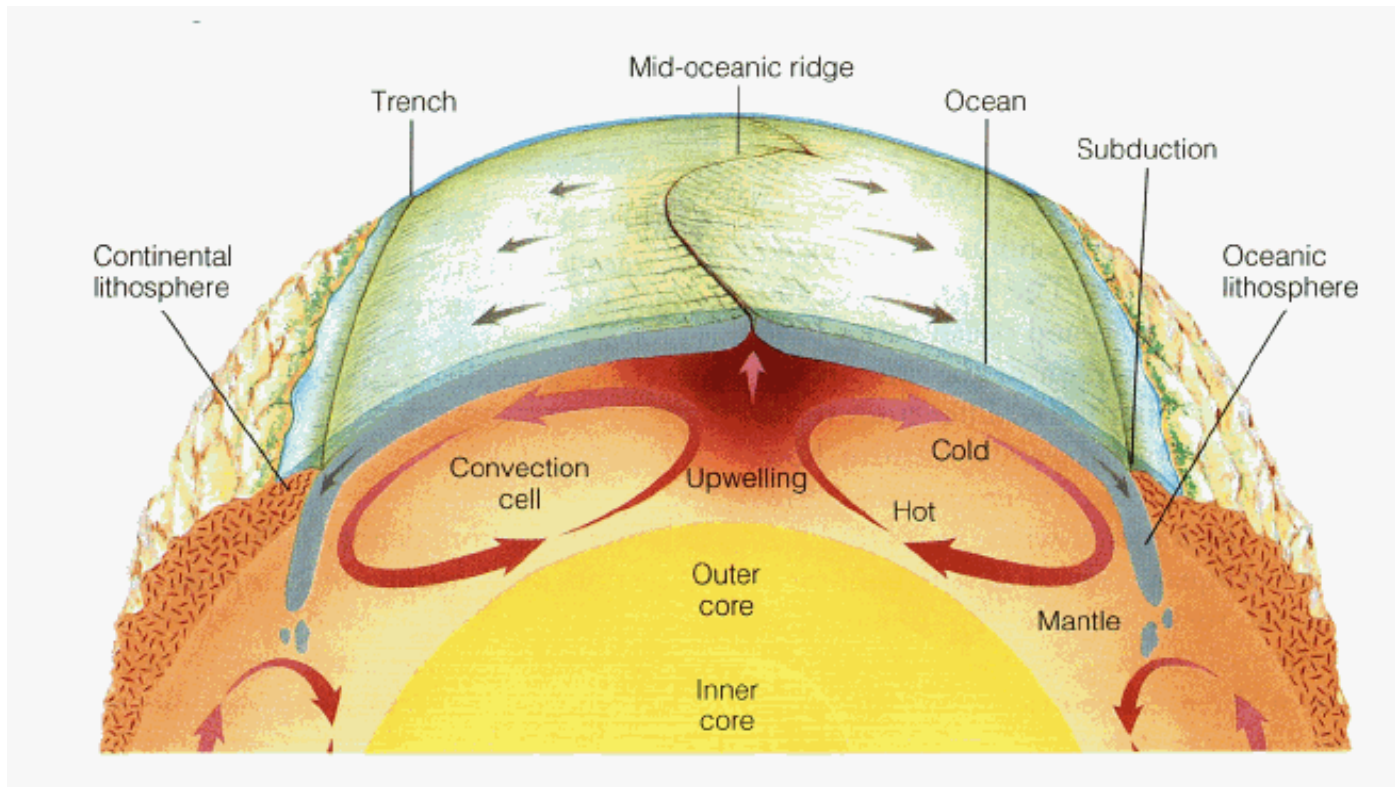


Image: <http://www.dstu.univ-montp2.fr/PERSO/bokelmann/convection.gif>

- The mantle appears to convect even though it is solid.
- This is responsible for plate tectonics and earthquakes.
- Oceanic crust is being renewed at mid-ocean ridges and recycled at trenches.

Remember, always,
the words of
Francis Birch (1952)



“Unwary readers should take warning that ordinary language undergoes modification to a high-pressure form when applied to the interior of the Earth. A few examples of equivalents follow:

<i>High-pressure form</i>	<i>Ordinary meaning</i>
certain	dubious
undoubtedly	perhaps
positive proof	vague suggestion
unanswerable argument	trivial objection
pure iron	uncertain mixture of all the elements

Very specific data on the Earth's interior is hard to collect

The only universal probe for the interior of the Earth is seismology. But this is only sensitive to the elastic properties of the rocks.

Nomenclature derives from the seismic boundaries. Composition is then guessed for the different regions that are assumed homogeneous in composition.

Seismically motivated nomenclature is then used at times to signify a region of a certain composition. This is not necessarily kosher.

OK, let's say that the Earth probably has the same composition as the Solar System

Meteorites

Achondrite, Ca-poor, Diogenite



Mantle-crust pieces (?)

Johnstown

Allende



chondrites

undifferentiated planets (?)

Carbonaceous chondrite (CV3)



Imilac

Pallasite: olivine and iron mixtures (CMB?)

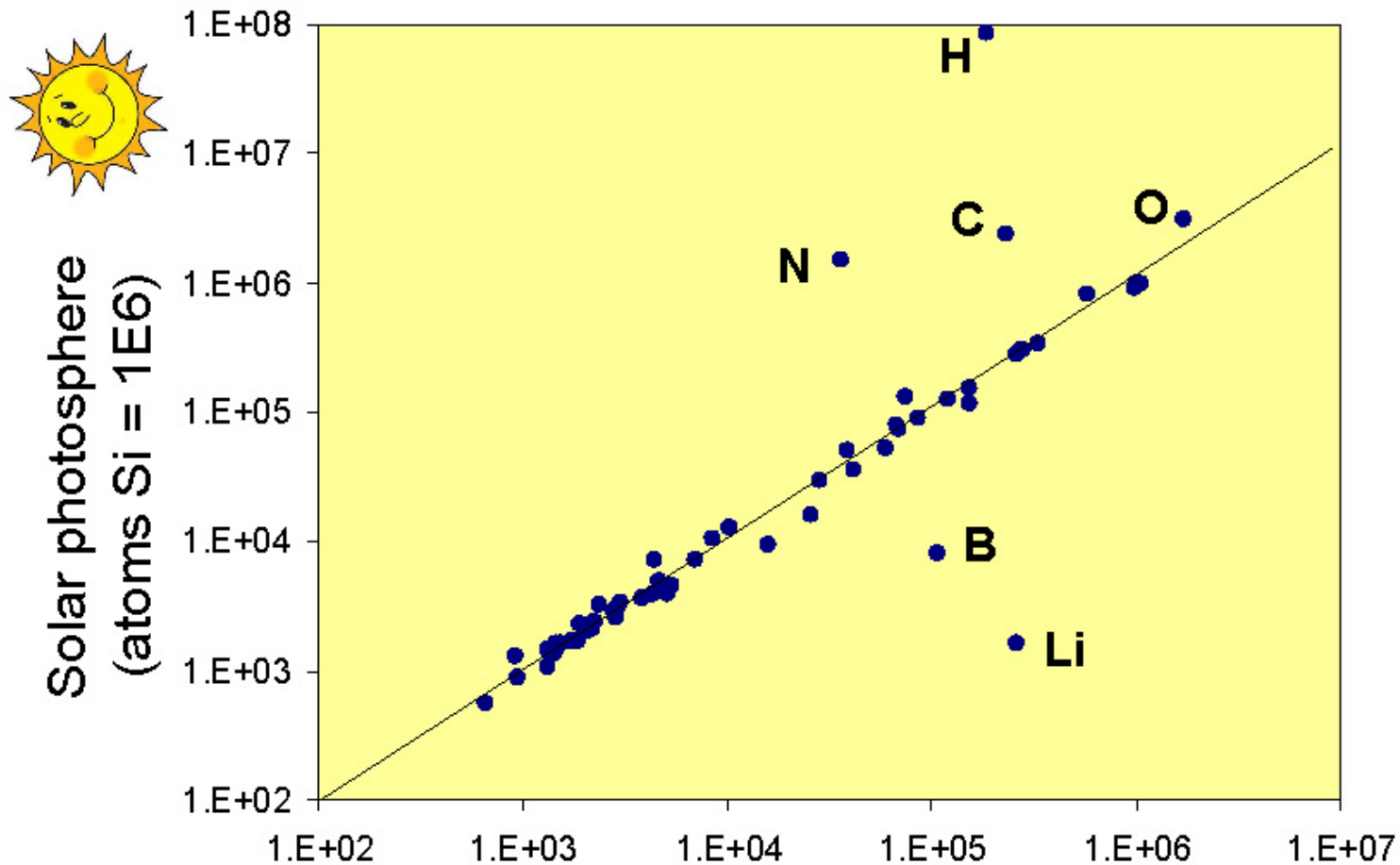


Henbury

IIIAB

Irons: pieces of core

C1 chondrites have very similar composition to the solar photosphere (except for peculiar light elements that are expected to be anomalous around the Sun)



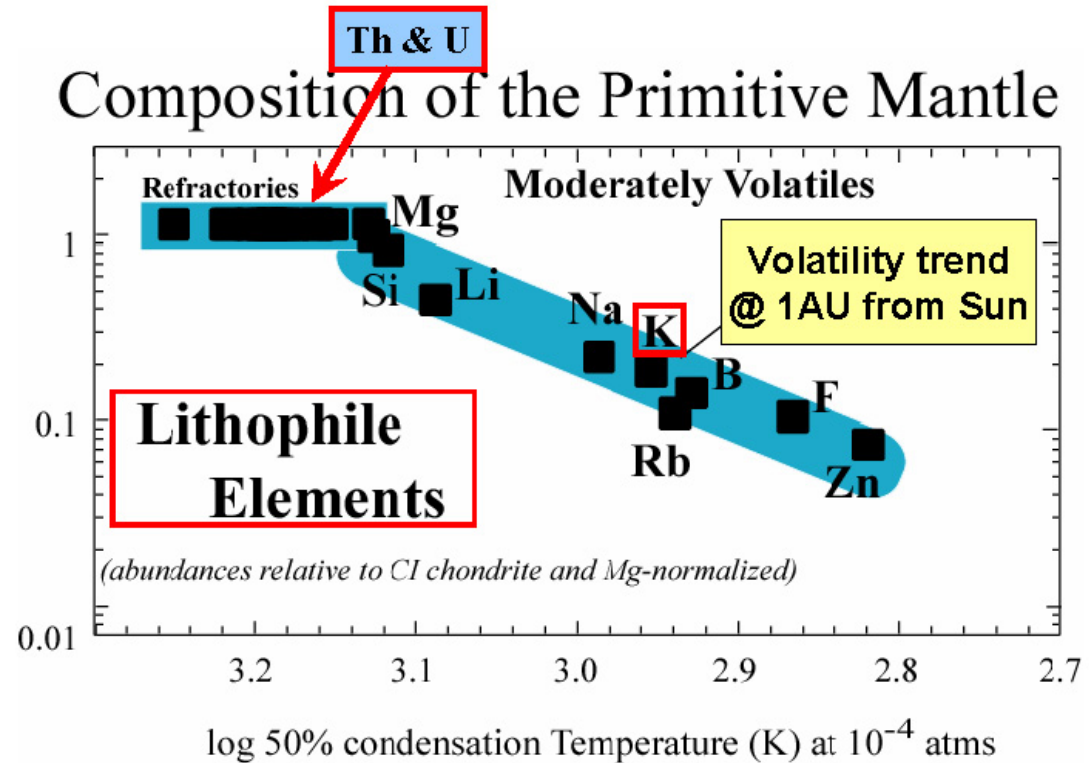
McDonough, Neutrino 2008



C1 carbonaceous chondrite
(atoms Si = 1E6)

Next:

1) Correct for the loss of volatile elements during the Earth's formation



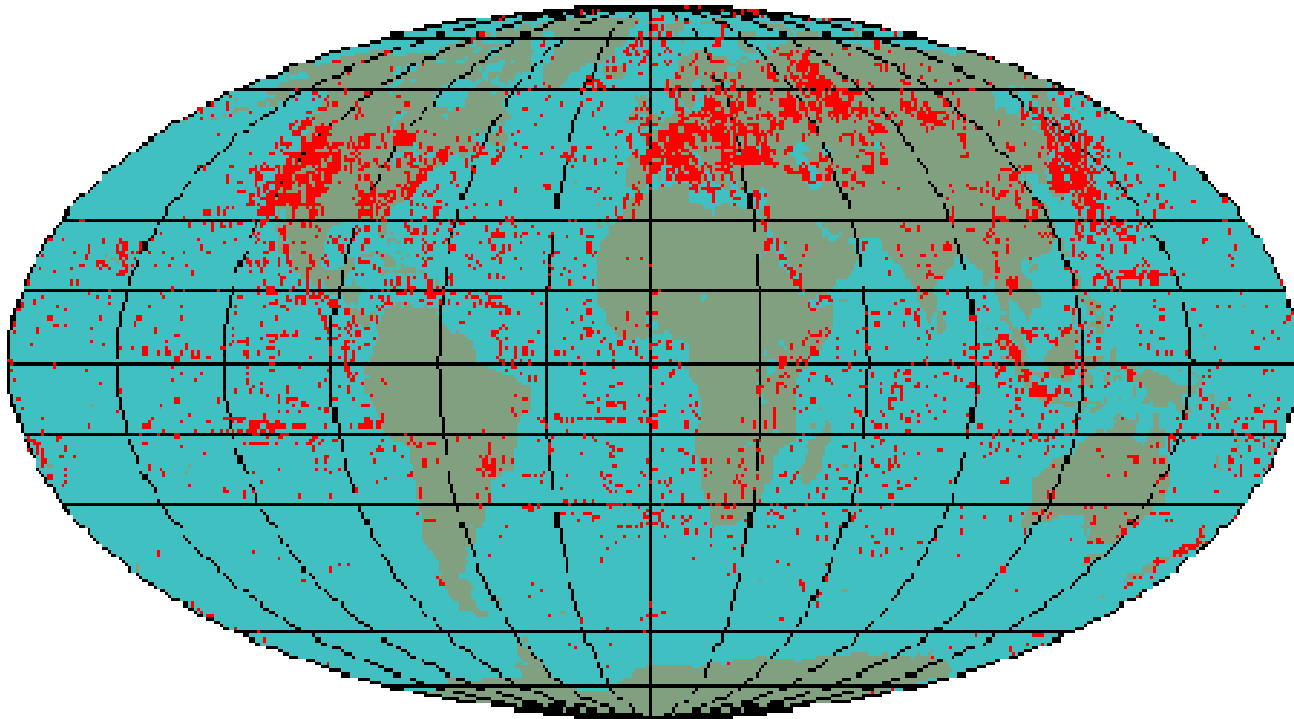
McDonough, Neutrino 2008

2) Based on chemical affinity estimate the composition of different regions

→ Core expected to have insignificant U, Th

→ Independently know that U, Th are ~1000x more common in the crust (ppm) than in the mantle (ppb)

Only a shallow layer has been sampled for chemical composition by drilling/sampling



Deepest bore-hole (12km)
is only $\sim 1/500$ of the Earth's radius.

Q: What powers the Eyjafjallajökull?



Q: What powers the Eyjafjallajökull?



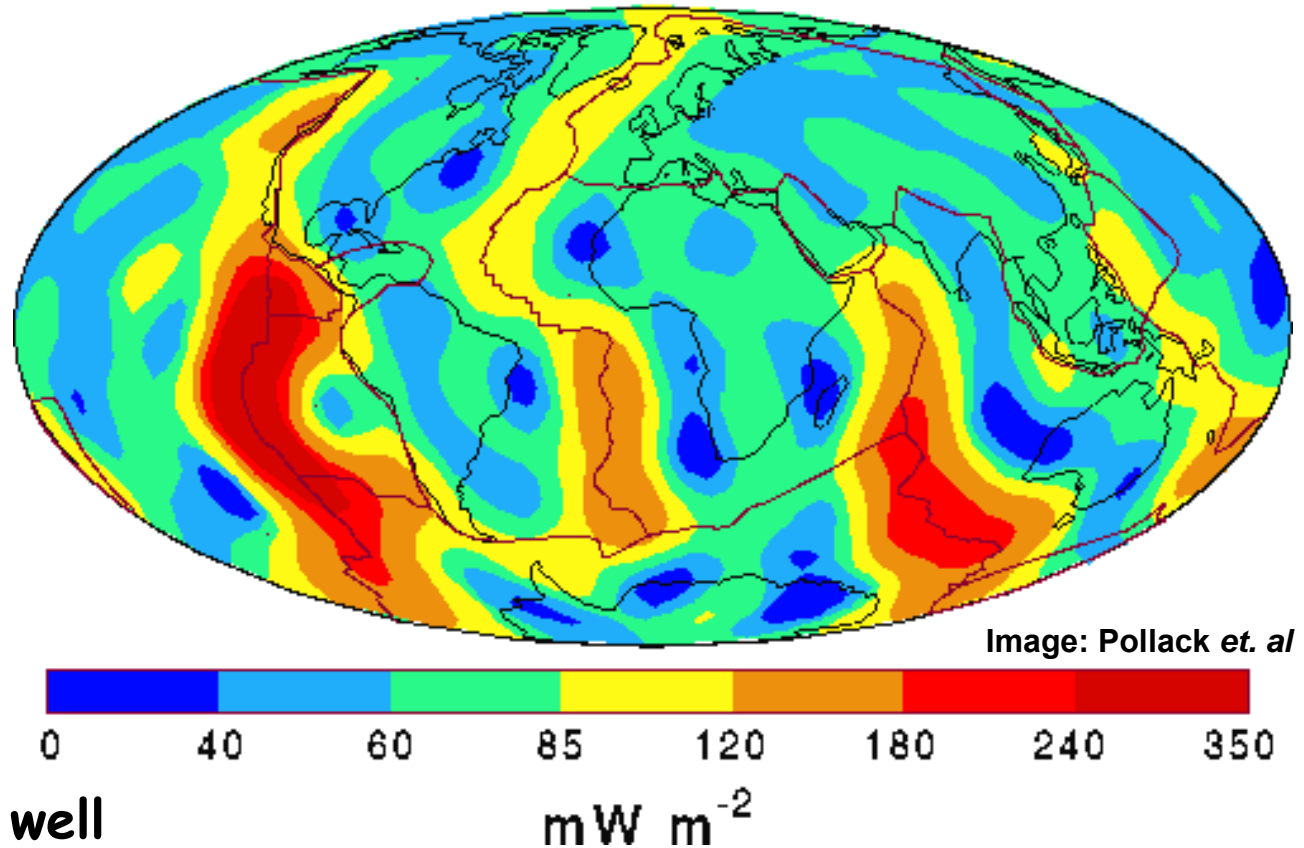
...more generally what powers plate tectonics
that powers volcanoes?

Same boreholes used to measure the heat flow from the Earth's interior

- ΔT_{hole} is measured between 2 points far away along the borehole
- Thermal conductivity C_{rock} of the rock is measured in the lab
- $Q = \Delta T_{\text{hole}} C_{\text{rock}}$ (assuming pure conduction)
- But in addition have to account for mantle convection
- *Get a total 46 ± 3 TW (100 mW/m²)*
- Error is small BUT recent analysis with different convection model gives 31 ± 1 TW (61mW/m²)...

Heat flow from the Earth

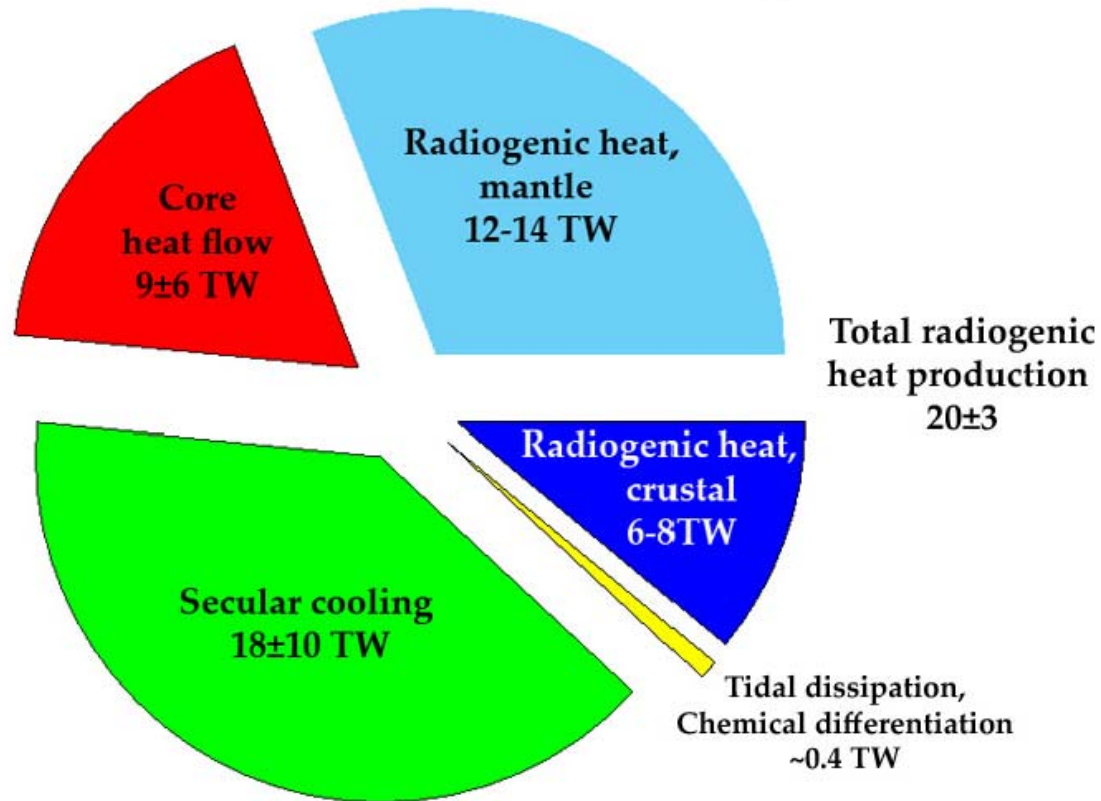
Note the large emission under the mid ocean ridges (~83% of the total heat!): this is where mantle convects this is also where the pure conduction model really does not work well



Note the units!
The Earth receives $\sim 1 \text{ kW m}^2$ from the Sun, so the surface temperature has nothing to do with the heat produced inside.

Note that we need 15 TW to run society. This is sizeable compared to the total output from the planet!

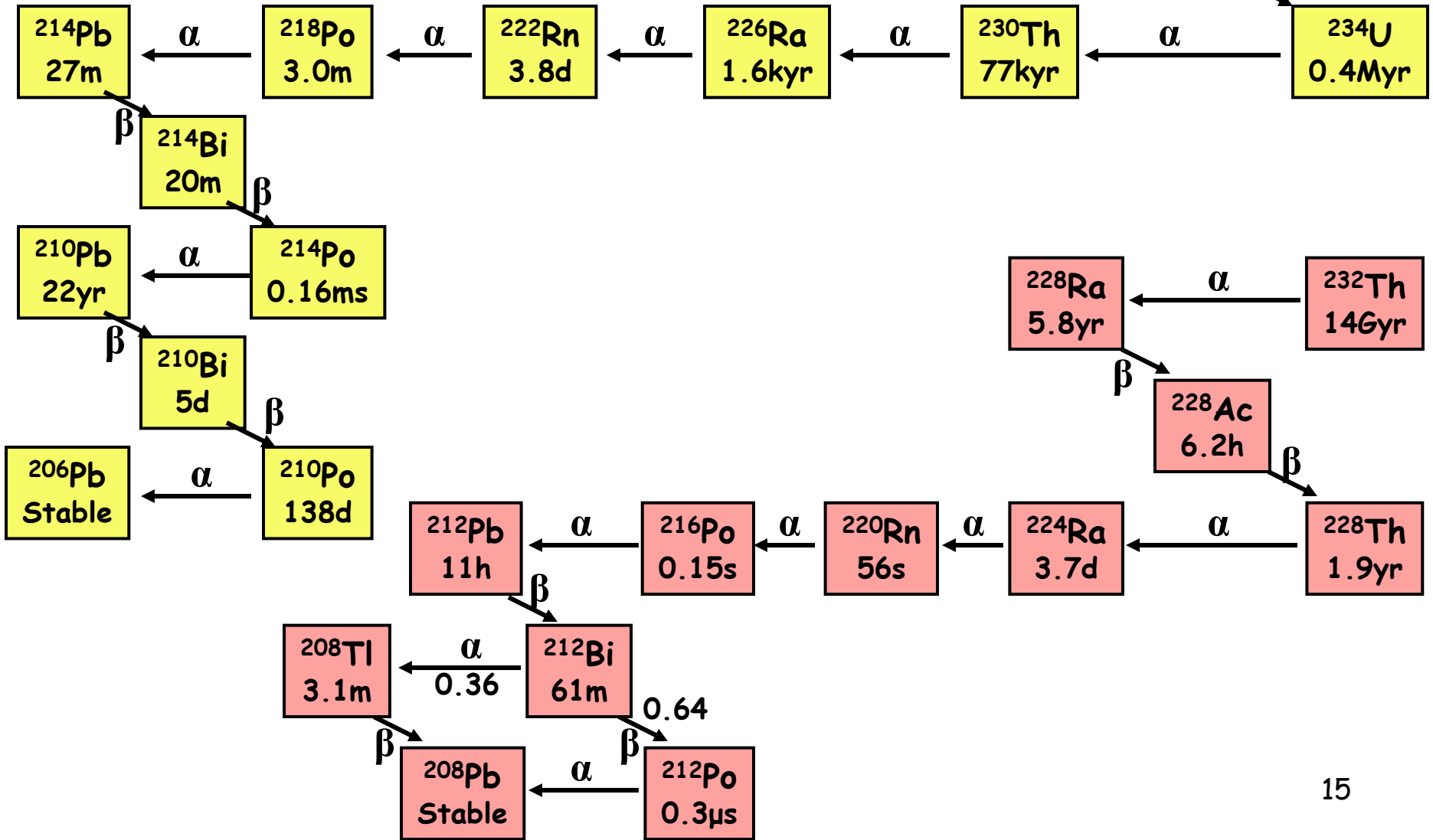
What produces the heat ?



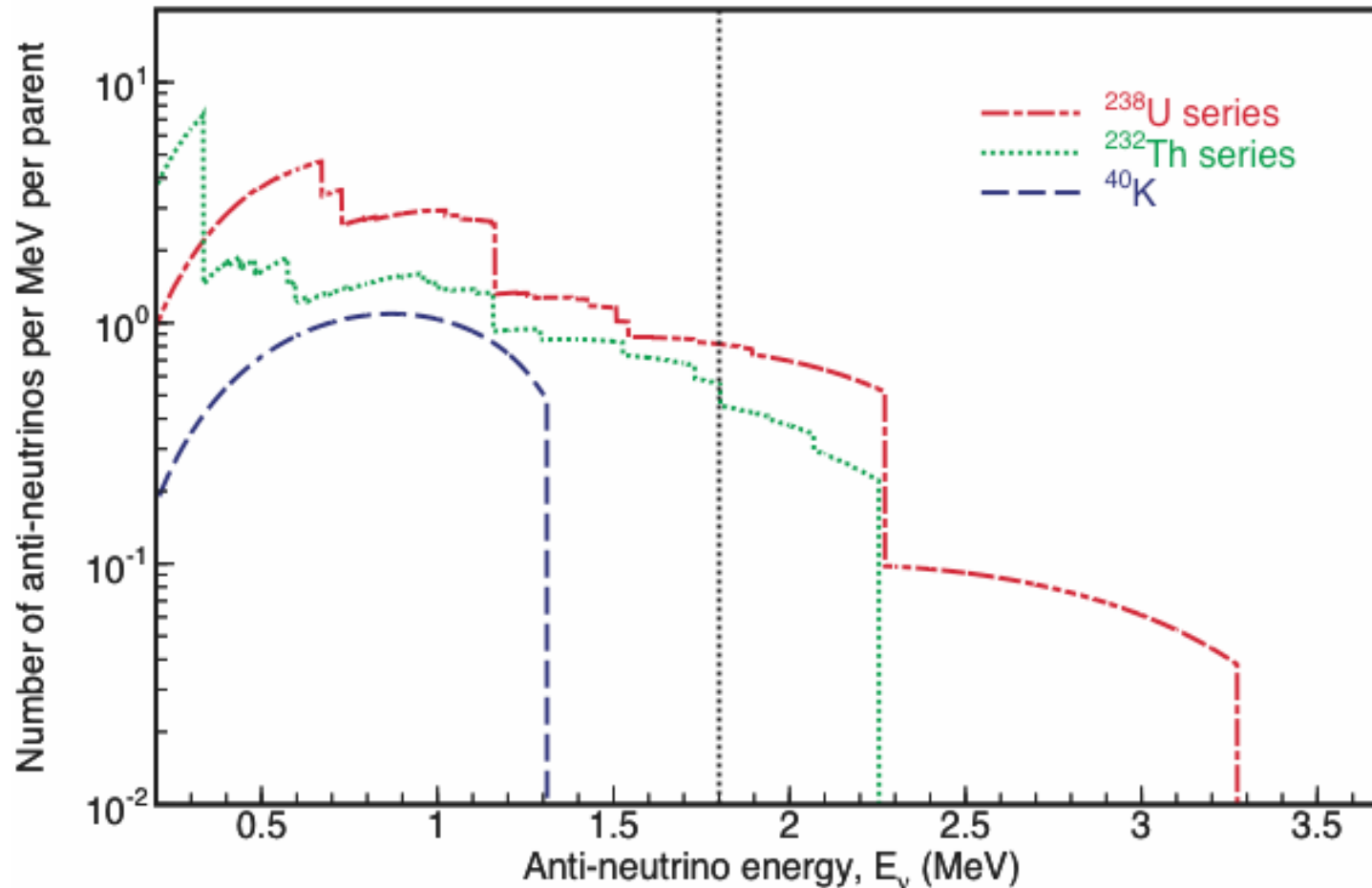
after Jaupart et al 2008 Treatise of Geophysics

**Radiogenic heat / Total heat
called "Urey ratio" → believed to be 0.3 to 0.7**

Both ^{238}U and ^{232}Th are primordial radioactive isotopes with long decay chains including β^- decays



$\bar{\nu}_e$ of different endpoint energy are emitted at each β^- decay step producing characteristic spectra for ^{238}U , ^{232}Th (and ^{40}K)



Geo-neutrinos are a rare witness of the chemical composition (at least for U, Th and K) of the Earth

Can they be detected?

Yesterday's background is today's signal...



Fred Reines (?) working at a neutrino detector (circa 1953)

G.Gratta

SSI 2010, 6

Dear Fred,
Just accved to me
that your background
neutrinos my just be comming
from high energy β -decaying
members of U and Th families
in the crust of the Earth. Do
not have on the train any
inform. to check it up, but it
seems the order of magn. is
reasonable. In fact the total energy
radioactive energy production
under one square foot of surface
may well be equal to the
energy of solar radiation falling
on ~~area~~ that surface ...
what do you think?

write to me at: The Union
Univ. of Mich. Ann Arbor. Mich

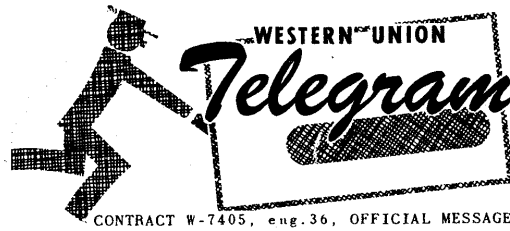
Yours GCO.



The Chief


...Well... not quite !

That detector was
some 5 orders of
magnitude too small



THIS MESSAGE IS TO BE SENT

- night letter ..
- day letter ..
- straight wire ..

TO: DR. GEORGE GAMOW
THE UNION
UNIVERSITY OF MICHIGAN
ANN ARBOR, MICHIGAN

~30 TW

MESSAGE:

FROM NUMBERS IN VREY BOOK ON THE PLANETS, EQUILIBRIUM HEAT LOSS
FROM EARTH'S SURFACE IS 50 ERGS/CM²SEC. IF ASSUME ALL DUE TO
BETA DECAY THEN HAVE ONLY ENOUGH ENERGY FOR ABOUT 10⁸, 1 Mev
NEUTRONS PER CM² AND SEC. THIS IS LOW BY 10⁵ OR SO. SHORT
HALF LIVES WOULD BE MADE BY COSMIC RAYS OR NEUTRONS IN EARTH.
IN VIEW OF RARITY OF COSMIC RAYS: I.E. ABOUT EQUAL TO ENERGY
OF STARLIGHT AND OF NEUTRONS IN EARTH THIS SOURCE OF NEUTRONS
SEEMS EVEN LESS LIKELY AS A SOURCE OF OUR SIGNAL.

RETURN ADDRESS OF SENDER:

Frederick Reines and Clyde L. Cowan, Jr.
Los Alamos Scientific Laboratory
P. O. Box 1663
Los Alamos, New Mexico

telephone ext. 2-3288

The above message is on OFFICIAL BUSINESS and is necessary for performance of Contract W-7405 eng. 36. The message to be transmitted cannot be performed by mail and is being sent in this manner in the interest of the work of the project.

G.Gratta

APPROVED..... DATE 6-26-53.....

Modern detectors are
 $\sim 10^5$ times larger !

Example: KamLAND

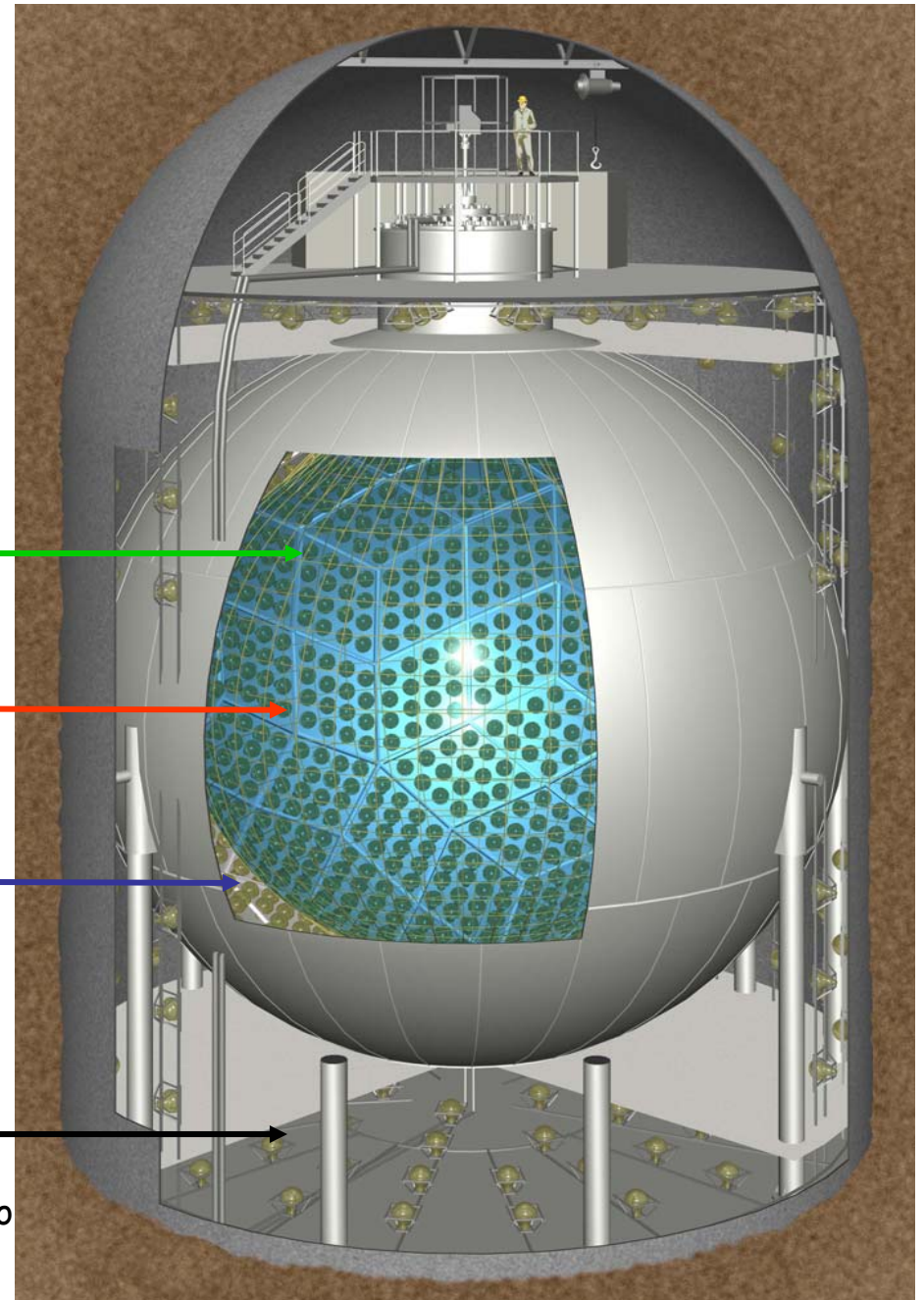
~ 2000 20" PMTs

1 kton liquid-scintillator

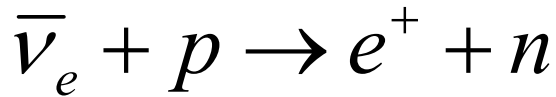
2.5 m-thick paraffin shielding

Water shield/Cherenkov veto

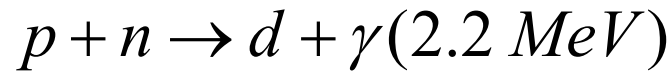
1km  overburden



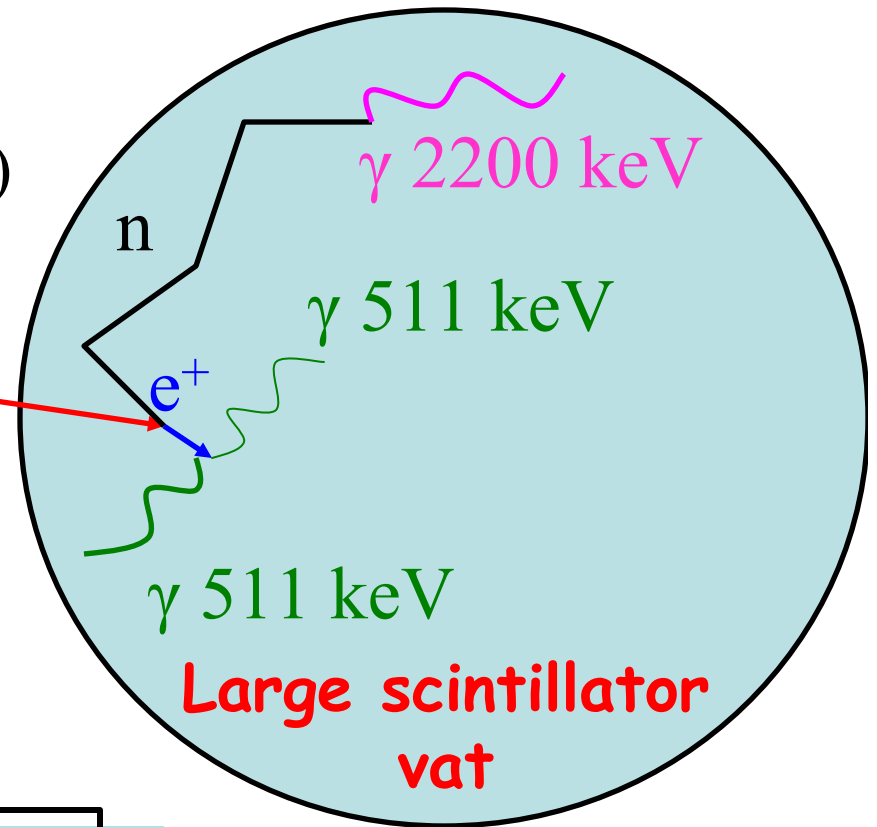
$\bar{\nu}_e$ are handy: can use inverse to suppress backgrounds



$\tau \approx 200 \mu s$



Event tagging by delayed coincidence
in **energy**, **time** and **space**



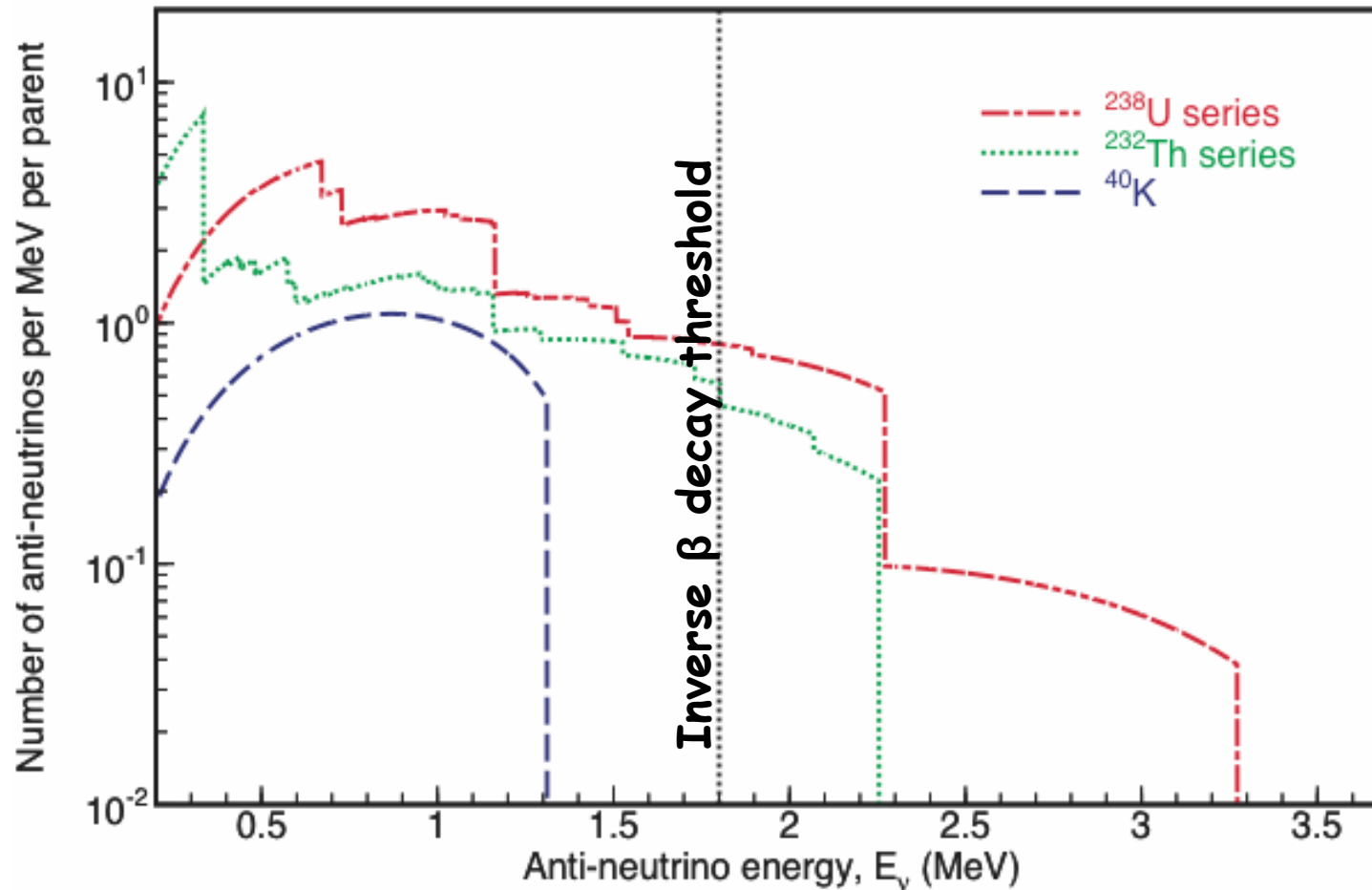
10-40 keV

800 MeV

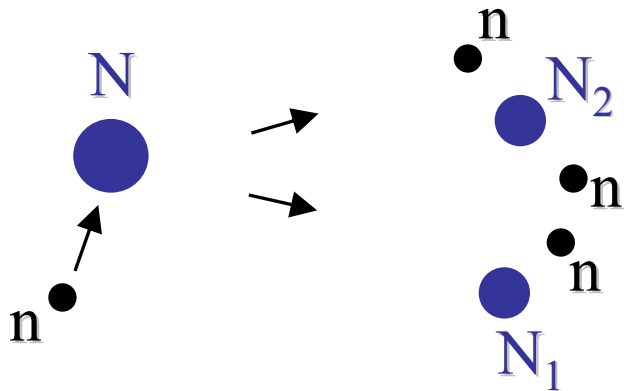
$$E_{\bar{\nu}} \cong E_{e^+} + E_n + (M_n - M_p) + m_{e^+}$$

E_{ν} measurement

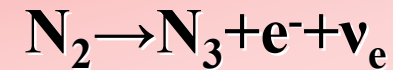
So there is a threshold for detecting neutrinos this way and ^{40}K $\bar{\nu}_e$ are all below threshold



Nuclear reactors are very intense sources of $\bar{\nu}_e$ deriving from beta-decay of the neutron-rich fission fragments



N_1 and N_2 still have too many neutrons and decay



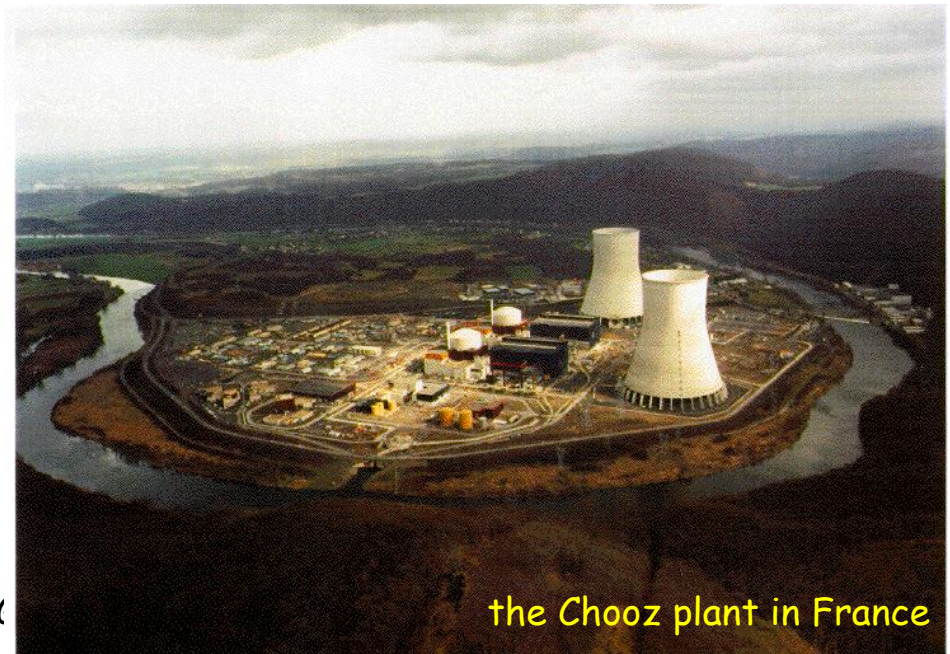
$200\text{MeV} / \text{fission}$

$6\bar{\nu}_e / \text{fission}$

A typical large power reactor produces
3 GW_{thermal} and
 $6 \cdot 10^{20}$ antineutrinos/s

G.Gratta

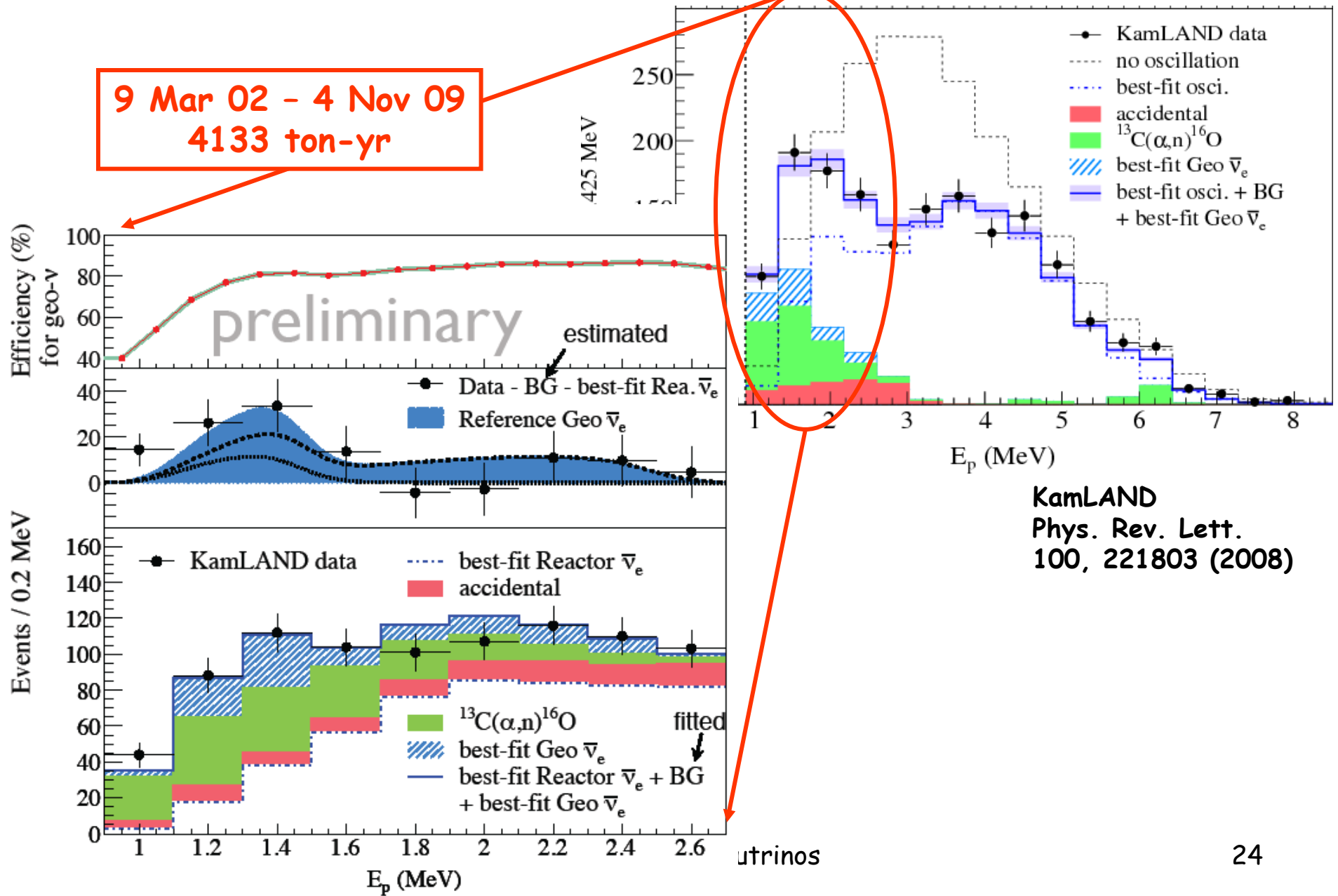
SSI 2010, (

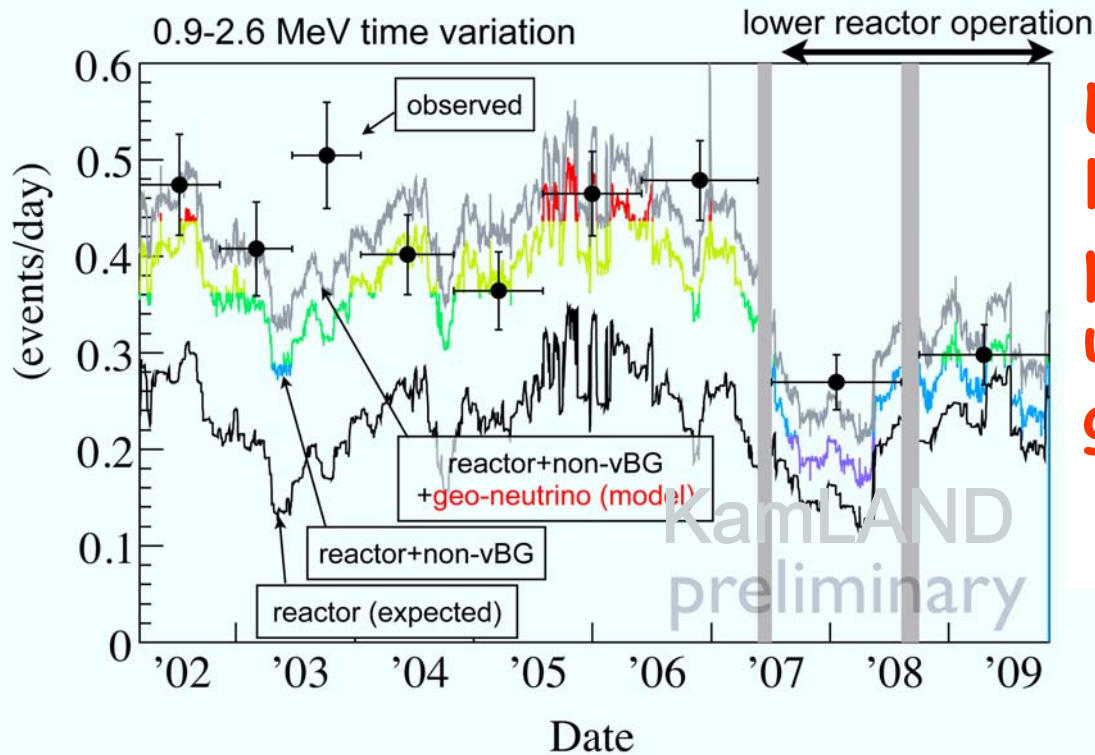


the Chooz plant in France

How does data look like? (KamLAND)

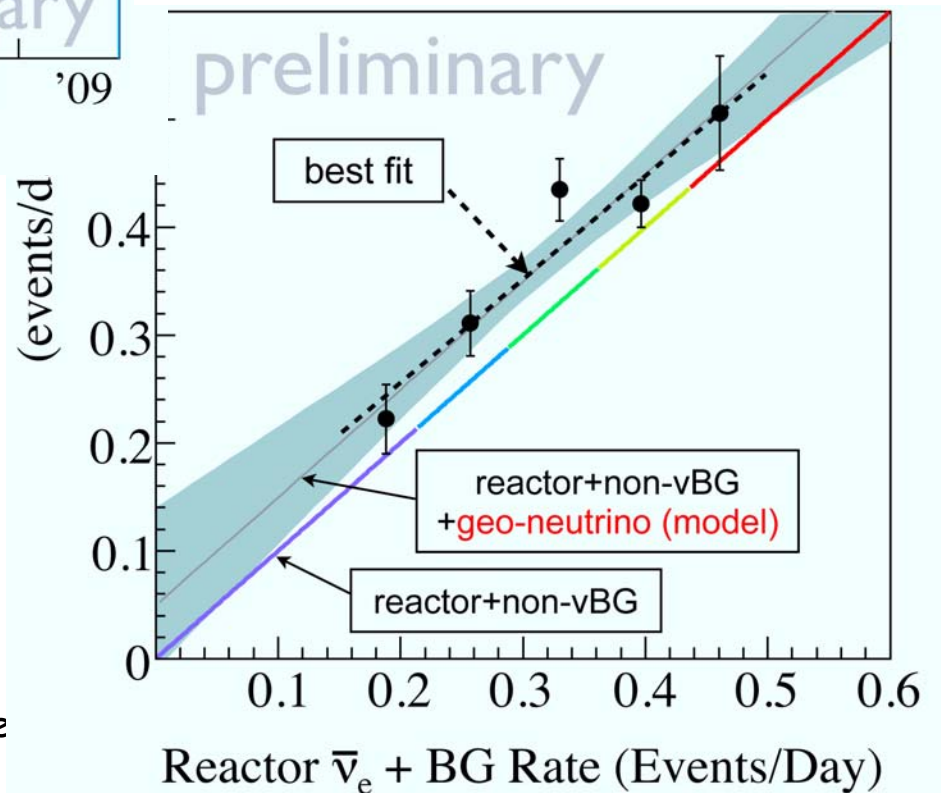
9 Mar 02 - 4 Nov 09
4133 ton-yr



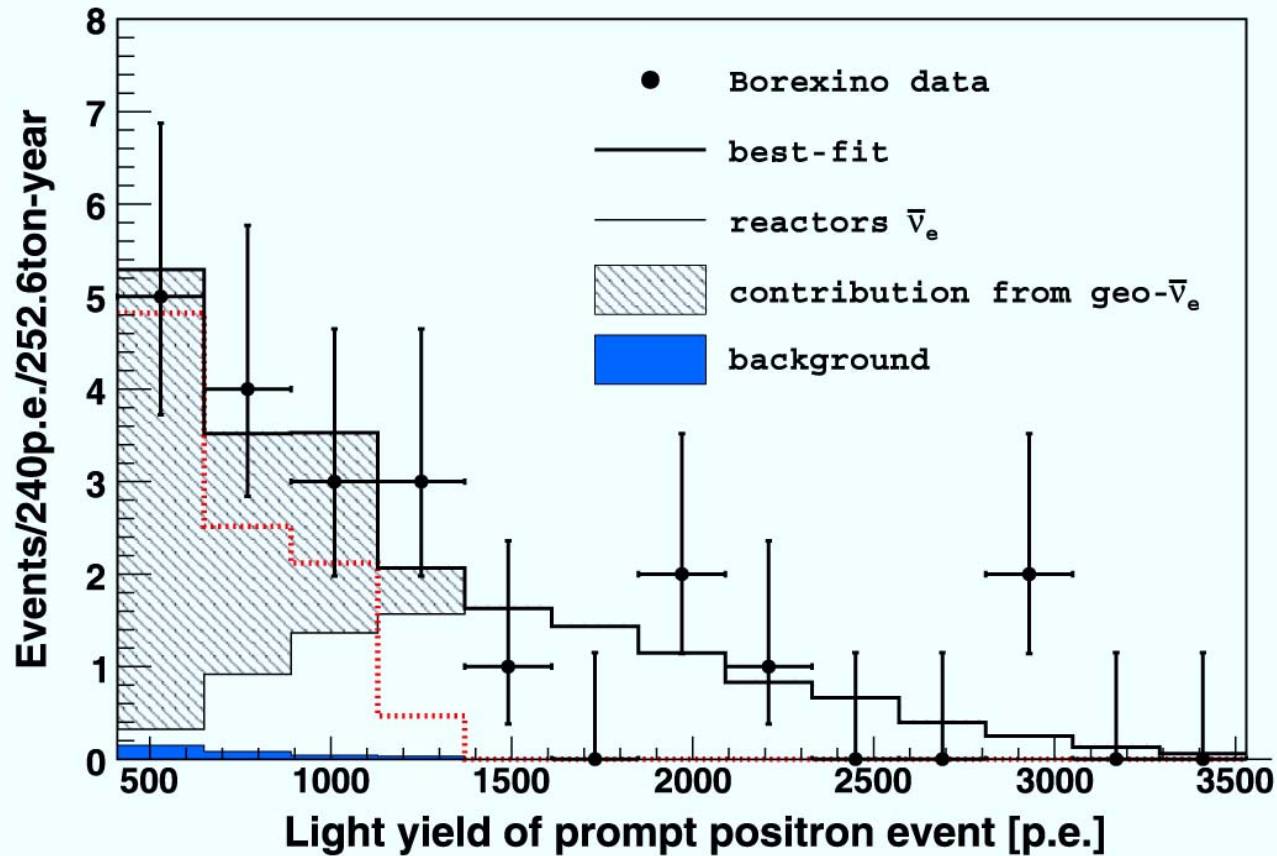


Even in presence of large reactor background power excursions can be used to separate the geoneutrino signal

Null hypothesis excluded at 99.997%CL



How does data look like? (Borexino)



Null hypothesis
excluded at
99.95%CL

- Borexino is ~3x smaller and run shorter time:
exposure is 253 ton-yr (~6% of KamLAND)
- Borexino is substantially cleaner than KamLAND and they have very little "background" from nuclear reactors

The geo-neutrino flux due to a particular isotope accumulated at a particular distance \vec{L} from the detector can be calculated as

$$\frac{d\Phi}{dE} \approx A \frac{dn}{dE} \Delta(E_\nu, \vec{L}) \int_{\oplus} d\vec{L} \frac{a(\vec{L})}{4\pi|\vec{L}|^2}$$

Where:

- A is the decay rate per unit mass for the isotope decay chain
- dn/dE is the anti-neutrino energy spectrum for the decay chain
- $\Delta(E_\nu, \vec{L})$ is a correction accounting for neutrino oscillations
- $a(\vec{L})$ is the amount of isotope at the distance

So from the measurement of $d\Phi/dE$ the amount of isotope $a(\vec{L})$ can be extracted by inverting this relationship above

The integral over the volume of the Earth introduces substantial degeneracy for a measurement done at a single site.

So, for the time being we'll use the flux measured at 2 sites to test models of U and Th distributions

A number of models have been developed over the last few years:

- *R. S. Raghavan et al. Phys. Rev. Lett. 80 (1998) 635*
- *C. Rothschild et al. Geo. Res. Lett. 25 (1998) 1083*
- *F. Mantovani et al. Phys. Rev. D 69 (2004) 013001*
- *G. Fiorentini et al. Phys. Rev. D 72 (2005) 033017*
- *K.A. Hochmuth, Prog. Part. Nucl. Phys. 57 (2006) 293*
- *G.L. Fogli et al. Earth, Moon, and Planets 99 (2007) 111*
- *S. Enomoto et al. Earth Planet. Sci. Lett. 258 (2007) 147*

Aside from the “interesting” part of estimating the global distributions of U and Th in core, mantle, continental and oceanic crusts, models require a more mundane, hi rez description of the distributions in the nearby crust
(because of the $1/L^2$ in the integral)

The expected rate at different world locations

Geoneutrino Event Rate (Crust+Mantle)

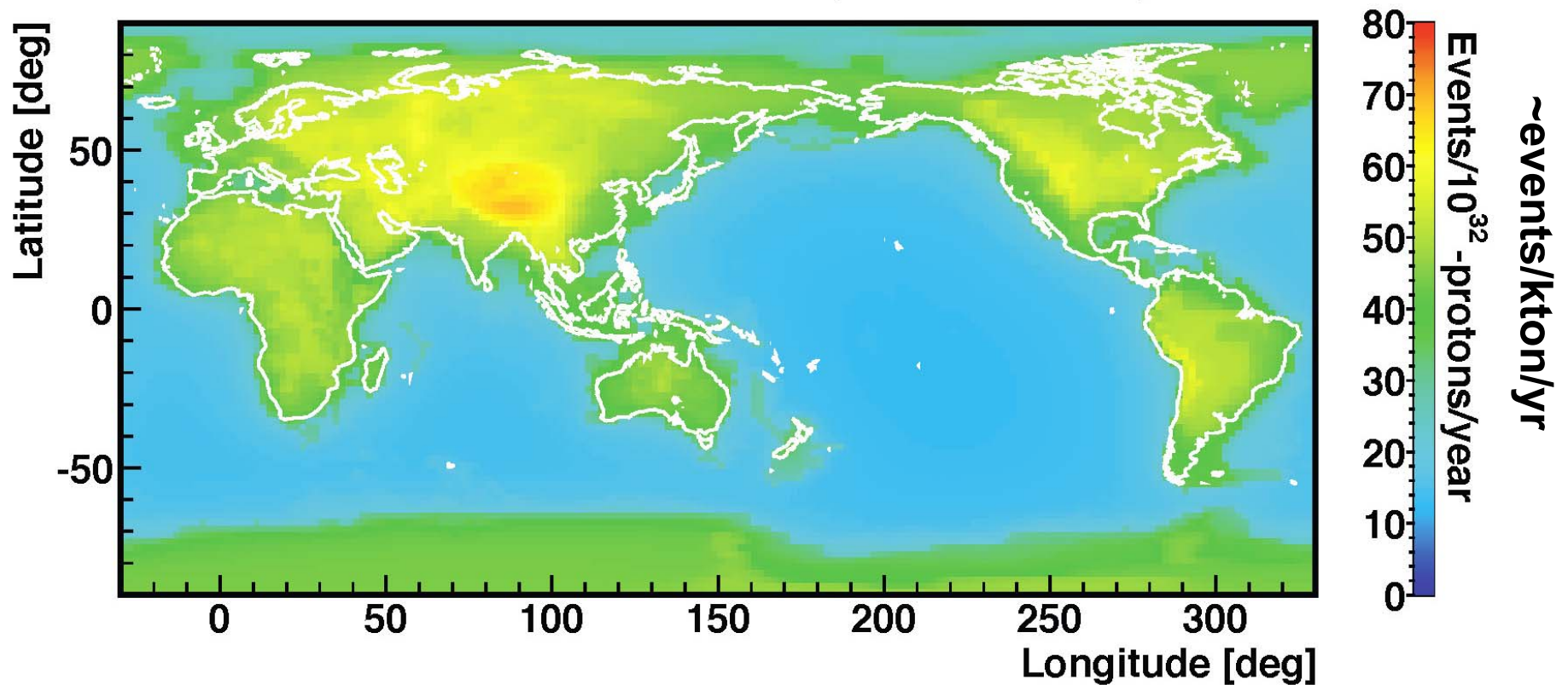


Image: S. Enomoto

Note the rate scale:

in most places expect 4 events/month in a 1kton detector!

Unfortunately where underground labs are this rate is mostly due to the crust

Crust is less interesting as it can be sampled directly

Geoneutrino Event Rate (Crust)

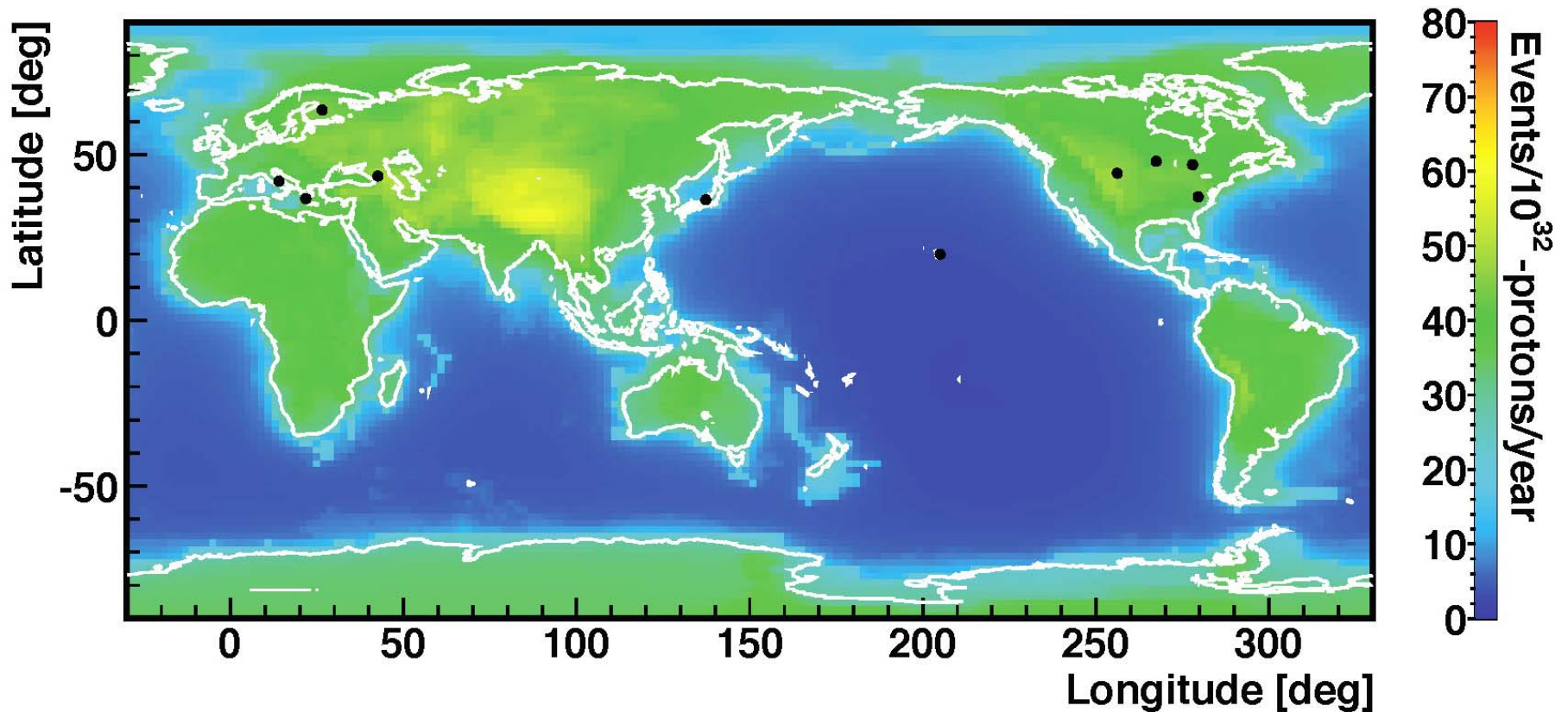
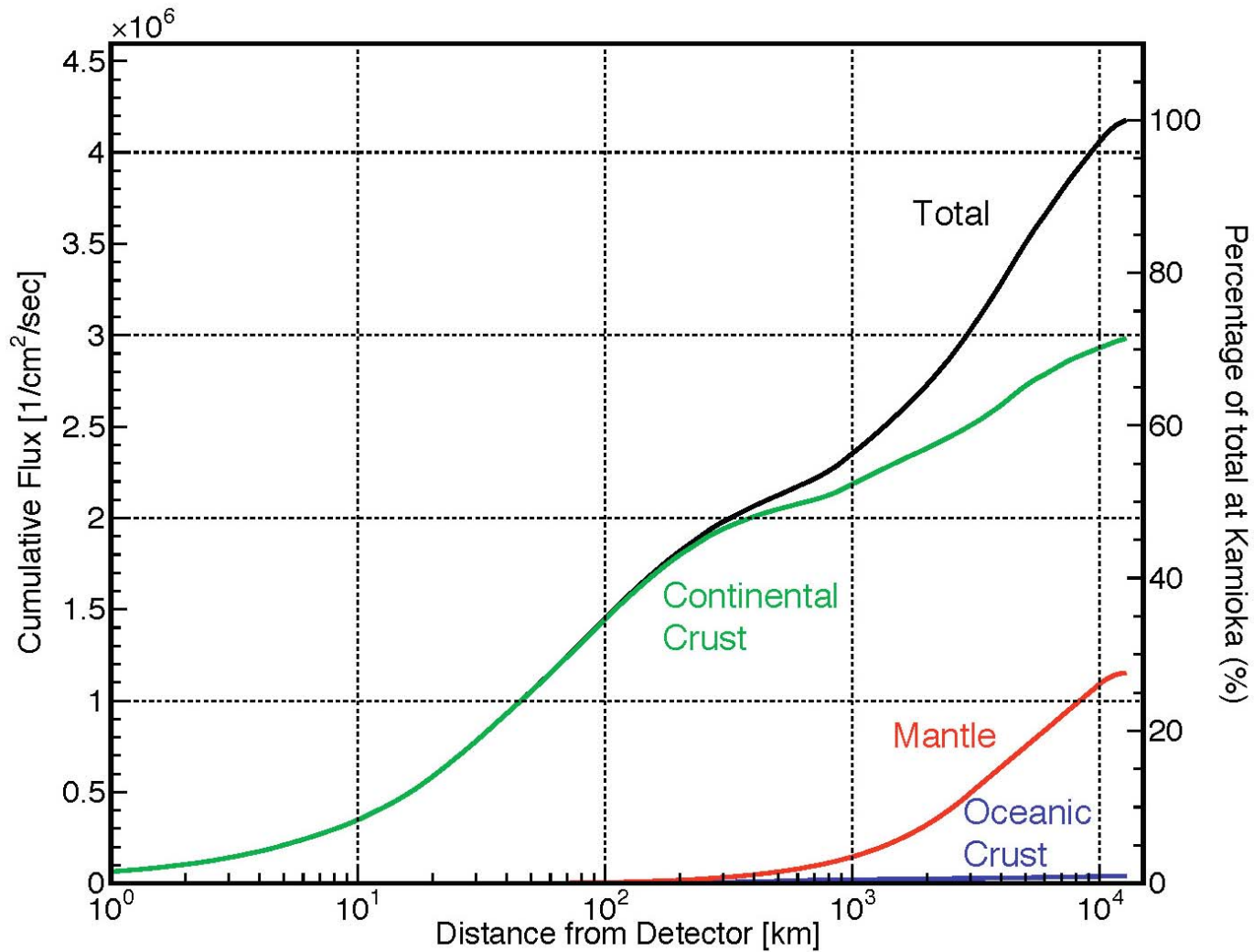


Image: S. Enomoto

The breakdown of various contributions at the Kamioka site



Background from reactors.

Note that in many locations this is a severe problem

Reactor Neutrino Event Rate ($1.8\text{MeV} < E < 3.3\text{MeV}$)

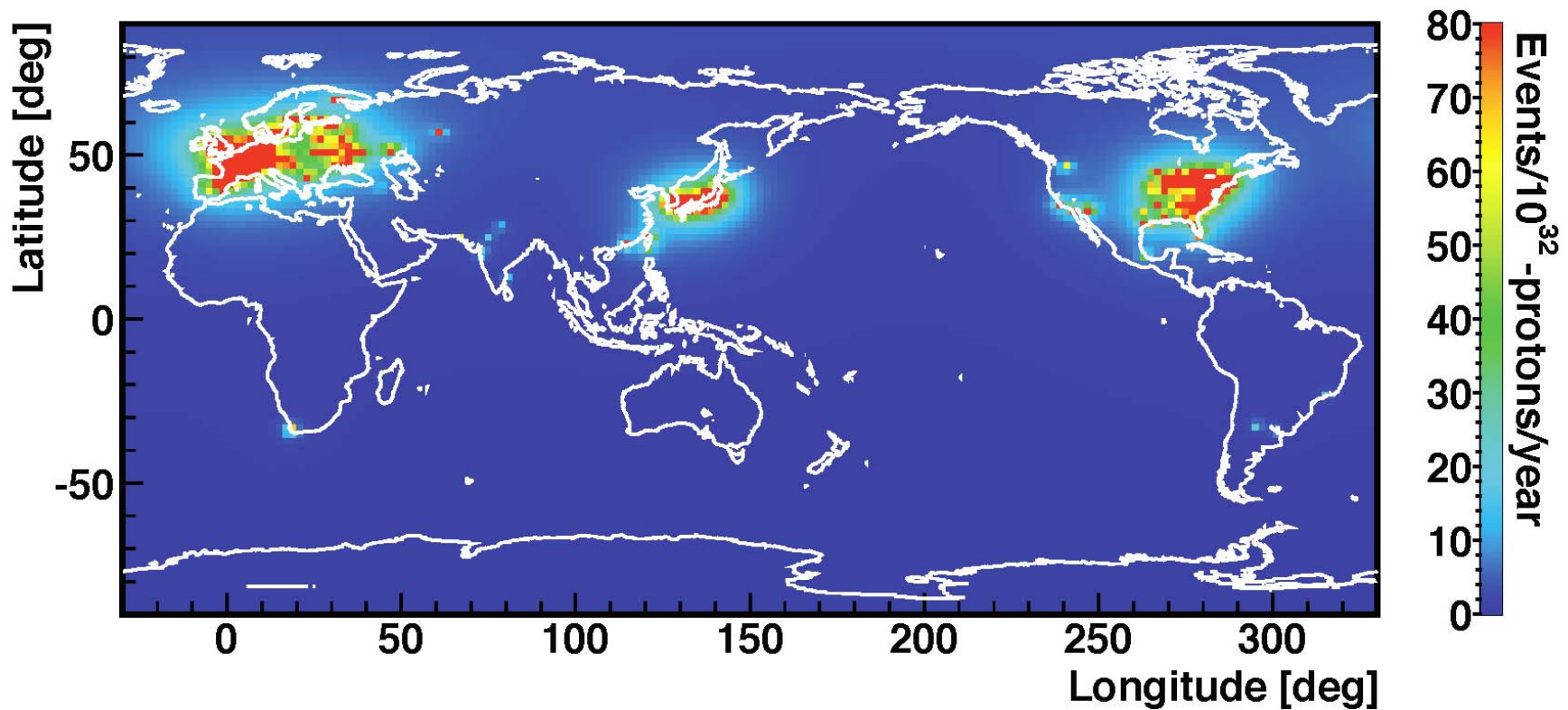


Image: S. Enomoto

The ideal location to study the Earth's mantle is the middle of an ocean, where there are no reactors and the crust is thinnest and depleted of Th & U

S/N Ratio: Mantle / (Crust + Reactor)

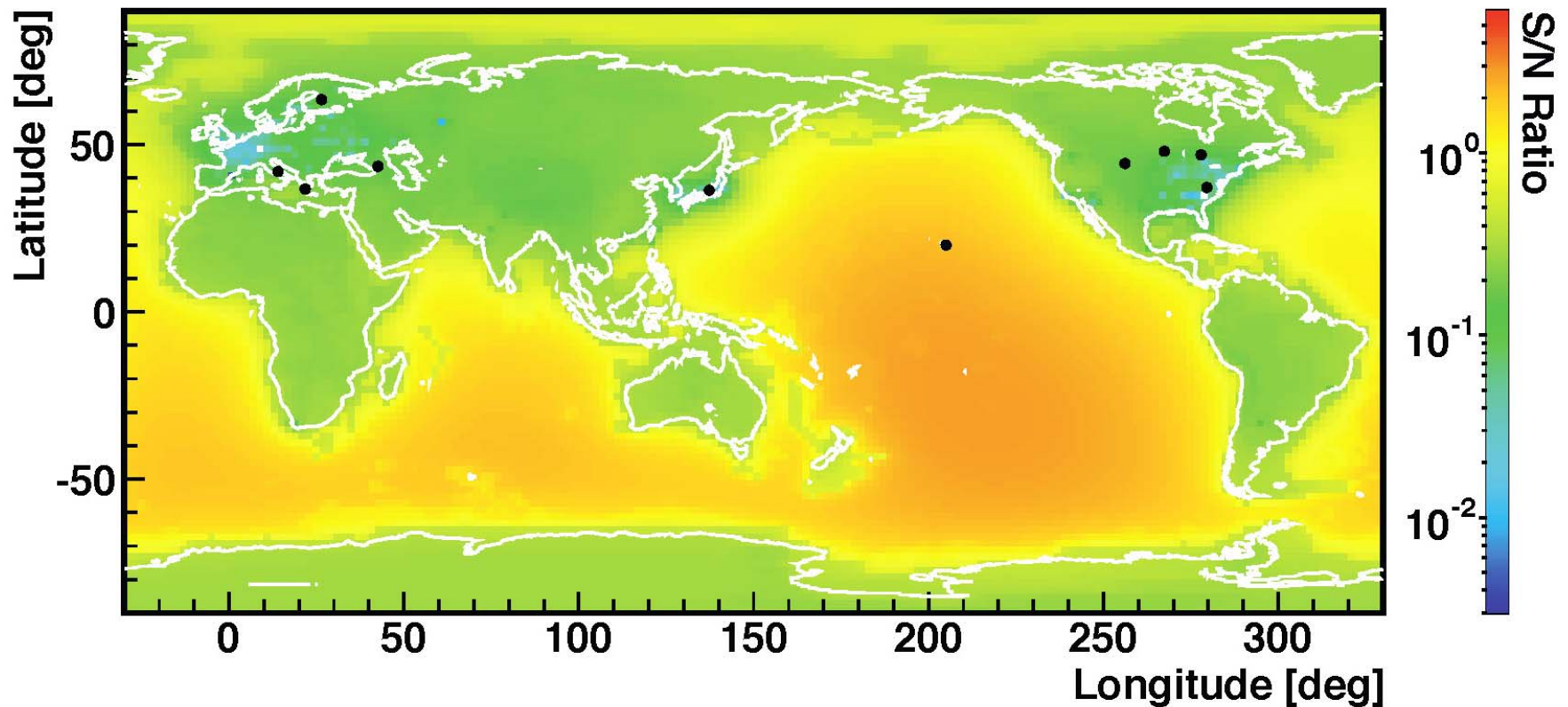
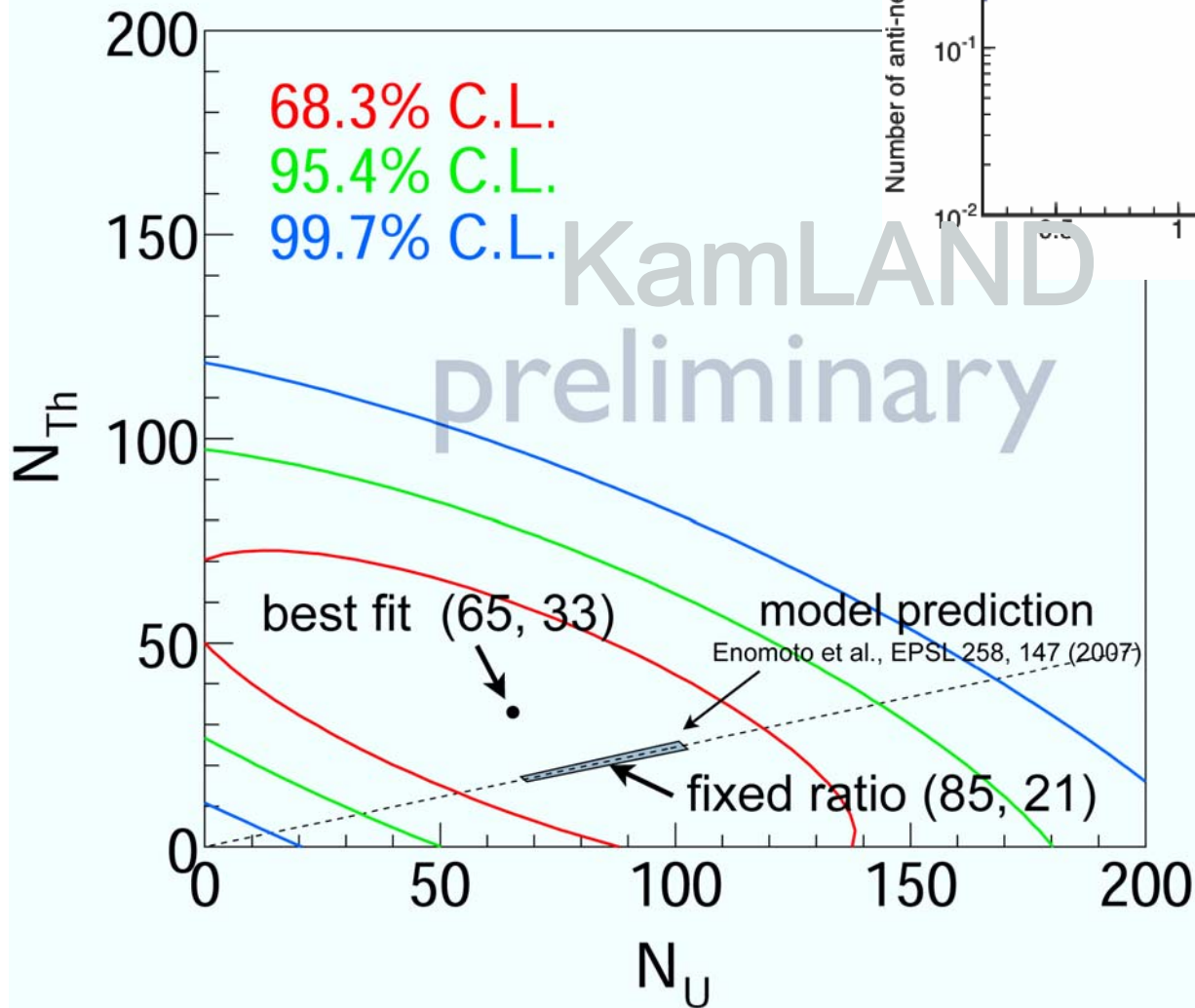
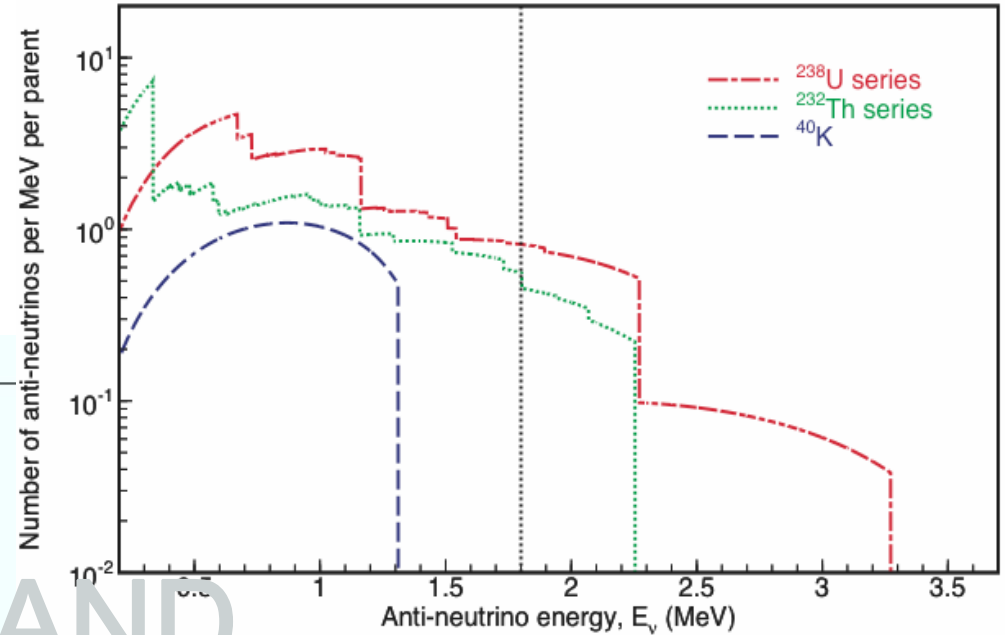


Image: S. Enomoto

Flux [$\times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$]

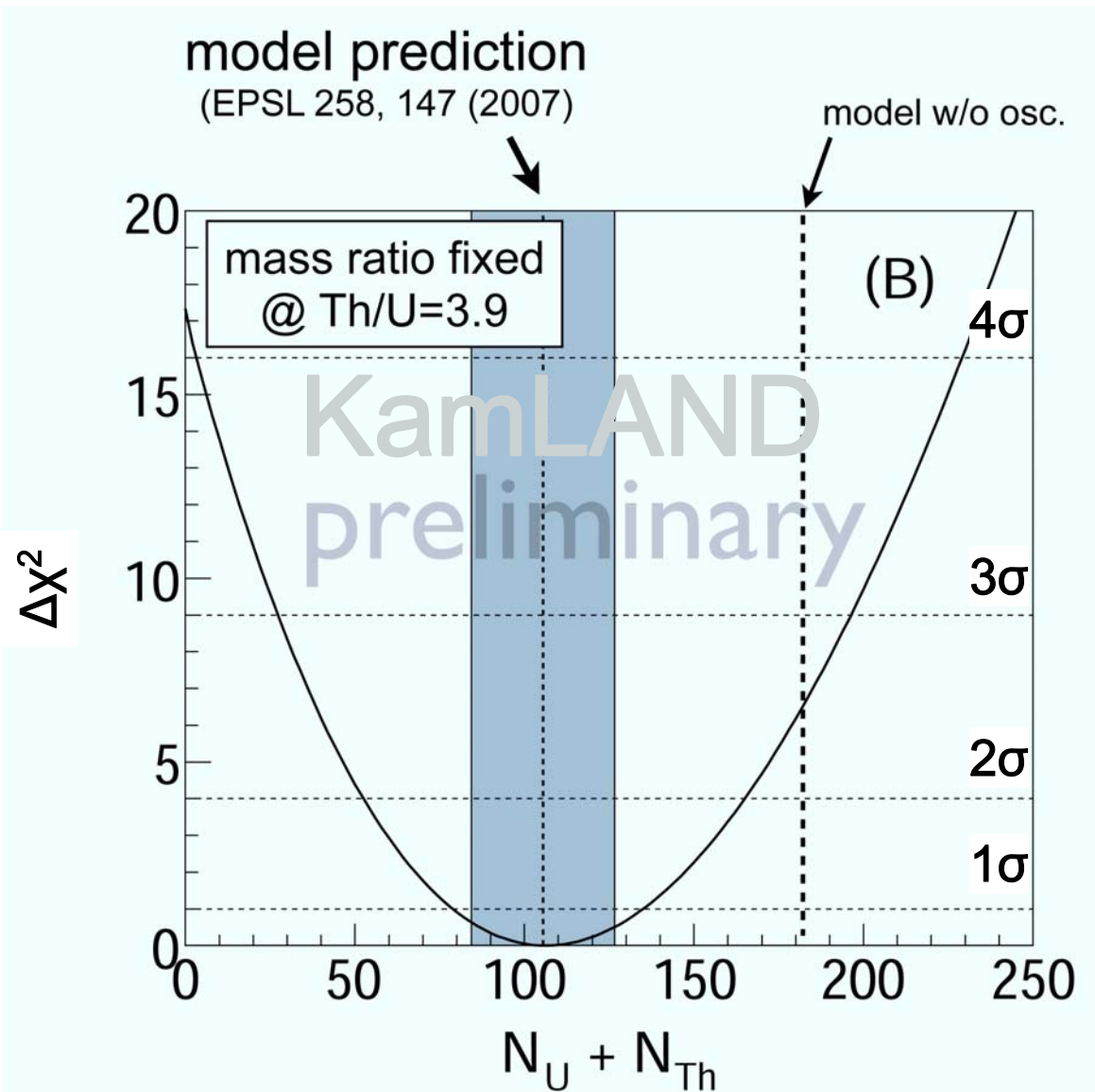
	Measured		Model (Enomoto / Mantovani)
KamLAND 2005*	$6.4^{+3.6}_{-3.4}$	Nature <u>436</u> , 499	4.4 / 4.0
KamLAND 2008	4.4 ± 1.6	PRL <u>100</u> , 221803	4.4 / 4.0
Borexino 2010*	$7.1^{+2.9}_{-2.4}$	Phys Lett B <u>687</u> , 299	5.2 / 4.6
KamLAND 2010	$4.3^{+1.2}_{-1.1}$	Preliminary, Neutrino 2010	4.4 / 4.0

Data is still pretty insensitive to the U/Th ratio

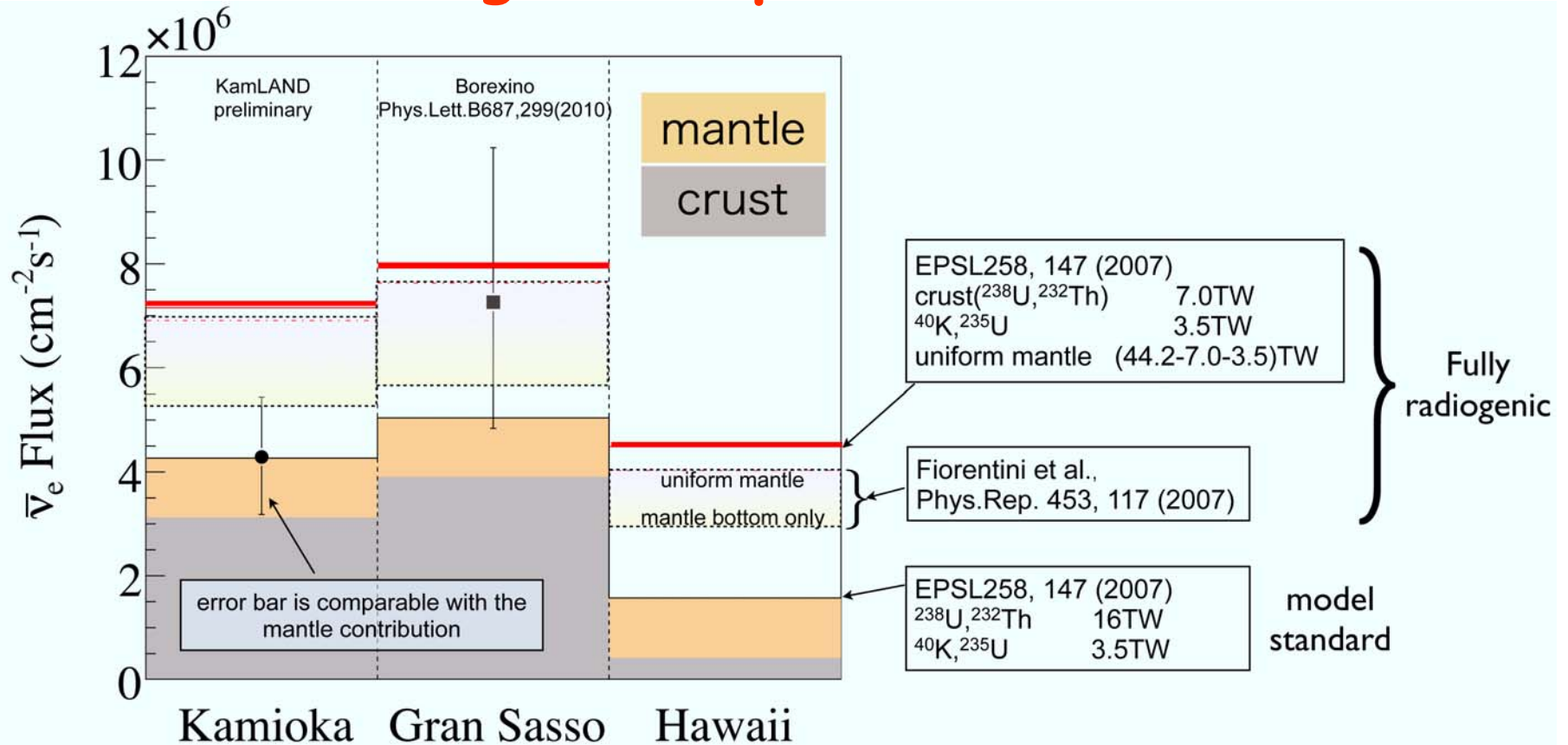


KamLAND preliminary

→ Take ratio from model
(basically from Chontritic Meteorites)



Summarizing the experimental situation



- Mantle contribution: 1σ contribution in KamLAND data
- KamLAND data $\sim 2\sigma$ away from fully radiogenic

Is there a large nuclear reactor inside the Earth?



While natural reactors are known to have existed on Earth a natural reactor today presents a number of challenges because of the low concentration of ^{235}U

Nevertheless large natural reactors have been proposed

- In the core

J.M. Herndon, PNAS USA
100, 3047 (2003)

- Near core-mantle boundary

V.D. Rusov et al. arXiv:0902.4092
R.J. deMeijer et al.

Radiat. Phys. Chem. 71, 769 (2004)



Part of the Oklo fossil reactor

A power of 5-10TW from fission may help explain heating, convection, ^3He anomaly and a number of other curiosities

Both models are strongly disfavored by geochemists

The evidence for/against such a reactor is indirect and clearly controversial.

The detection of an excess of anti-neutrinos with reactor-like spectrum (hence at higher energy than that of geoneutrinos) would provide solid evidence

In this case Borexino is clearly in a better situation because of the low reactor background.

Both experiments have set limits that start probing the interesting regime

KamLAND $P_{\text{reactor}} < 6.2 \text{ TW at } 90\% \text{ CL}$
Phys.Rev.Lett. 100, 221803 (2008)

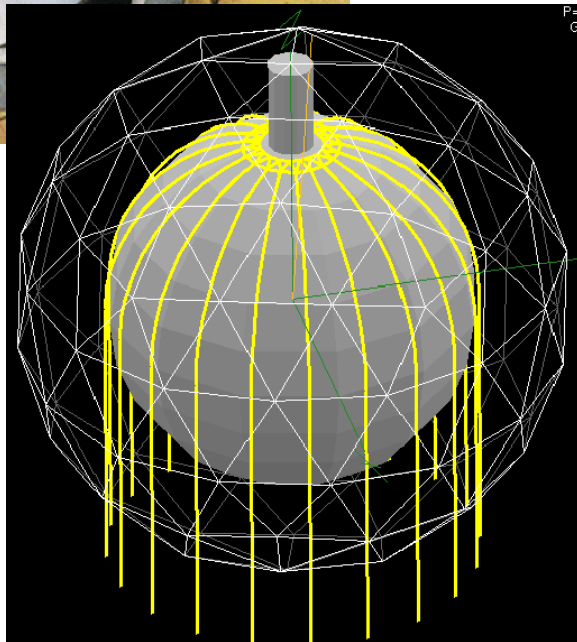
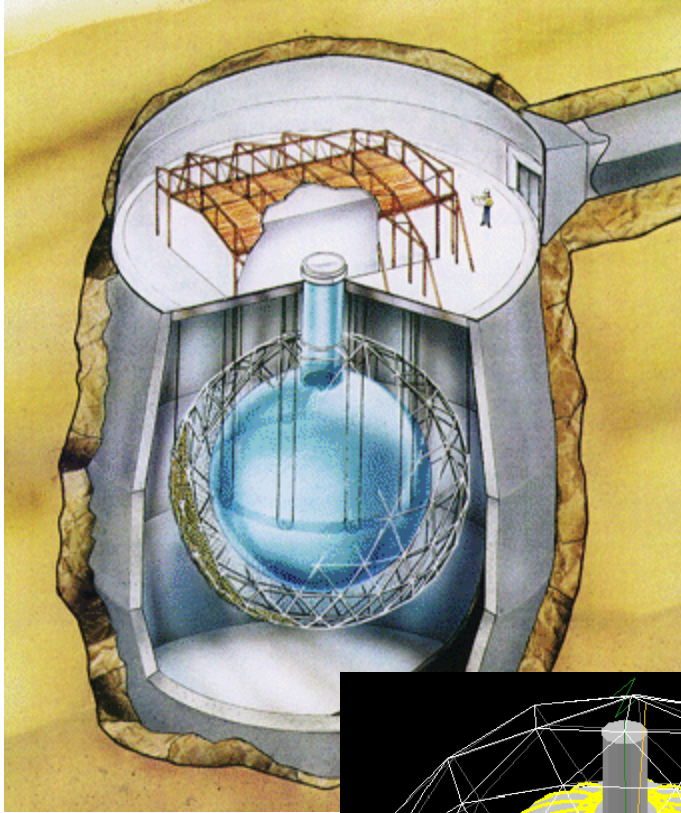
Borexino $P_{\text{reactor}} < 3 \text{ TW at } 95\% \text{ CL}$
Phys.Lett. B687, 299 (2010)

Wish list for future geoneutrino detectors:

- Better statistics (larger detectors)
- Multiple sites
- Oceanic site (this is very challenging but also very important)
- Pointing ability (very good position resolution)
- Lower threshold (For a number of reasons K concentration in mantle very important for geophysics)

The near future: SNO+

Some more remote future: LENA



~1kton
SNO site

G.Gratta

DETECTOR LAYOUT

Cavern

height: 115 m, diameter: 50 m
shielding from cosmic rays: ~4,000 m.w

Muon Veto

plastic scintillator panels (on top)
Water Cherenkov Detector
1,500 phototubes
100 kt of water
reduction of fast neutron background

Steel Cylinder

height: 100 m, diameter: 30 m
70 kt of organic liquid
13,500 phototubes

Buffer

thickness: 2 m
non-scintillating organic liquid
shielding external radioactivity

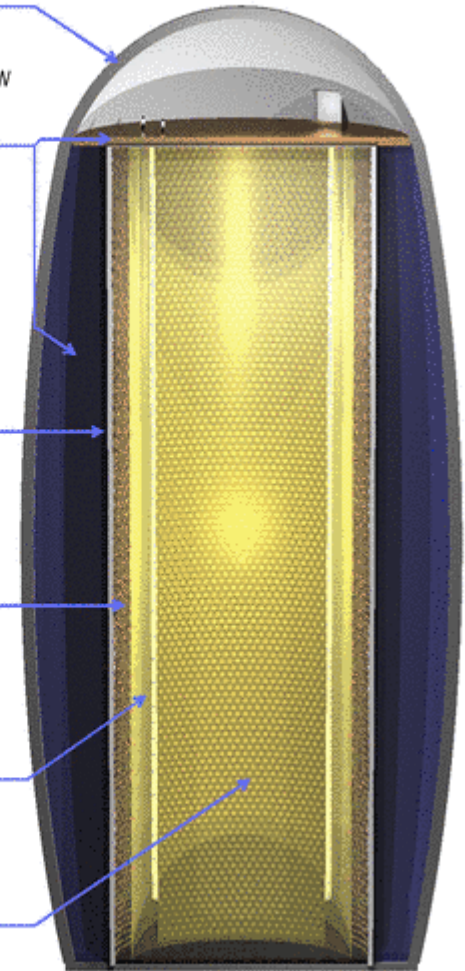
Nylon Vessel

parting buffer liquid
from liquid scintillator

Target Volume

height: 100 m, diameter: 26 m
50 kt of liquid scintillator

vertical design is favourable in terms of rock pressure and buoyancy forces

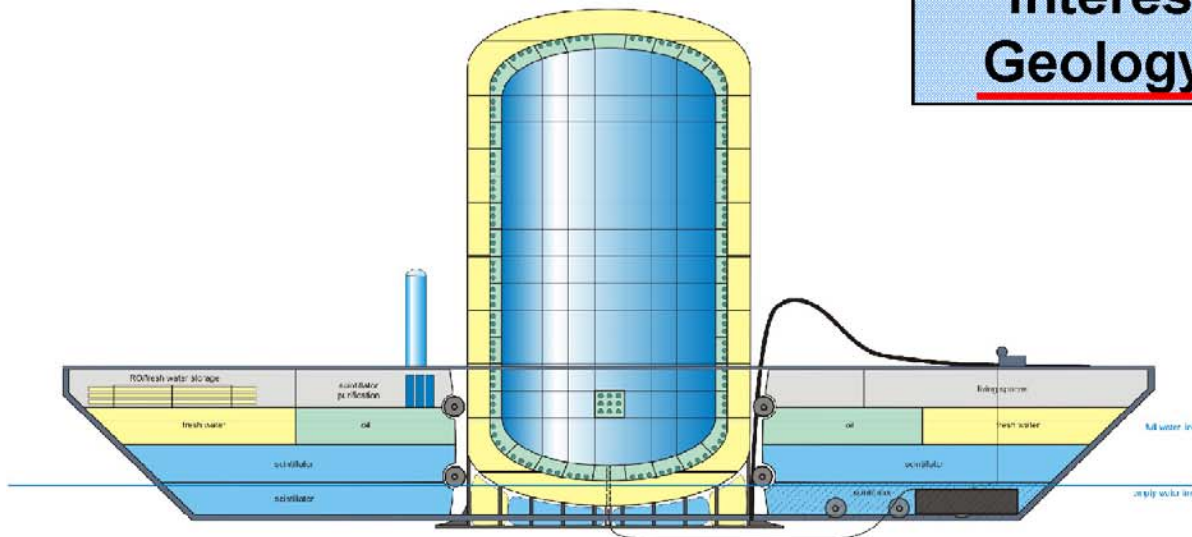


GeoNeutrinos

A 10kton dedicated detector that can be deployed in the ocean

Hanohano

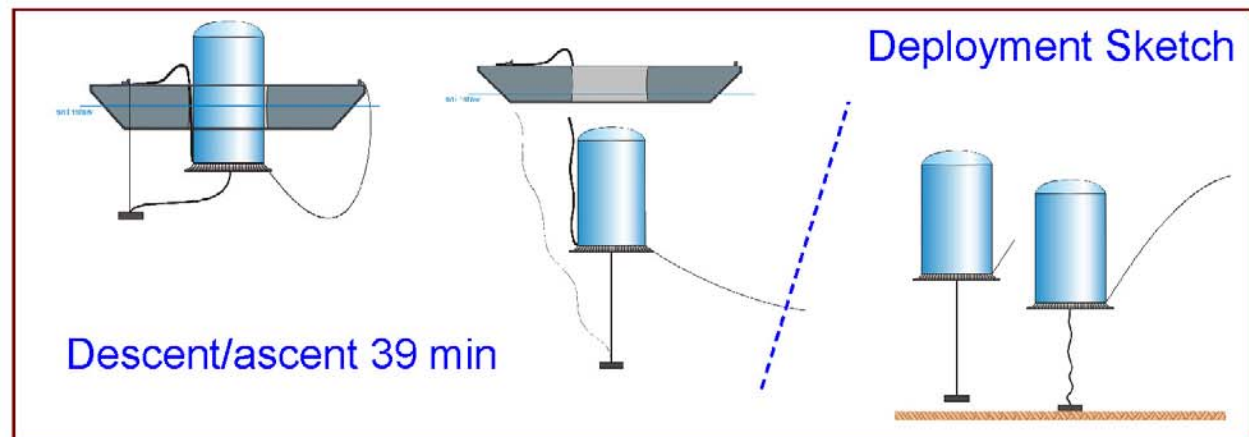
An experiment with joint interests in Physics, Geology, and Security



- multiple deployments
- deep water cosmic shield
- control-able L/E detection

Deep Ocean

$\bar{\nu}_e$ Observatory



Conclusions

- First example of “Applied Neutrino Physics” !!
- Clear detection of geoneutrinos by KamLAND and Borexino
- First chemical analysis of the mantle of the Earth
- Consistent with the current geological models
- With data coming from two continents start constraining different models
- SNO+ to join the club soon
- A network of detector and an oceanic detector would drastically advance the knowledge of our planet