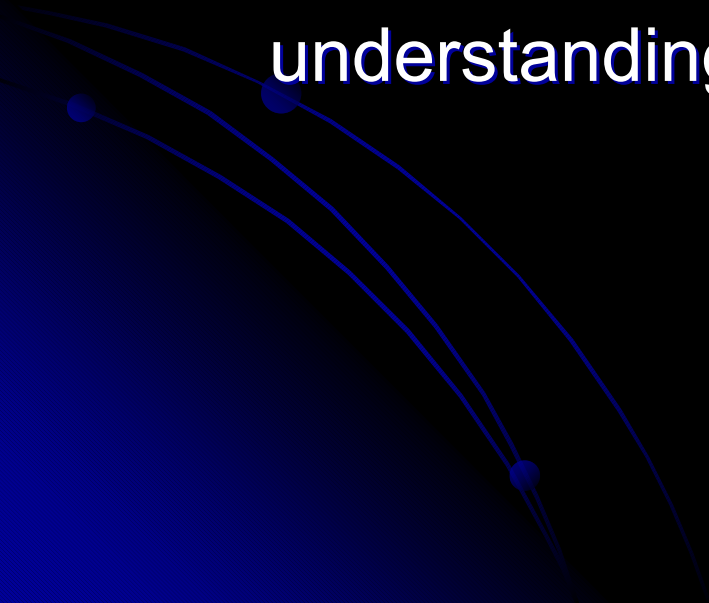


# Scattering X-Ray Polarimeters

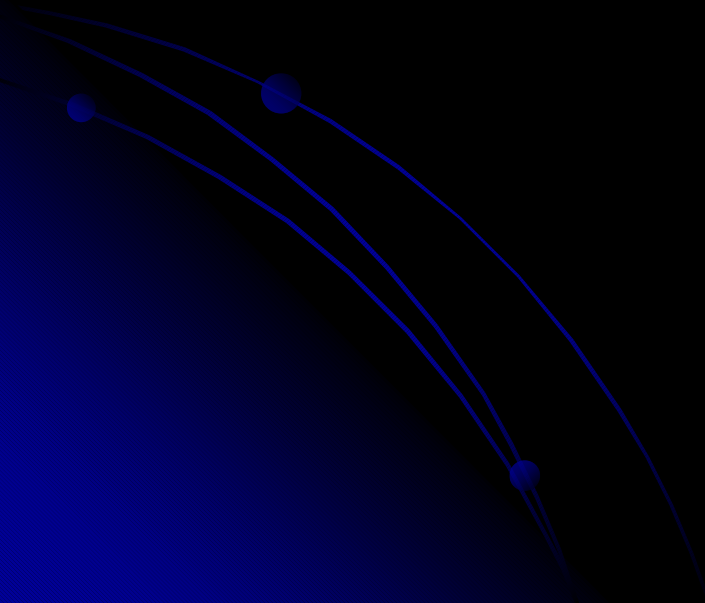
M.C.Weisskopf, R.F.Elsner, S.L. O'Dell &  
B.D.Ramsey

# Advantages

- Inherently broad band device
    - Astrophysical non-thermal spectra are characteristically broad band
    - Astrophysical diagnostics (model discriminators) may/should benefit from understanding the energy dependence
- 

# Basic Principles

- Based on the angular dependence of the incoherent (and coherent) scattering cross-section on the linear polarization of the incident photon



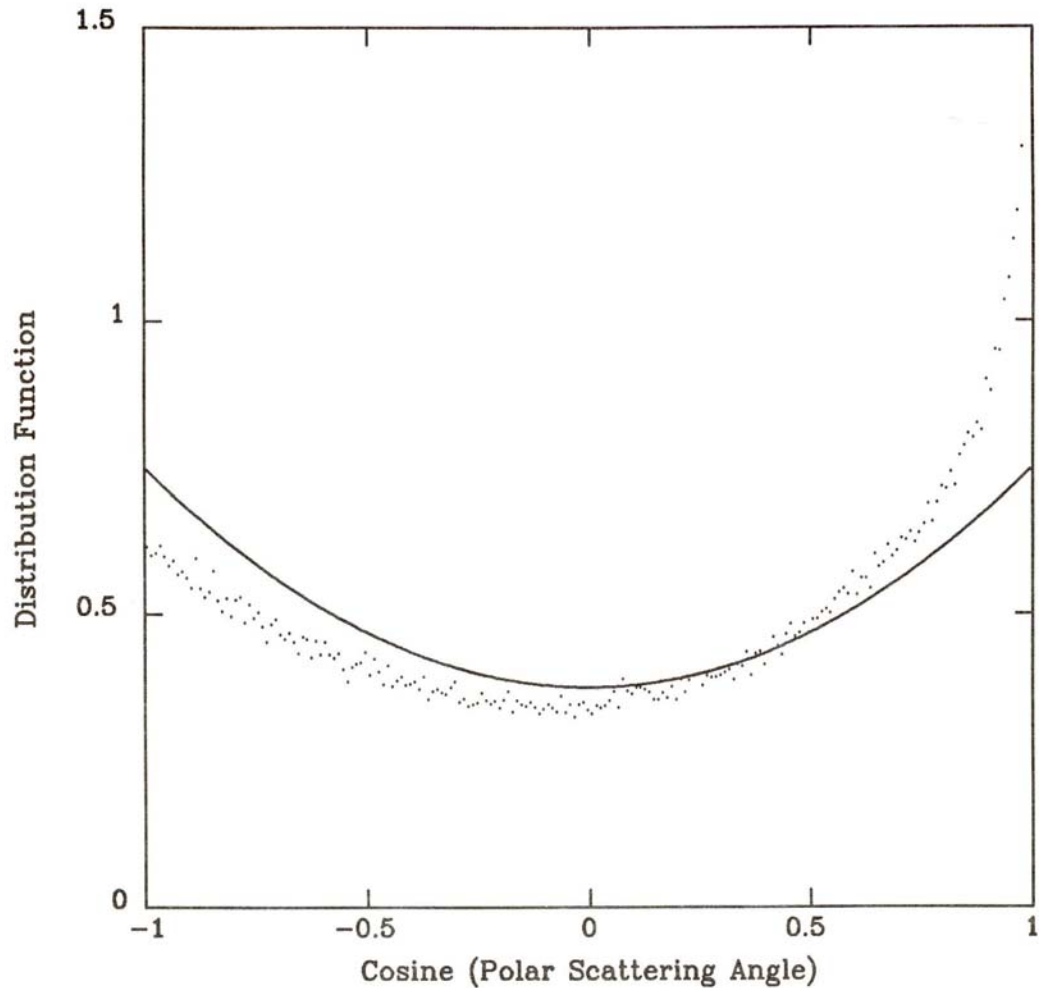
# Basic Principles

- Thomson cross-section illustrates the angular dependence

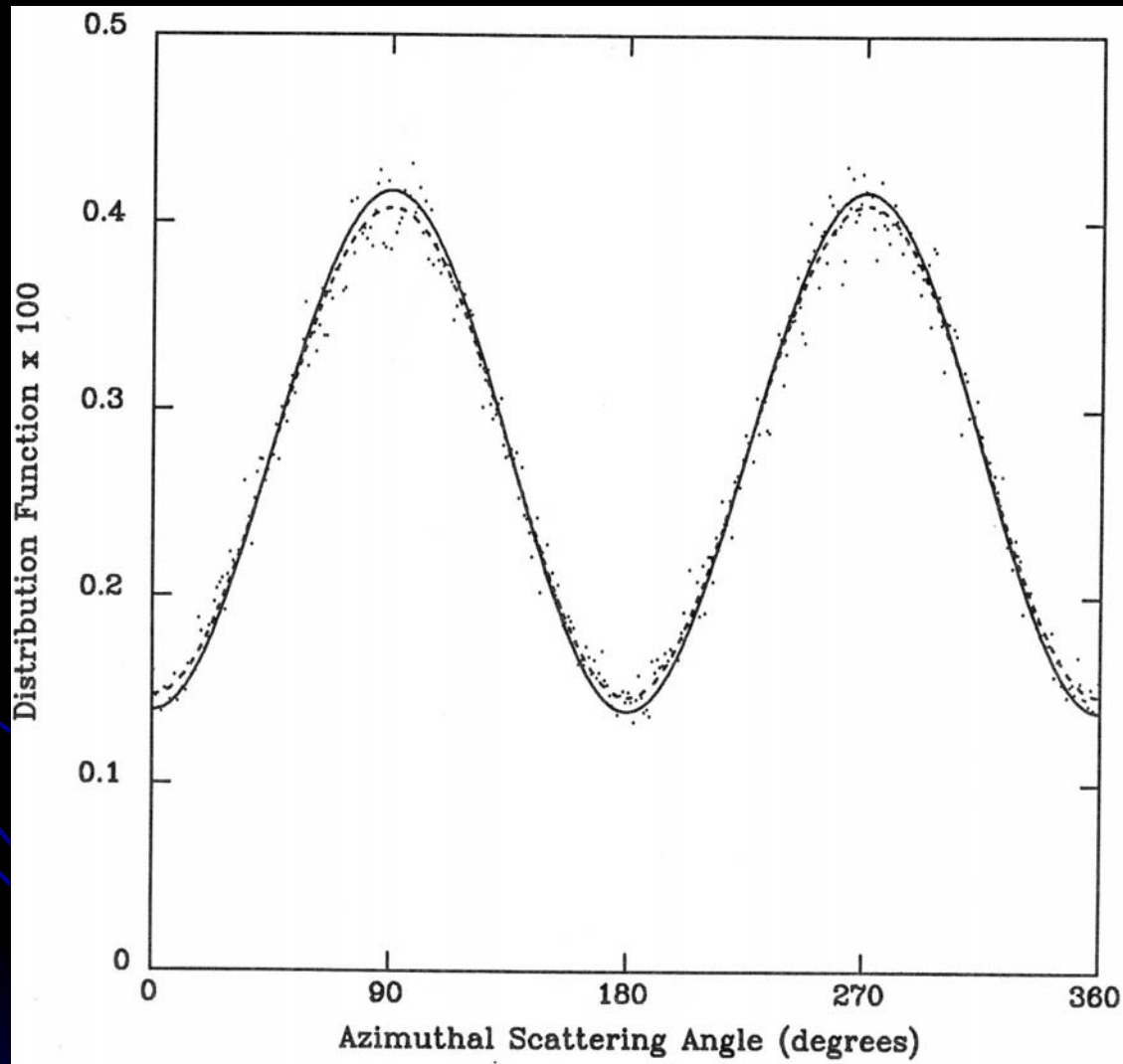
$$d\sigma / d\Omega = (e^2 / mc^2)(\cos^2 \theta \cos^2 \varphi + \sin^2 \varphi)$$

- For scattering from bound electrons one must account for both coherent and incoherent scattering and photoelectric absorption

# Thompson Approximation



# Thompson Approximation



# Considerations

- Scatter as much incident flux as possible
- Avoid multiple scattering
- Achieve as large an M as possible

$$MDP(\%) = (4.29 \times 10^4 / M(\%)) \sqrt{(R_S + R_B)} / \sqrt{R_S^2 t}$$

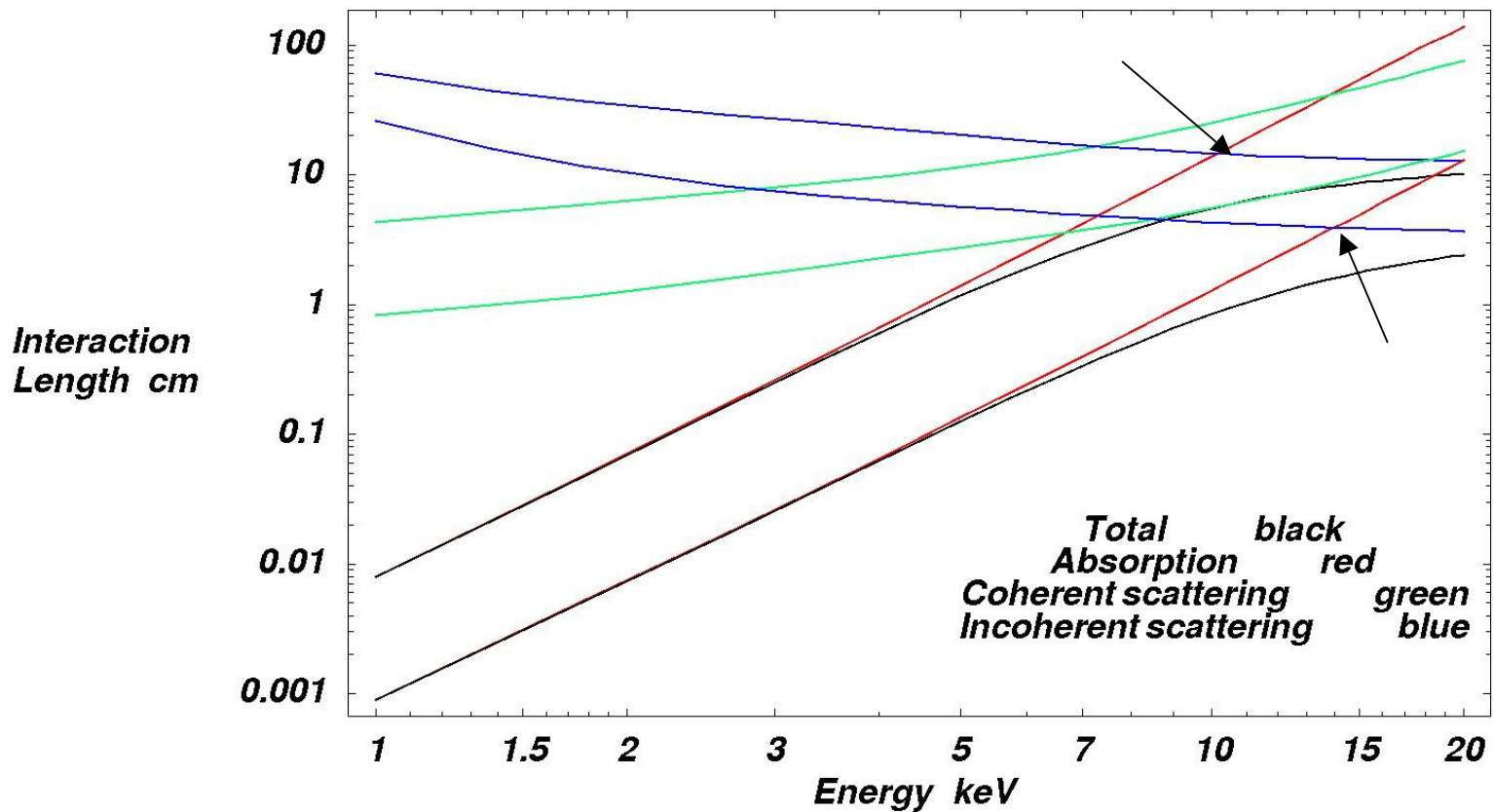
- Collect as many scattered X-Rays as possible
- Minimize the background

# The Polarimeter Conundrum

- The scattering material should be thick (deep) in order to effectively provide for interaction with all the incident photons.
- The scattering material should be thin (narrow) in order to allow the scattered photon to easily escape.
- (Similar conundrums apply as well to other approaches to polarimetry.)

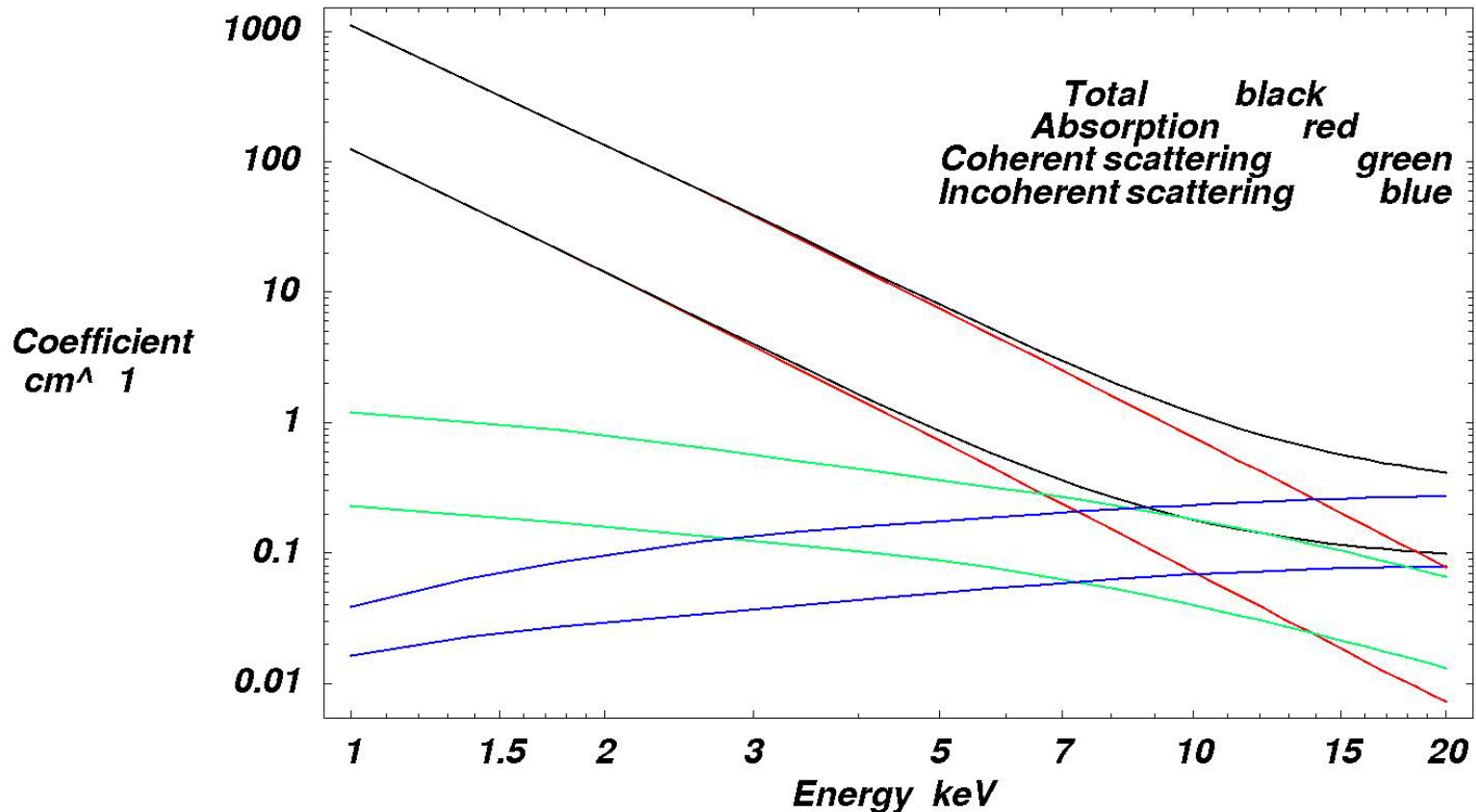
# Interaction Lengths

*Interaction Lengths for Lithium and Beryllium  
lower curves for beryllium and higher curves for lithium*

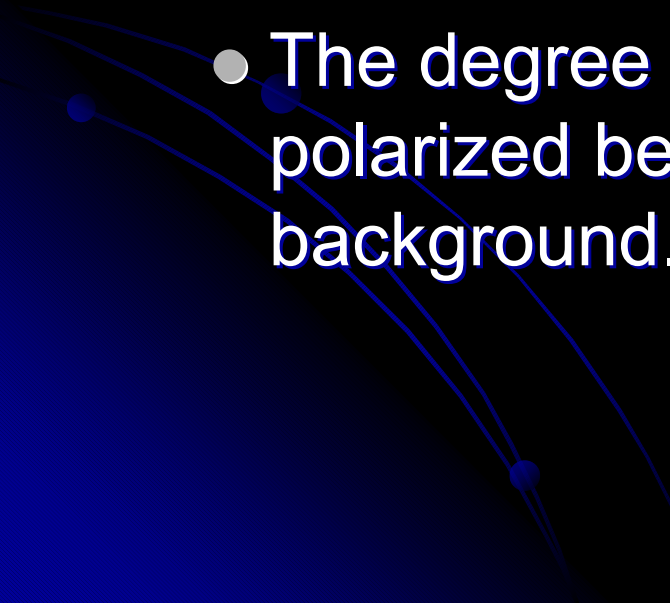


# Scattering Coefficients

*Linear Interaction Coefficients for Lithium and Beryllium  
higher curves for beryllium, lower curves for lithium*

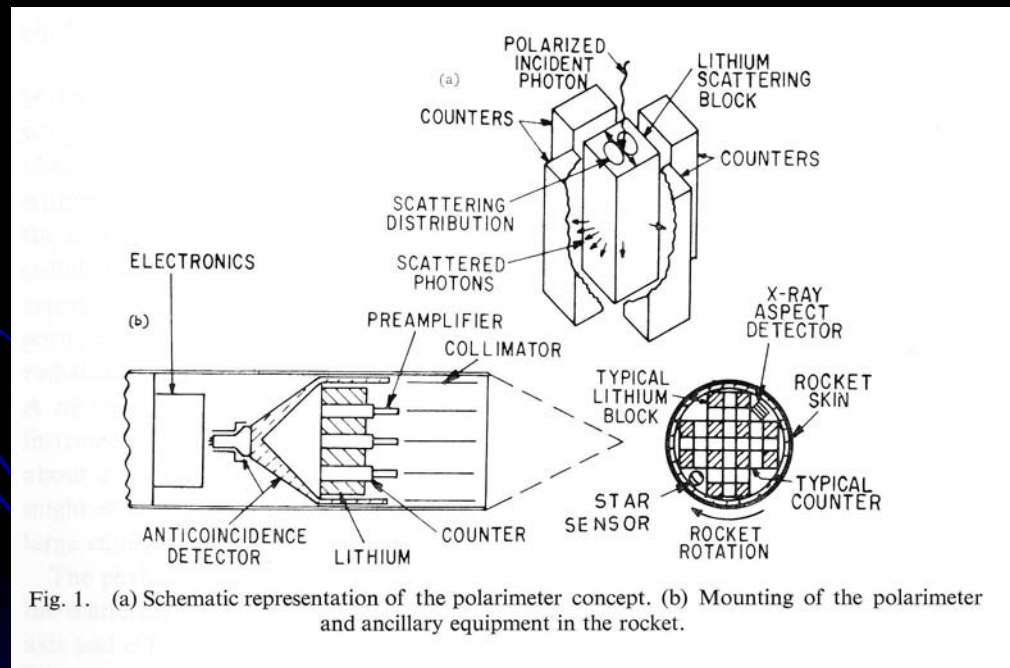


# Thus...

- Photo-electric absorption is a real nuisance!!
  - Multiple scattering needs to be minimized as this negatively impacts the “modulation factor”
    - The degree of modulation for a 100% polarized beam in the absence of any background.
- 

# In the beginning.....

- July 1968 – Lithium-block, Thomson-scattering polarimeter flown on an Aerobee -150 sounding rocket
  - Target was Sco X-1



# In the beginning....

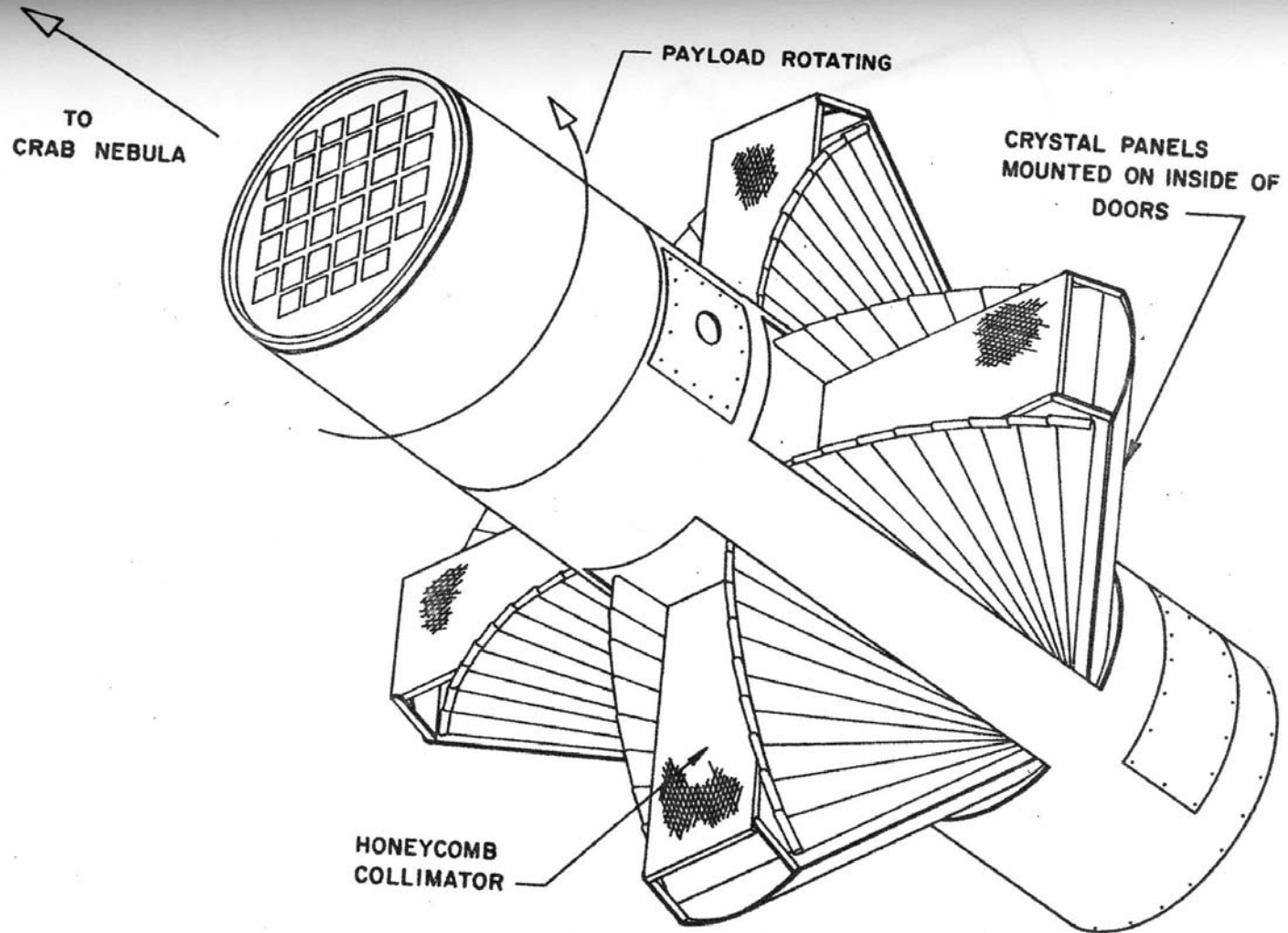
- March 1969 - Lithium-block, Thomson-scattering polarimeter flown on an Aerobee -150 sounding rocket
  - Target was the Crab Nebula
- February 1971 Lithium-block, Thomson-scattering polarimeter flown on an Aerobee -350 sounding rocket
  - Target was the Crab Nebula
  - (Included crystal polarimeter)

Three rockets in 21 months

# Rocket 17.09

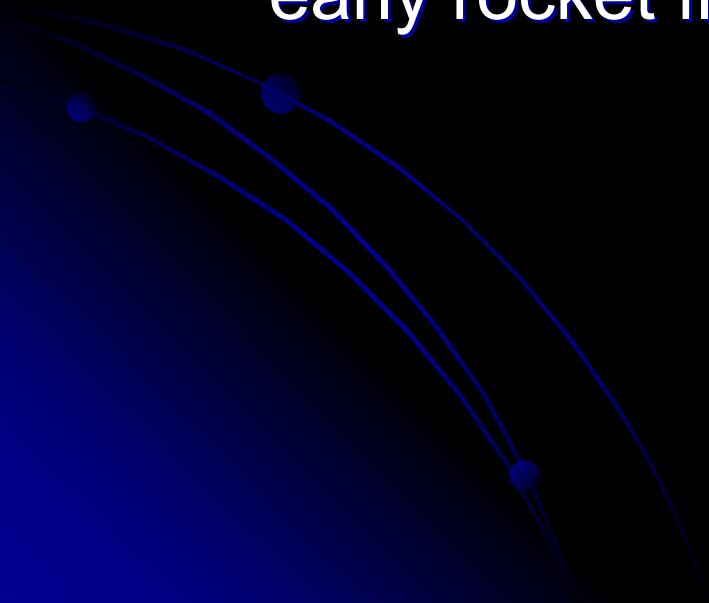


# Rocket 17.09

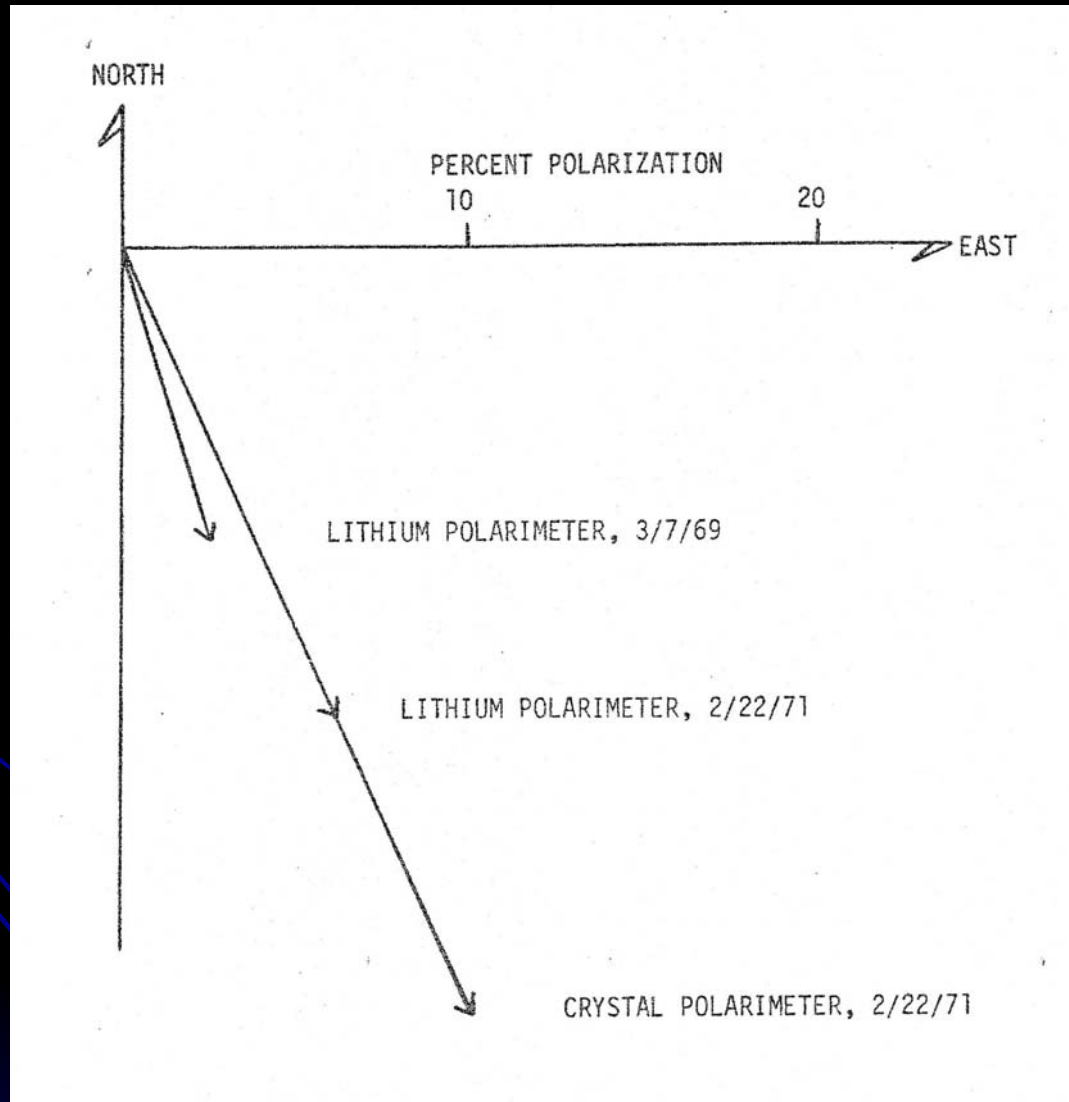


# Systematic Effects

- Critical design consideration
- Astrophysical Polarization may be/will be small
  - E.g. “east-west” effect was of concern in these early rocket flights



# The Results



# Letters ... We send Letters...

UNIVERSITY OF ILLINOIS AT URBANA-CHAMPAIGN  
DEPARTMENT OF PHYSICS  
URBANA, ILLINOIS 61801

11/9/73

Dear Bob,

This took a little longer to get out than we had anticipated; we hope that it will still prove of help. Indeed, it should - since it seems to me quite likely that polarization measurements may be of significant assistance in deciding whether some of the sources are neutron stars or black holes.

Best

M. Rees

Madingley Road  
Cambridge CB3 0HA  
Telephone (0223) 62204

University of Cambridge  
Institute of Astronomy

Director:  
Professor D. Lynden-Bell

6 November 1973

Dr J. Naugle, Code S  
Associate Administrator  
Office of Space Science  
National Aeronautics and Space Administration  
Washington, D. C. 20546  
U.S.A.

Dear Dr Naugle,

The purpose of this letter is to emphasise the increasingly important potential role of polarimetric studies in a balanced program of cosmic X-ray astronomy. My only credentials for venturing an opinion on this matter are that I am in close contact with many members of the X-ray astronomy community in

- 3 -

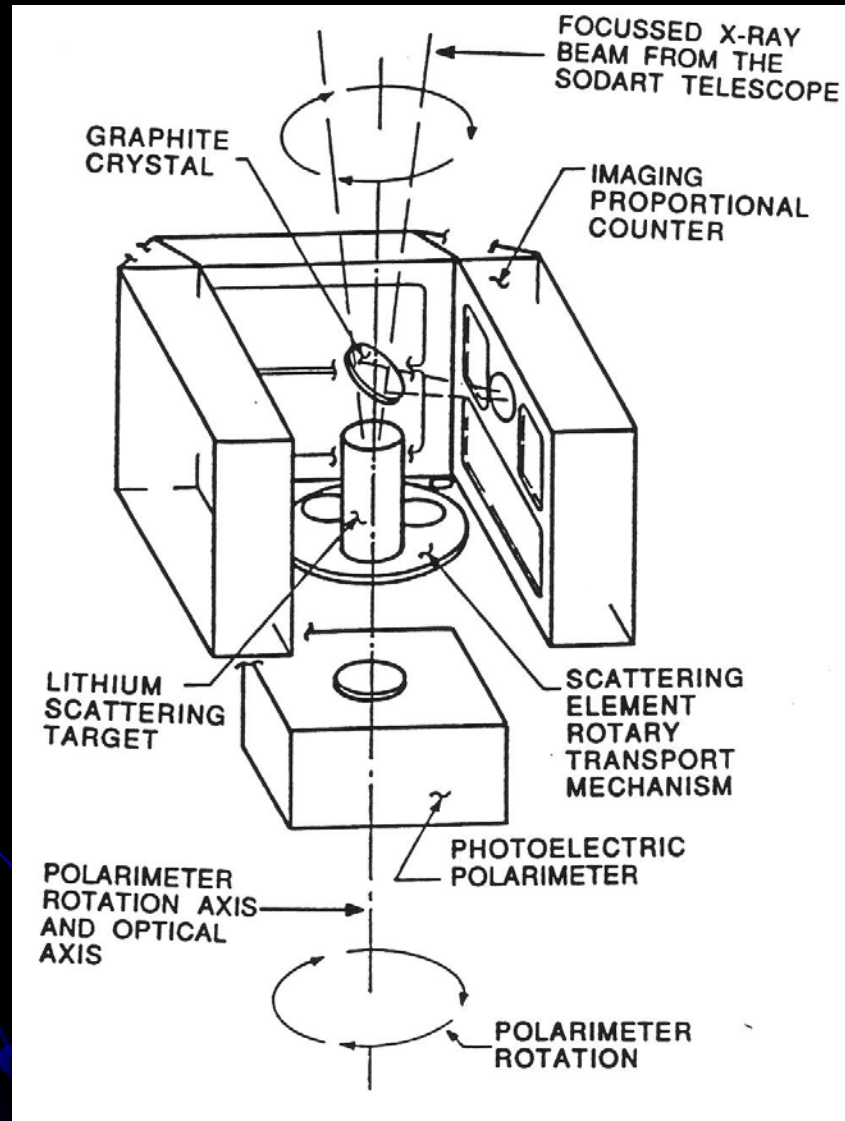
me to express my personal hope that the claims of polarimetry will not be overlooked when new plans are being drawn up. Improved polarimetric observations will be as essential to our understanding of cosmic phenomena as are improvements in positional accuracy, sensitivity and spectral resolution, and should certainly be a part of a balanced and successful X-ray astronomy program.

Yours sincerely,

M. Rees

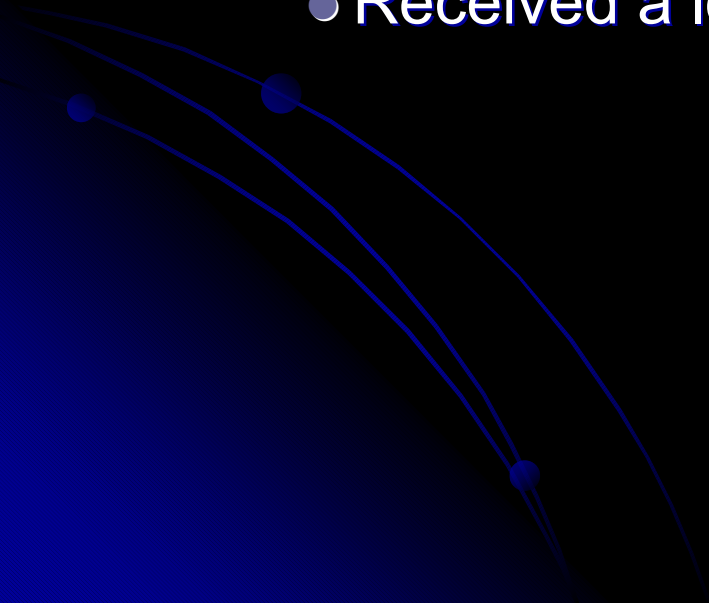
Martin Rees  
(Professor of Astronomy)

# New Concepts - Spectrum-X



# Spectrum-X

- Lithium (and crystal) scattering block at the focus of a telescope
- New systematic effects to worry about
  - Impacts of off-axis pointing
    - Received a lot of attention

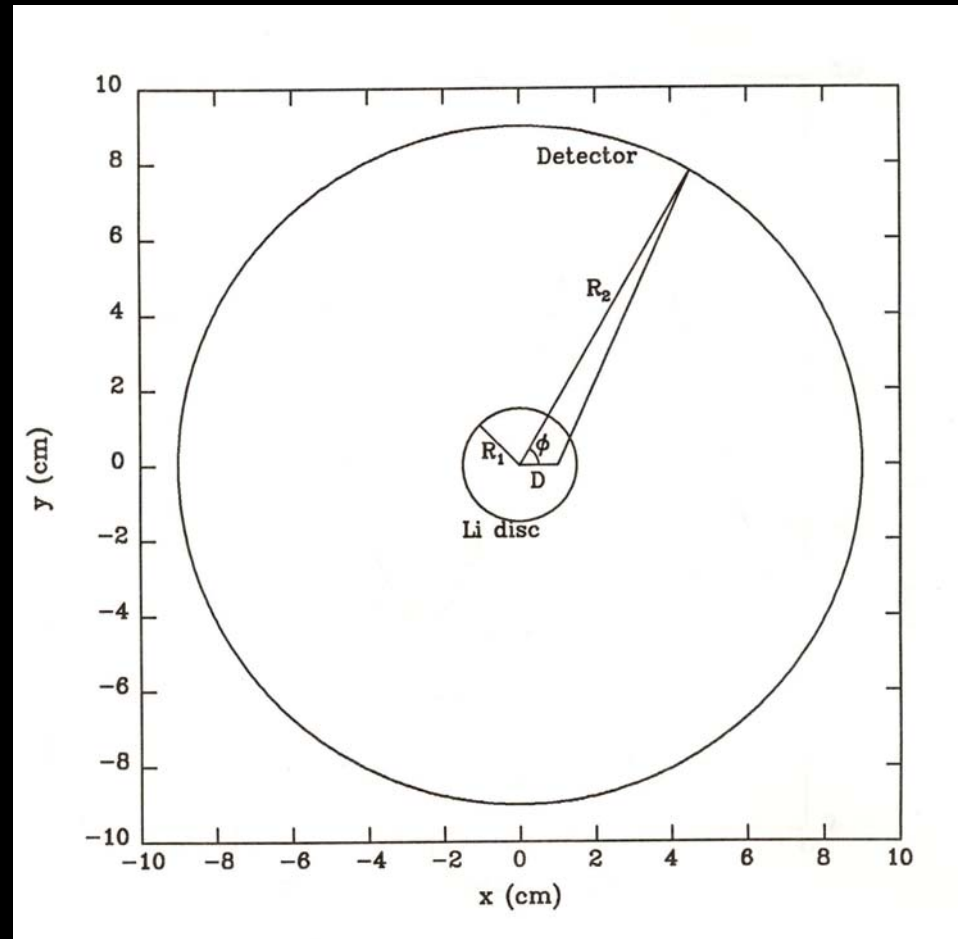


# Off-Axis Pointing

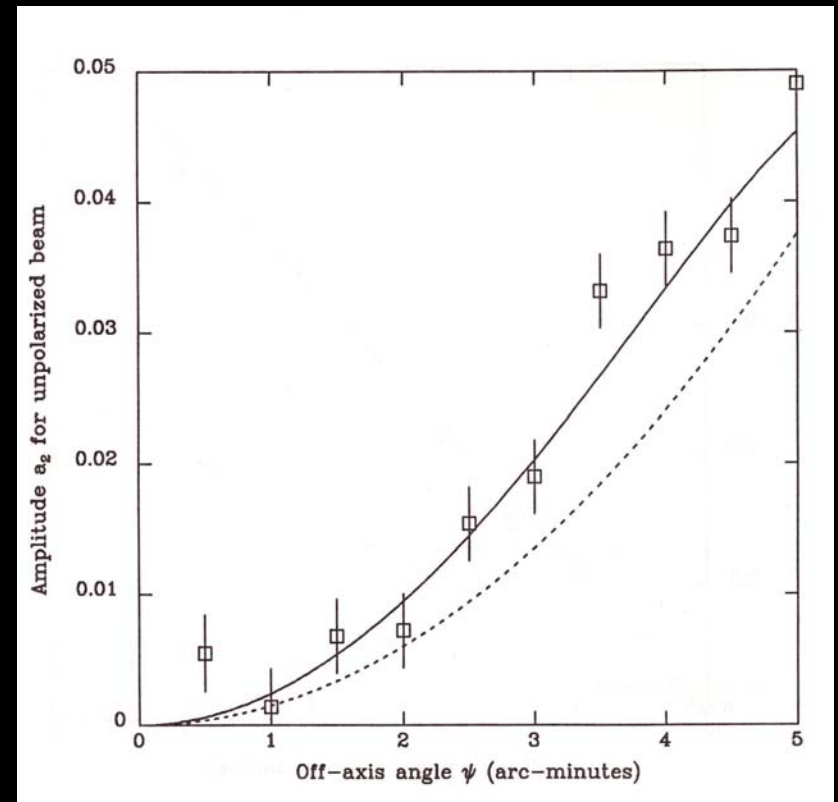
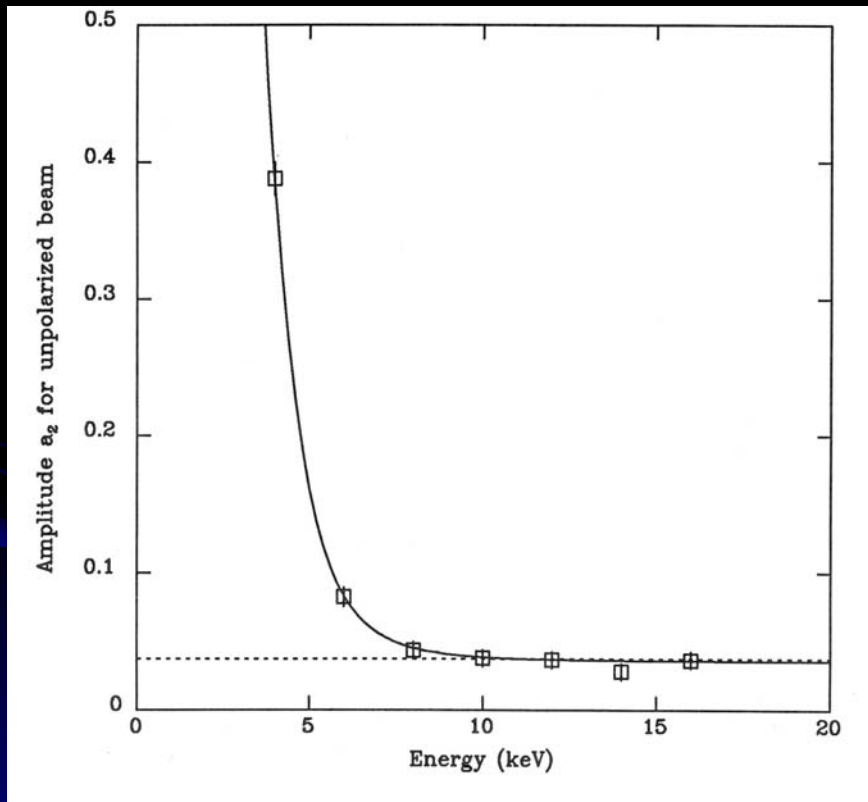
- A solid angle effect

$$P \cong 100(a_2 / \mu)$$

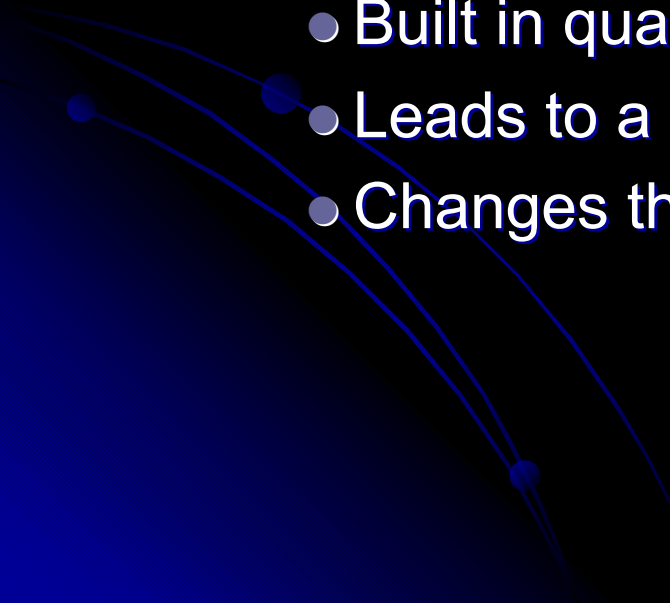
$$\cong [(\psi / \theta_{res})^2 / \mu]$$



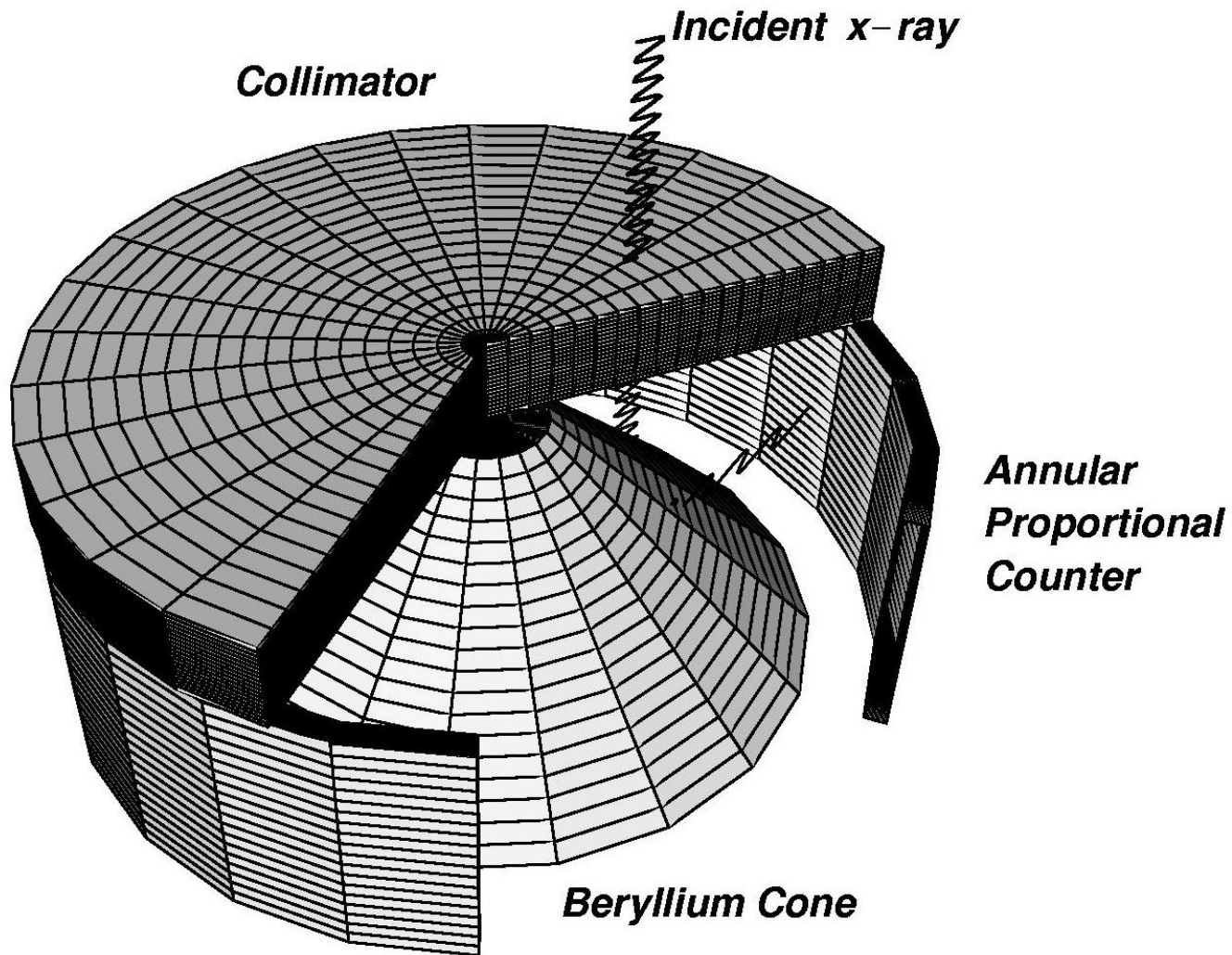
# Off-Axis Pointing - 1



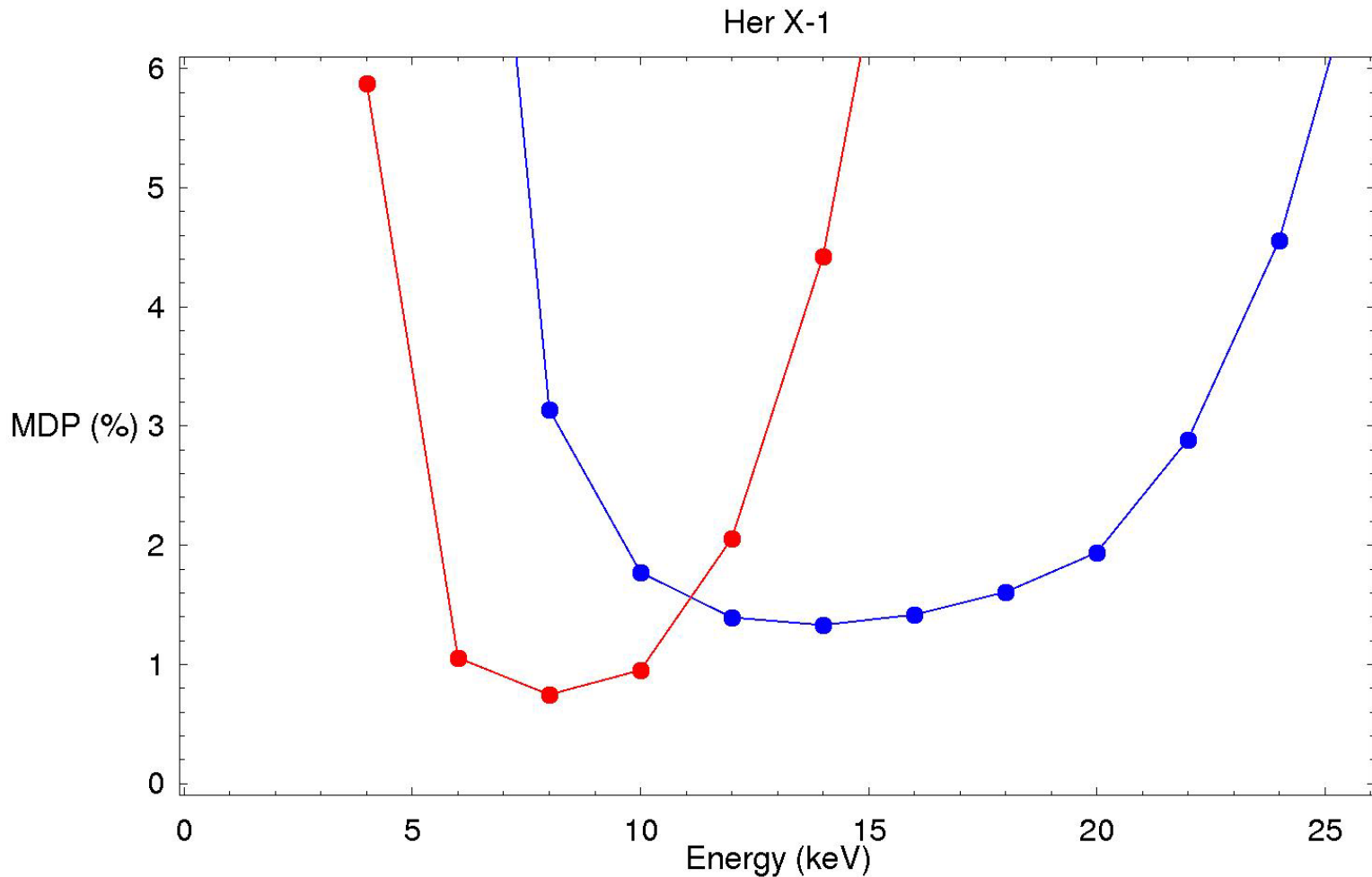
# Other Systematic Effects

- Off-axis pointing also polarizes an unpolarized beam
    - Symmetry broken even if the telescope is perfectly symmetric
  - Impact of the asymmetries in the telescope
    - Built in quadrants!
    - Leads to a second component at “ $2\omega$ ”
    - Changes the statistics, sensitivity, etc.
- 

# New Concepts - XPE



# MDP for SXRP & XPE ( $10^5$ seconds)



# Conclusions

- Inherently broad band
  - Based on sound physical principles
  - Have been built and flown
    - Worked as predicted
  - Relatively inexpensive
- 