







## How can we search for double beta decay? Carter Hall University of Maryland









# **Neutrinoless Double Beta Decay** ( $\beta\beta0\nu$ )

Forbidden if neutrino mass is Dirac only



we can fight it with Avagadro's number

## **Two-Neutrino Double Beta Decay:**

	e <sub>L</sub> -	
2n W-	V VR	e <sub>L</sub> -
to two protons and four leptons.		$\overline{v_{\rm p}}$
First direct observation by Moe, Elliott, and Hahn in <sup>100</sup> Mo (1988)	2p	

No direct implications for neutrino physics, but useful for constraining the nuclear matrix element calculations

# $\beta\beta0\nu$ strategy: search for a peak in the summed electron energy spectrum at the known Q value



Summed electron energy in units of the kinematic endpoint (Q)

### **Choosing a double beta decay source isotope:**



 $N(Z,A) \rightarrow N(Z+2,A)e^{-}e^{-}$ : daughter nucleus must have a smaller mass than the parent for the decay to occur. <sup>5</sup>

### **Choosing a double beta decay source isotope**

Decay candidate	Q value (MeV)	natural abundance (%
<sup>48</sup> Ca→ <sup>48</sup> Ti	4.271	0.187
<sup>76</sup> Ge→ <sup>76</sup> Se	2.040	7.8
<sup>82</sup> Se→ <sup>82</sup> Kr	2.995	9.2
<sup>96</sup> Zr→ <sup>96</sup> Mo	3.350	2.8
<sup>100</sup> Mo→ <sup>100</sup> Ru	3.034	9.6
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}Cd \rightarrow ^{116}Sn$	2.802	7.5
$^{124}Sn \rightarrow ^{124}Te$	2.228	5.64
<sup>130</sup> Te→ <sup>130</sup> Xe	2.533	34.5
<sup>136</sup> Xe→ <sup>136</sup> Ba	2.479	8.9
$^{150}Nd \rightarrow ^{150}Sm$	3.367	5.6

N→N'e⁻e⁻

Q value =  $\dot{M(N)} - M(N')$ The electron energies must sum to the Q value by energy conservation

A large Q value is desirable because the decay rate is faster (larger phase space) and the radioactive backgrounds are smaller.

Large natural abundance make the experiment cheaper.

 $\beta\beta0\nu$  candidate isotopes: Q value and natural abundance



High Q value reduces backgrounds and increases the phase space & decay rate,large abundance makes the experiment cheaper.7

 $\beta\beta0\nu$  candidate isotopes: Q value and natural abundance



High Q value reduces backgrounds and increases the phase space & decay rate,large abundance makes the experiment cheaper.8

# Virtually all materials contain small amounts of radioactive isotopes:



Natural radioactive decay chains

Compton scattering:  $\gamma e^- \rightarrow \gamma e^-$ broad energy spectrumPhotoelectric effect:  $\gamma e^- \rightarrow e^-$ monochromaticPair production:  $\gamma \rightarrow e^+ e^-$ monochromatic

# Shielding a detector from gammas is difficult because the absorption cross section is small.



**Example:** γ interaction length in Germanium is 4.6 cm, <sub>10</sub> comparable to the size of a germanium detector.



# Why go underground?

- Studies for rare events, either decays (eg proton or 0vββ) or weak interactions (dark matter, natural or generated neutrino), require very radio-quiet environments to undertake searches
- Deep underground facilities provide significant rock overburden and commensurate reduction in c.r. flux, and c.r.-spallation induced neutrons
  - Additional science programmes possible with such infrastructure - nuclear astrophysics, extreme biosystems, geology, geophysics, ...



# The most sensitive double beta decay experiments to date are based on 76-Germanium.



Half-life limit:  $1.9 \ge 10^{25}$  years (H-M and IGEX) Majorana neutrinos ruled out for masses greater than ~0.35-1.0 eV

# Energy resolution of $\beta\beta0\nu$ candidate isotopes

Gotthard TPC (136Xe) energy spectrum (1998)



Superior energy resolution: 76Ge (diode): 0.2% FWHM 130Te (bolometer): 0.4% FWHM Modest energy resolution: 136Xe (liquid TPC): 3.3% FWHM 100Mo, 82Se (plastic scintillator): ~14% FWHM

# $\beta\beta0\nu$ discovery claim – 2001-2006



HV. Klapdor-Kleingrothaus, et. al, Nuclear Instruments and Methods A **513** (2004) 371-406<sup>14</sup>

## ββ0v discovery claim - 2004



# **Discovery of the \Omega-**

VOLUME 12, NUMBER 8

PHYSICAL REVIEW LETTERS

24 FEBRUARY 1964



FIG. 2. Photograph and line diagram of event showing decay of  $\Omega^-$ .

The statistical significance of a signal is determined by how strongly you reject the null hypothesis.

# Germanium-76

### Heidelberg-Moscow, IGEX, GERDA, MAJORANA experiments



2-3 kg Ge diodes, 80% <sup>76</sup>Ge

### Lead shielding

### Cu cryostats

Fantastic energy resolution (4 keV FWHM).Cooled with Liquid Nitrogen to suppress thermal noise.Pulse shape analysis rejects multiple site events within a single crystal.Suffers from low Q value (2039 keV), and cosmogenic activation of germanium and copper cryostats.

#### MPIK Heidelberg

# MAJORANA project status

- Demonstrator approved for FY 2010-2013
  - 30 kg <sup>nat</sup>Ge & 30 kg <sup>enr</sup>Ge
  - Running 3 years (90 kg·y)  $\rightarrow$  T<sub>1/2</sub>  $\ge$  10<sup>26</sup> y (90% CL))
  - $B = 10^{-3} cts/(kg \cdot keV \cdot y)$
- Objective: Demonstrate background low enough to justify building a ton scale Ge experiment

## • Schedule:

- Start of Cu electroforming deep underground at DUSEL this year
- First cryostat with 20 kg of <sup>nat</sup>Ge modified BEGe p-type detectors ready in fall 2011





See posters 4, 95 & 120



Marik Barnabé Heider

**MPIK Heidelberg** 



# **GERDA** status

- Summer/autumn '09: Integration test of Phase I detector string, FE, lock, DAQ
- Nov/Dec.'09: Liquid argon filling
- Apr/May'10: Installation of 1-string lock in the GERDA cleanroom
- May '10: Deployment of FE & detector mock-up, followed by first deployment of a of non-enriched detector
- June '10: Water tank filling
- June '10: Commissioning run with <sup>nat</sup>Ge detector string



#### **MPIK Heidelberg**



# **Tellurium basics: low temperature bolometers**

Take advantage of tiny heat capacity of crystals at low temperature to measure energy deposition with a thermometer(!).



Technique applicable to many isotopes – currently 130Te is used to take advantage of its high isotopic abundance (30%).

Drawbacks: no information beyond energy is available, like particle ID ( $\alpha, \beta, \gamma$ ), event location, or topology.

# **130Te : CUORICINO**

Italy, US, Netherlands, Spain, at Gran Sasso, Italy



#### tower under construction



CUORICINO is a 40.7 kg tower of TeO2 crystals (34% 130Te) in operation at Gran Sasso.



Maura Pavan – Università di Milano Bicocca and sez. INFN - Neutrino 2010 - Athens - June 14-19

#### CUORICINO: On DBD result

TOTAL EXPOSURE 19.75 [kg(<sup>130</sup>Te) yr]







Maura Pavan – Università di Milano Bicocca and sez. INFN - Neutrino 2010 - Athens - June 14-19





### CUORE goal 5 years sensitivity

Background	$\Delta E_{_{FWHM}}$	$ au_{\frac{1}{2}}^{0v}$	m <sub>ee</sub> [meV]			
[c/keV/kg/y]	[keV]	[y] @ 68%C.L.	$R(QRPA)^1$	pn(QRPA) <sup>2</sup>	ISM <sup>3</sup>	IBM-2 <sup>4</sup>
0.01	5	<b>2.1×10</b> <sup>26</sup>	35÷66	41÷67	65÷82	41
0.001	5	6.5×10 <sup>26</sup>	20÷38	23÷38	37÷47	23

NME bibliography:

1 Šimkovic et al., PRC 77 (2008) 045503 2 Civitarese et al., JoP:Conference series 173 (2009) 012012 3 Menéndez et al., NPA 818 (2009) 139 4 Barea and Iachello, PRC 79 (2009) 044301

# **100Mo: NEMO-3**

France, Russia, Japan, US, Czech Republic, currently taking data in Frejus, France.



Passive isotope on thin foils surrounded by Geiger mode drift cells for electron tracking and plastic scintillator for energy measurement

NEMO is designed to investigate the  $\beta\beta0\nu$  mechanism: multiple isotopes and electron angular distribution.

# **100Mo: NEMO-3**



# **100Mo: NEMO-3**

25 Gauss magnetic field measures  $\beta$  charge.



6.9 kg of 100Mo and 0.93 kg of 82Se for  $\beta\beta$ 0v search.

Other isotopes for  $\beta\beta2\nu$  measurements and background studies.

# 100Mo and 82Se: NEMO-3

Data for 100Mo  $\beta\beta 2\nu$  agree beautifully with simulations .



Super-NEMO: 100 kg with neutrino mass sensitivity of 50 meV.

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From K. Nakamua, Neutrino 2010

# SNO+

### 1000 t D<sub>2</sub>O will be replaced by Nd loaded LS 0.1 wt% = 780 kg Nd(natural)

= 44 kg Nd-150

#### 9500 PMTs -

Energy res = 5 %@1 MeV

7000 t pure water shield

Hold down ropes will be installed

### From K. Nakamua, Neutrino 2010



### From K. Nakamua, Neutrino 2010

# Sensitivity



### From K. Nakamua, Neutrino 2010 KamLAND-Zen



<sup>136</sup>Xe 400 kg loaded LS in mini-balloon, R=1.7m

### <sup>136</sup>Xe 400 kg:

2.7 wt% dissolved into LS easy handling/ enrichment (90%) longer 2∨ beta decay life time T<sup>2∨</sup> >10<sup>22</sup> years (cf: ~10<sup>19-20</sup>)
KamLAND exists:

ultra pure environment (U/Th~10<sup>-17</sup> g/g) LS techniques Balloon experience LS Density control techniques Reactor/Geo neutrino



## From K. Nakamua, Neutrino 2010 Sensitivity











University of California, Irvine













# Xenon can be continuously purified of chemical and radioactive contaminants

LXe purity with recirculation, May 17-25



Fluids often have extraordinarily low radioactivity, and noble gases are particularly simple to purify.

# Liquid xenon calorimetry



Measure the event energy by collecting the ionization on the anode and/or observing the scintillation.

## Liquid xenon data show an anti-correlation between ionization and scintillation



Factor of two better than most recent Xe experiment

# The crown jewels of EXO



200 kg of xenon enriched to 80% in <sup>136</sup>Xe:

11 times larger than previous double beta decay experiments. 40

# **EXO-200:** the first 200 kg $\beta\beta0\nu$ experiment



# EXO-200 Cryostat & Lead (March 2008)



## **EXO-200 TPC Construction**





## **EXO-200 – Final TPC installation in January 2010**



### **EXO-200: First data expected in September**



# **Sensitivity of EXO-200**

Case	Mass	Eff.	Run	σ <sub>E</sub> /Ε @	Radioactive	<b>T</b> <sub>1/2</sub> <sup>0v</sup>	Majorana mass	
	(ton)	(%)	Time	2.5MeV	Background	(yr,	(m	eV)
			(yr)	(%)	(events)	90%CL)	QRPA <sup>1</sup>	NSM <sup>2</sup>
EXO-200	0.2	70	2	1.6*	40	6.4*10 <sup>25</sup>	109	135

1) Simkovic et al. Phys. Rev. C79, 055501(2009) (use RQRPA and  $g_A = 1.25$ )

2) Menendez et al., Nucl. Phys. A818, 139(2009), (use UCOM results)

# Xe offers a new tool to reduce background:

 $^{136}$ Xe  $\rightarrow$   $^{136}$ Ba<sup>++</sup> final state can be identified using optical spectroscopy (M.Moe PRC44 (1991) 931)

Ba<sup>+</sup> system best studied (Neuhauser, Hohenstatt, Toshek, Dehmelt 1980) Very specific signature "shelving" Single ions can be detected from a photon rate of 10<sup>7</sup>/s

Barium tagging would eliminate all radioactive backgrounds, leaving only  $2\nu\beta\beta$ .



Level structure for Ba+

# **EXO Ba+ trapping Experiment**



RF quadrupole trap loaded in UHV from a Ba dispenser and e-beam ionizer. Xe can be injected while observing the ions at pressures from  $10^{-10}$  torr to 0.1 torr=

# **EXO spectroscopy lab**



650 nm: External Cavity Diode Laser (ECDL)

493 nm: Frequency doubled 986 nm

Ba Oven



# Ba<sup>+</sup> Tagging: Ion Trap + fluorescence



Dolinski

16 June 2010

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# Ba<sup>+</sup> Tagging: RIS



16 June 2010

# Ba<sup>+</sup> Tagging: RIS



"single ion mode" setup about to start taking data.

holds Ba atoms

# Sensitivity of ton-scale EXO with barium tagging

Case	Mass	Eff.	Run	σ <sub>E</sub> /E @	2νββ	T <sub>1/2</sub> <sup>0v</sup>	Majora	ina mass
	(ton)	(%)	Time	2.5MeV	Background	(yr,	(meV)	
			(yr)	(%)	(events)	90%CL)	QRPA	<sup>1</sup> NSM <sup>2</sup>
Conserva tive	1	70	5	1.6*	0.5 (use 1)	<b>2*10</b> <sup>27</sup>	19	24
Aggressi ve	10	70	10	1†	0.7 (use 1)	4.1*10 <sup>28</sup>	4.3	5.3

- 1) Simkovic et al. Phys. Rev. C79, 055501(2009)
- 2) Menendez et al., Nucl. Phys. A818, 139(2009)

# What to expect in the next decade



Figure from Strumia & Vissani, Nucl. Phys. B 726 294 (2005)